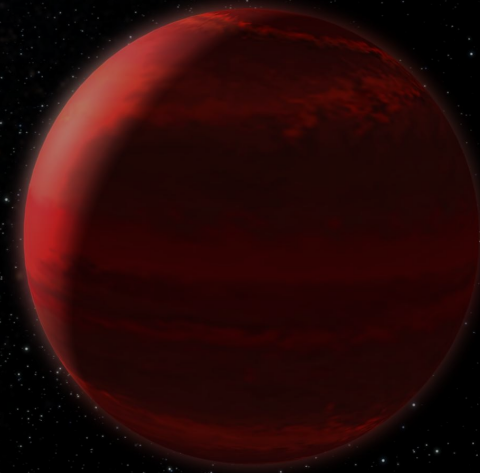


Searching for Habitable Worlds and Life in the Universe with Optics



Contact: oliv.guyon@gmail.com

What makes planets habitable ?

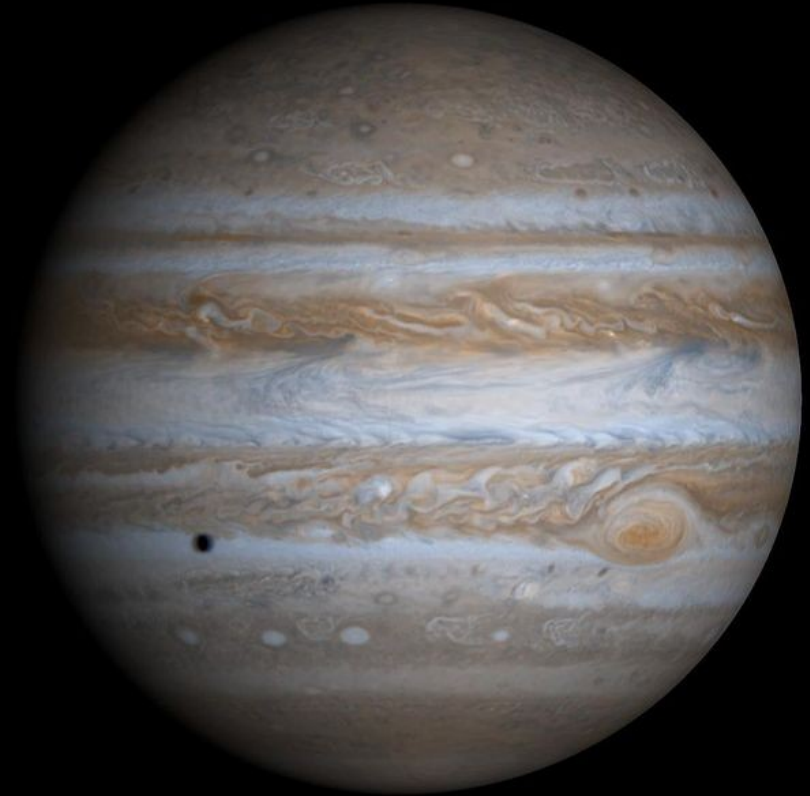
Size matters:
not too big, not too small



**Moon: too small
No atmosphere**



Earth



**Jupiter: too big
Mostly gas**

What makes planets habitable ?

The planet must be in the habitable zone of its star: not be too close or too far



Venera 13 lander, survived 127mn at 457 C, 89 atm

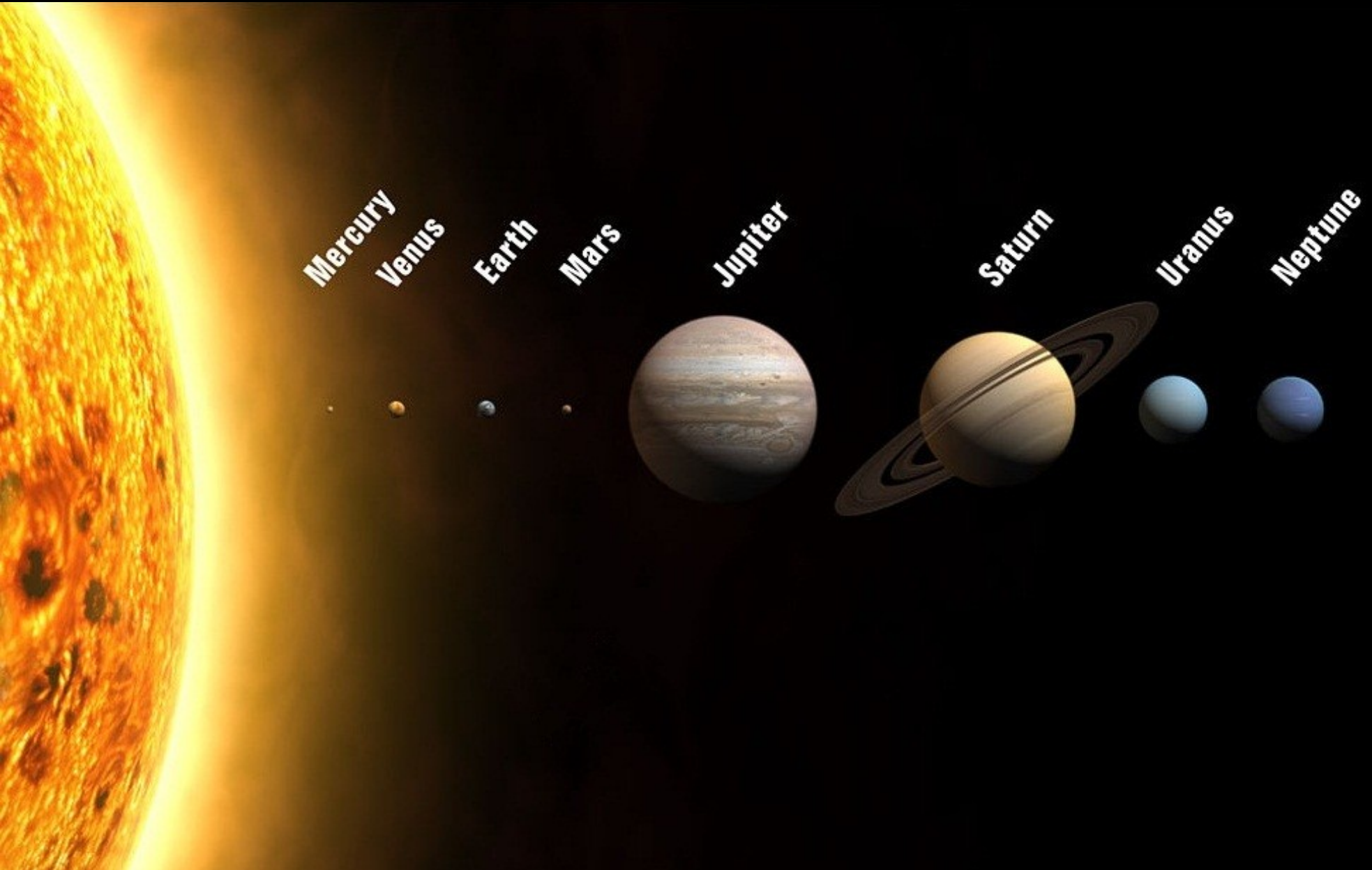
Venus: too close, too hot



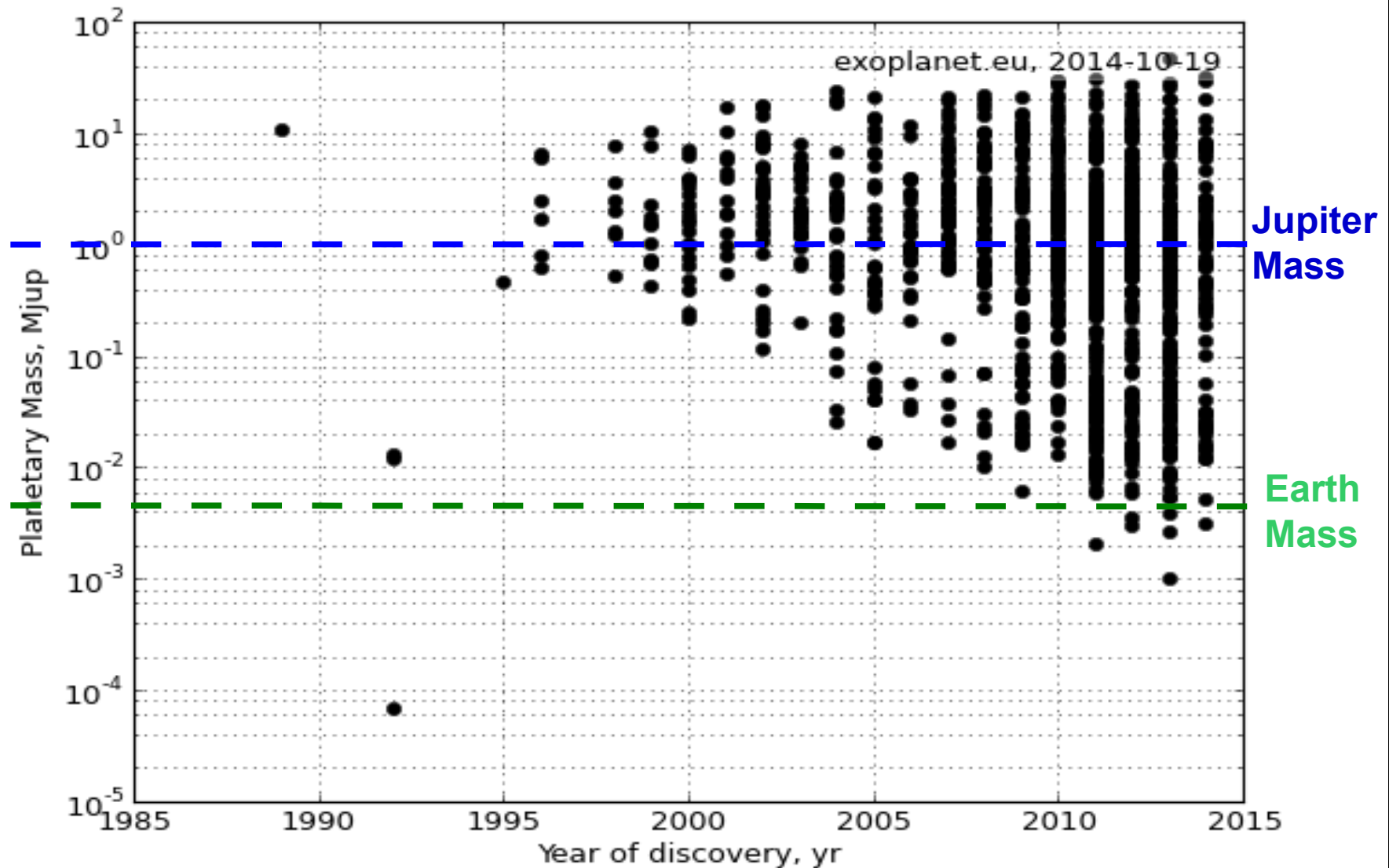
Image credit: NASA/JPL-Caltech/MSSS

Mars: too far, too cold

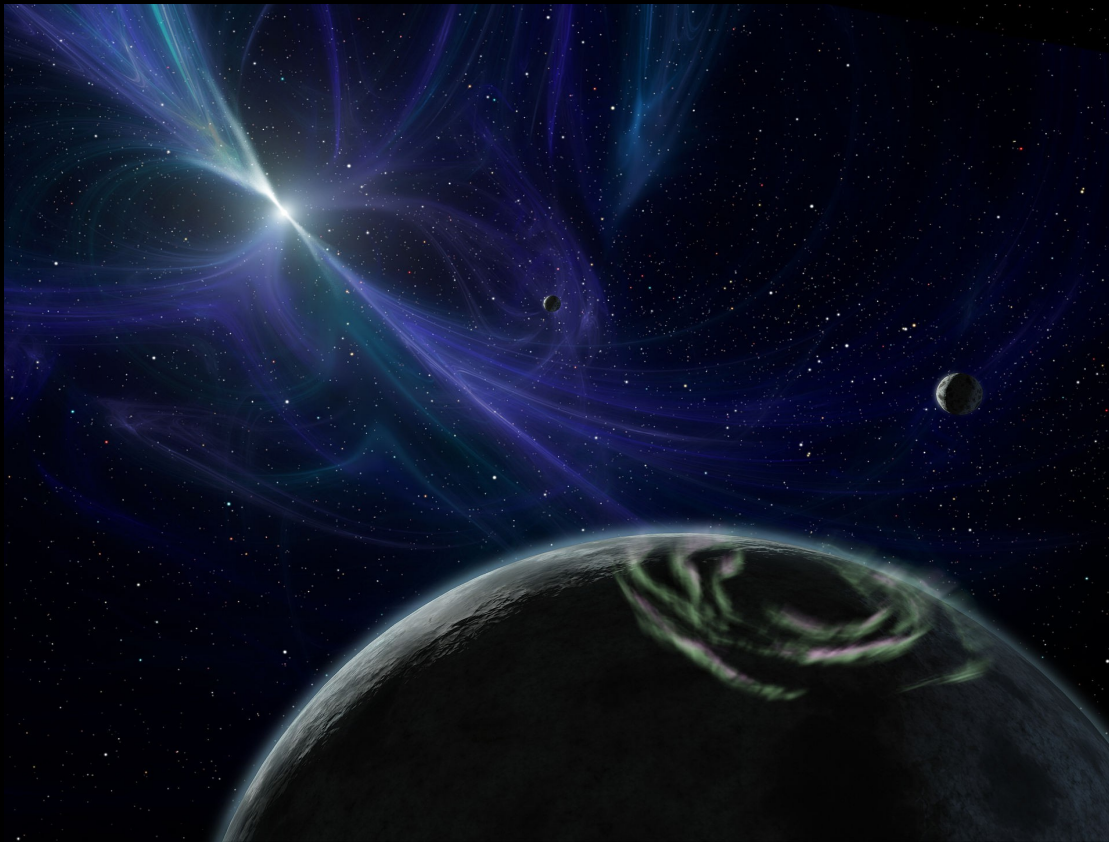
ALL known Planets until 1989



Exoplanets identified – we are now starting to identify Earth-size planets



First discoveries: Oddballs



PSR 1257+12 : 3 planets around a pulsar

How could planets survive a supernova ?

Planets may have formed from debris after the supernova explosion



51 Peg b : The first confirmed “Hot Jupiter”

A Jupiter-size planet at only 5.3% of the Sun-Earth distance from its host star, orbiting in 4.2 days

Approximately 10% of stars have a habitable planet

200 billion stars in our galaxy

→ approximately 20 billion habitable planets

Imagine 200 explorers, each spending 20s on each habitable planet, 24hr a day, 7 days a week.

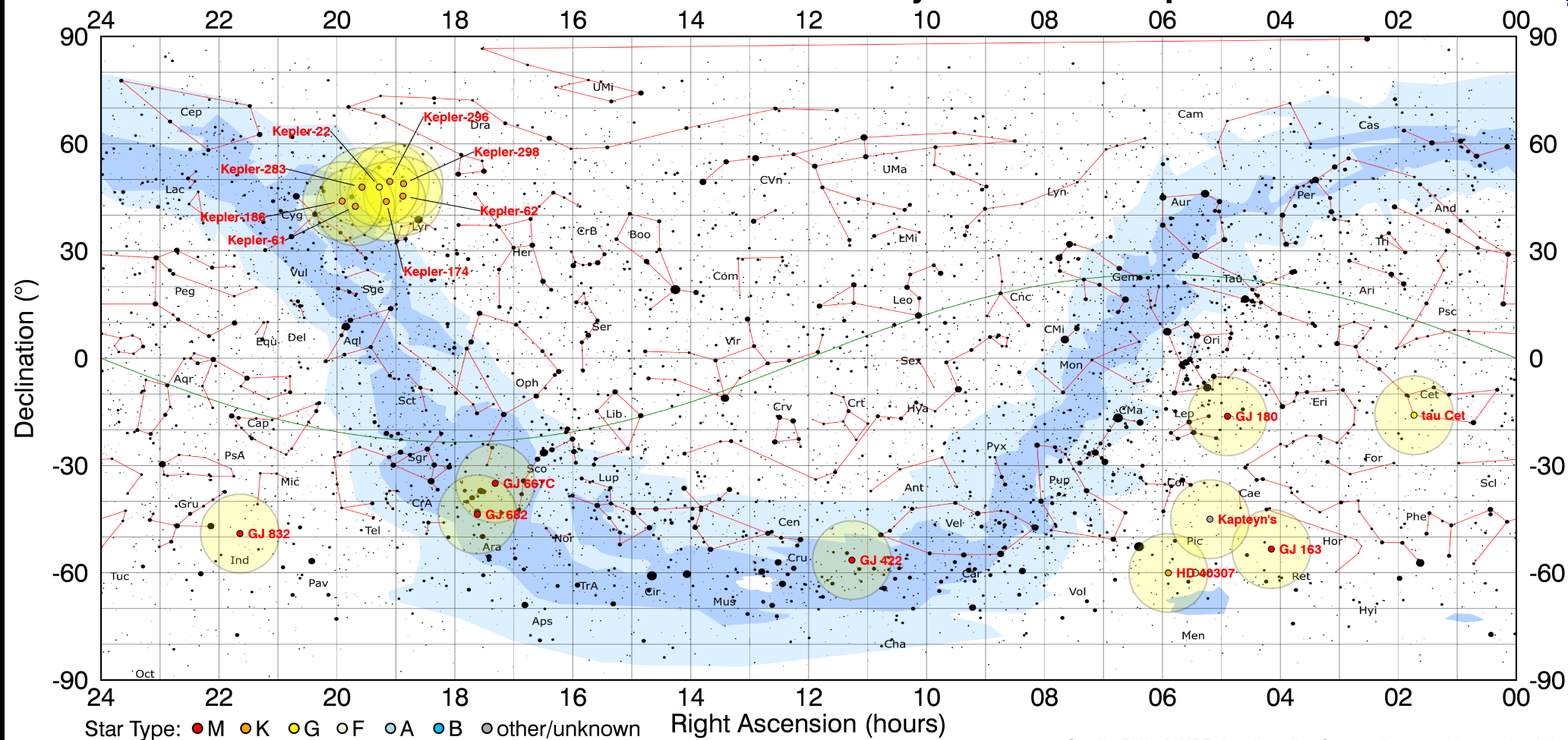
It would take >60yr to explore all habitable planets in our galaxy alone.

x 100,000,000,000 galaxies in the observable universe





Location of the Stars with Potentially Habitable Exoplanets



Credit: PHL @ UPR Arcibo, Jim Cornmell

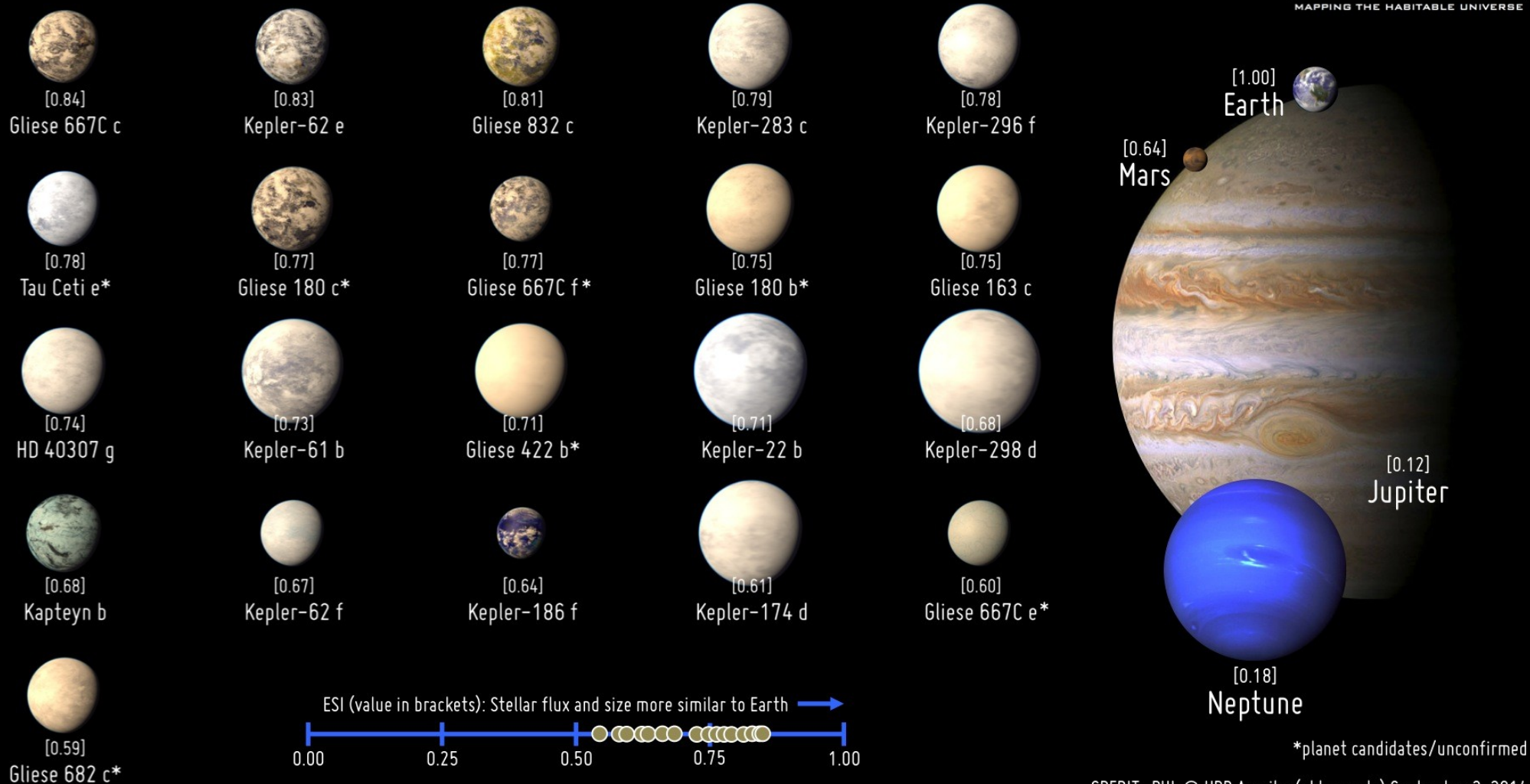
phl.upr.edu, Jul 2014

>~10% of stars have potentially habitable planet

First potentially habitable planets now identified

Current Potentially Habitable Exoplanets

Ranked by the Earth Similarity Index (ESI)



How do astronomers identify exoplanets ?

HIGH PRECISION OPTICAL MEASUREMENTS OF STARLIGHT (indirect techniques)

Earth around Sun at ~30 light year

→ ***Star position moves by 0.3 micro arcsecond
(thickness of a human hair at 20,000 miles)***

→ ***Star velocity is modulated by 10cm / sec
(changes light frequency by 1 part in 3,000,000,000)***

***If Earth-like planet passes in front of Sun-like star, star dims by
70 parts per million
(12x12 pixel going dark on a HD TV screen 70 miles away)***



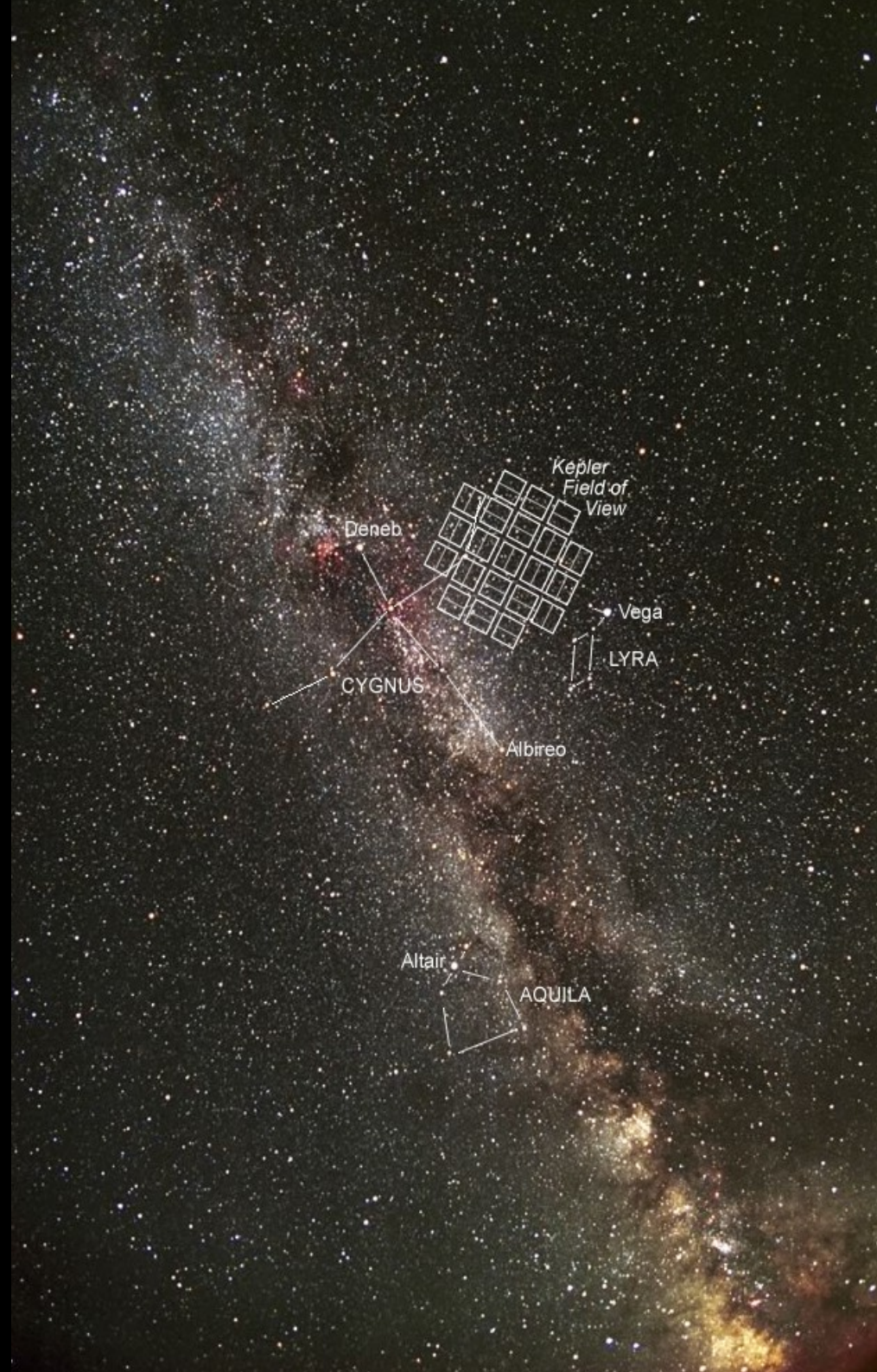
Exoplanet transit

If the planet passes in front of its star, we see the star dimming slightly

Transit of Venus, June 2012



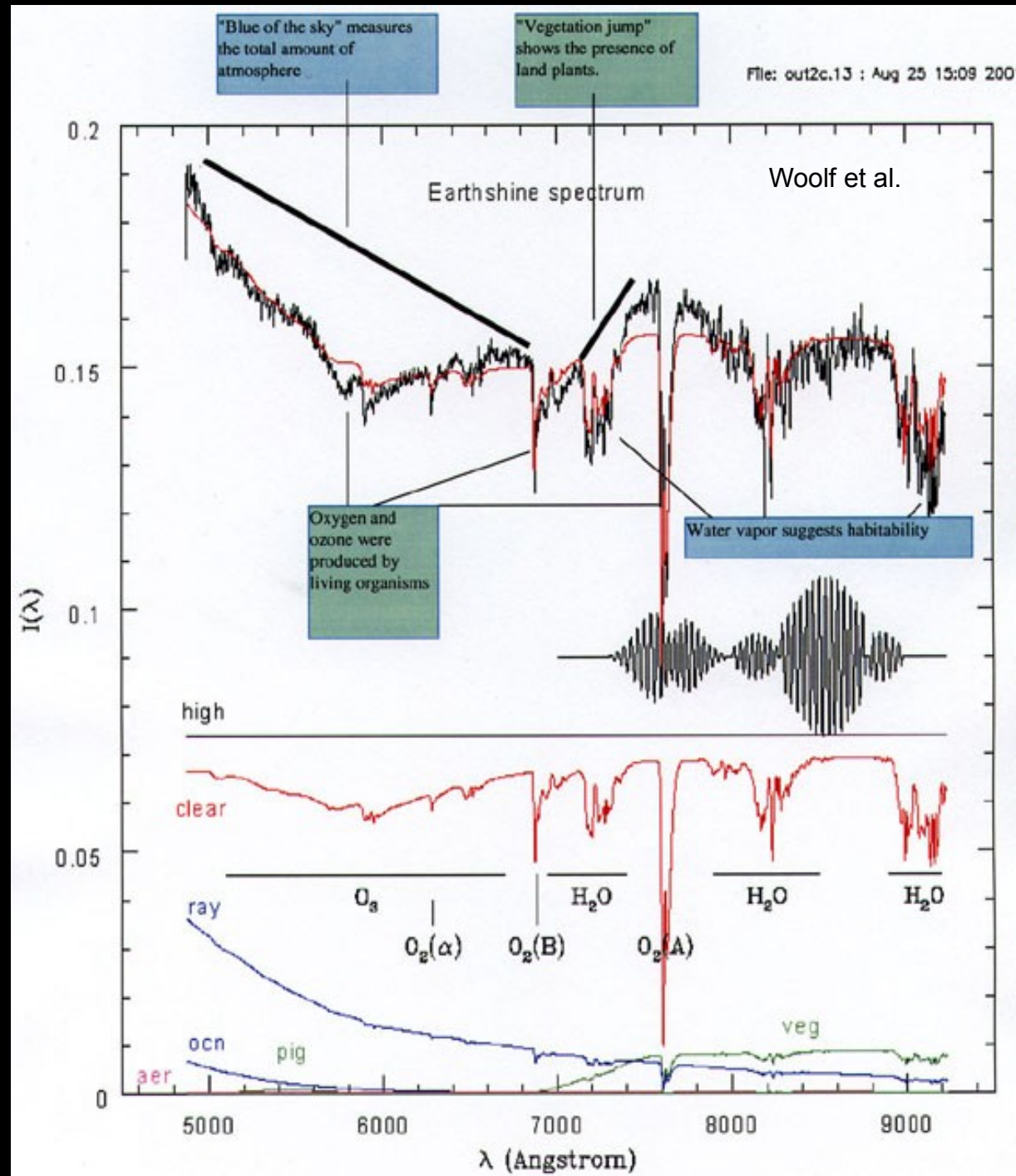
Kepler (NASA)



Directly imaging planet is necessary to find life

We need to take spectra of habitable planets

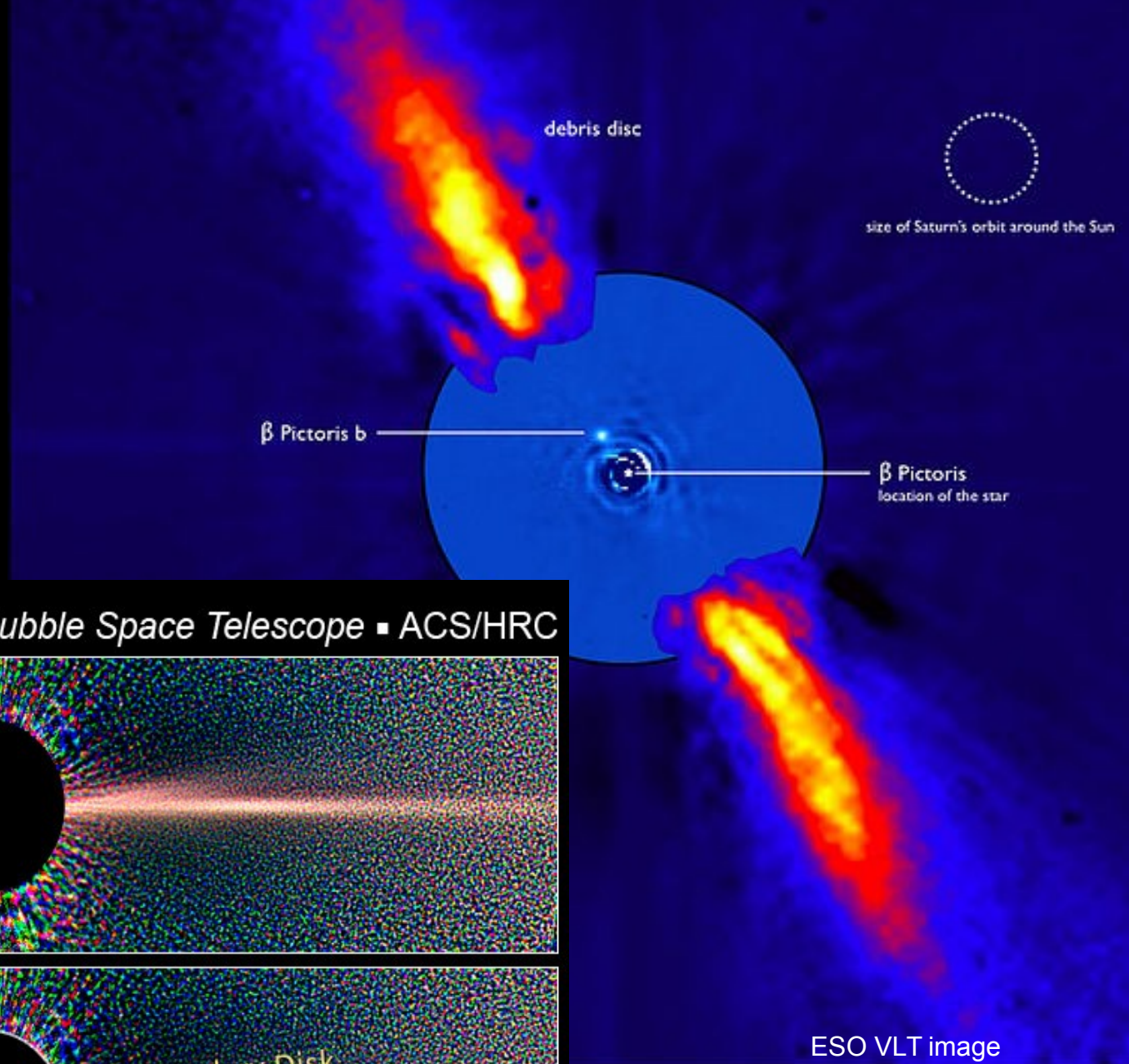
Spectra of Earth (taken by looking at Earthshine) shows evidence for life and plants



Beta Pictoris

8 Jupiter mass planet

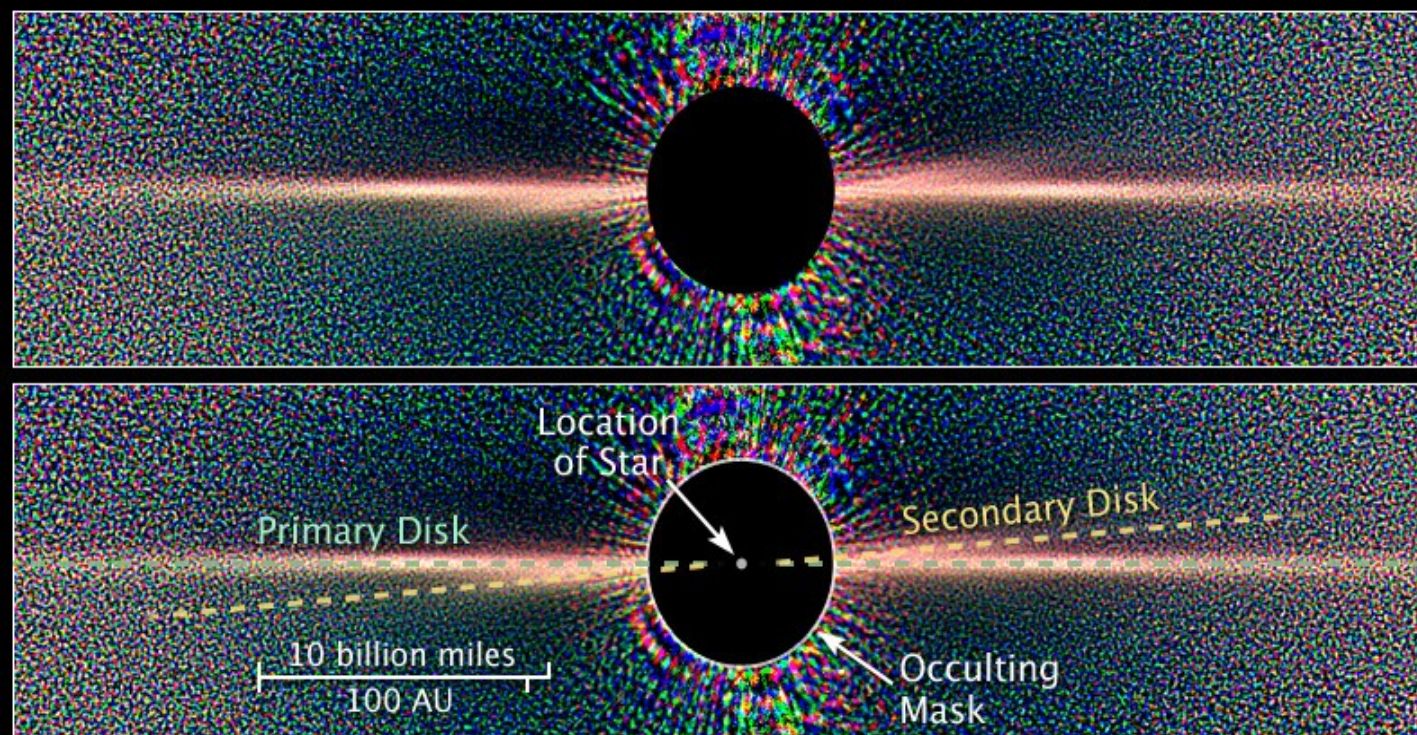
Orbits young massive star in ~20yr



ESO VLT image
(Lagrange et al.)

Beta Pictoris

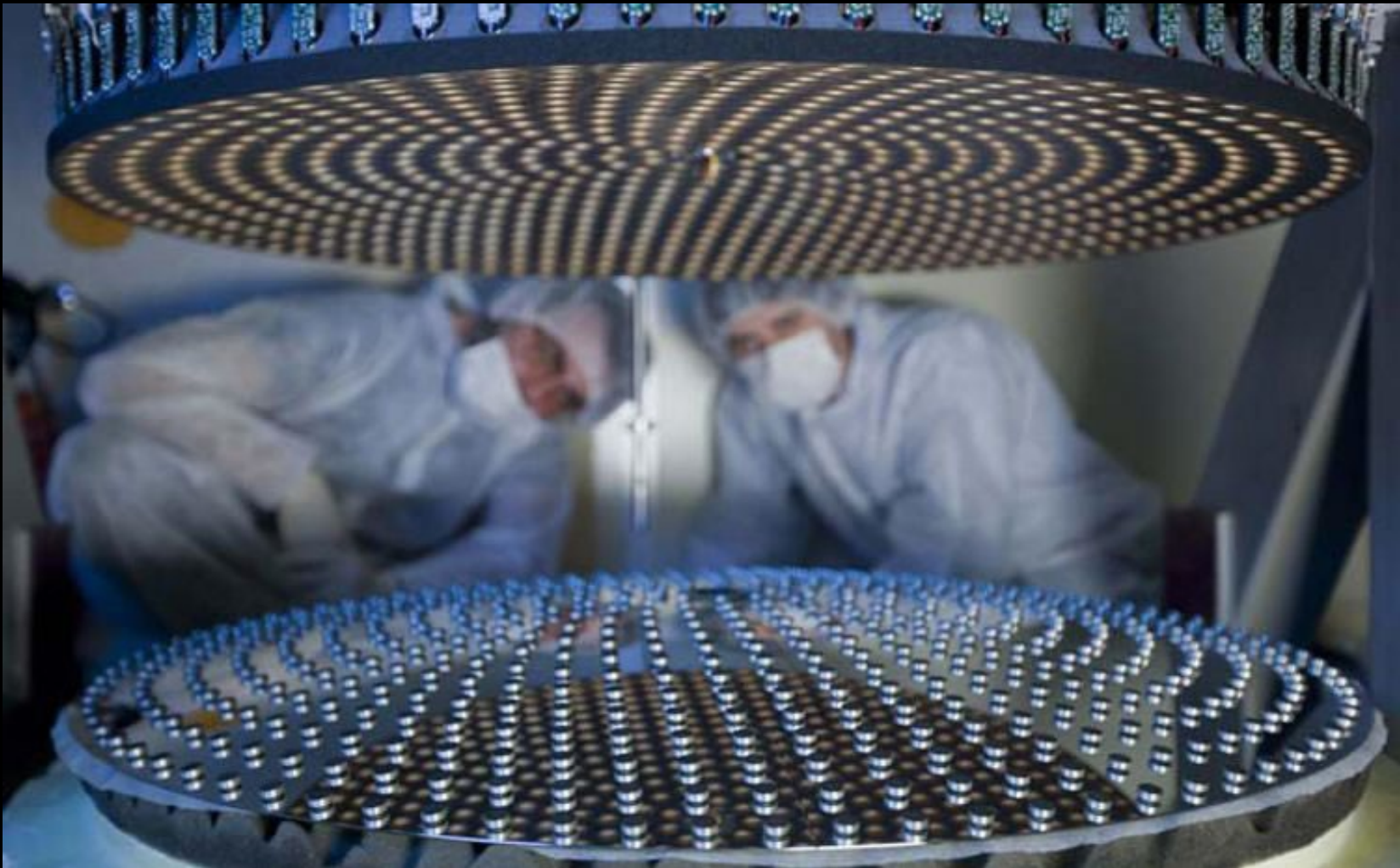
Hubble Space Telescope ■ ACS/HRC



Large Binocular Telescope, AZ



Ground-based telescopes use adaptive optics to mitigate atmospheric turbulence

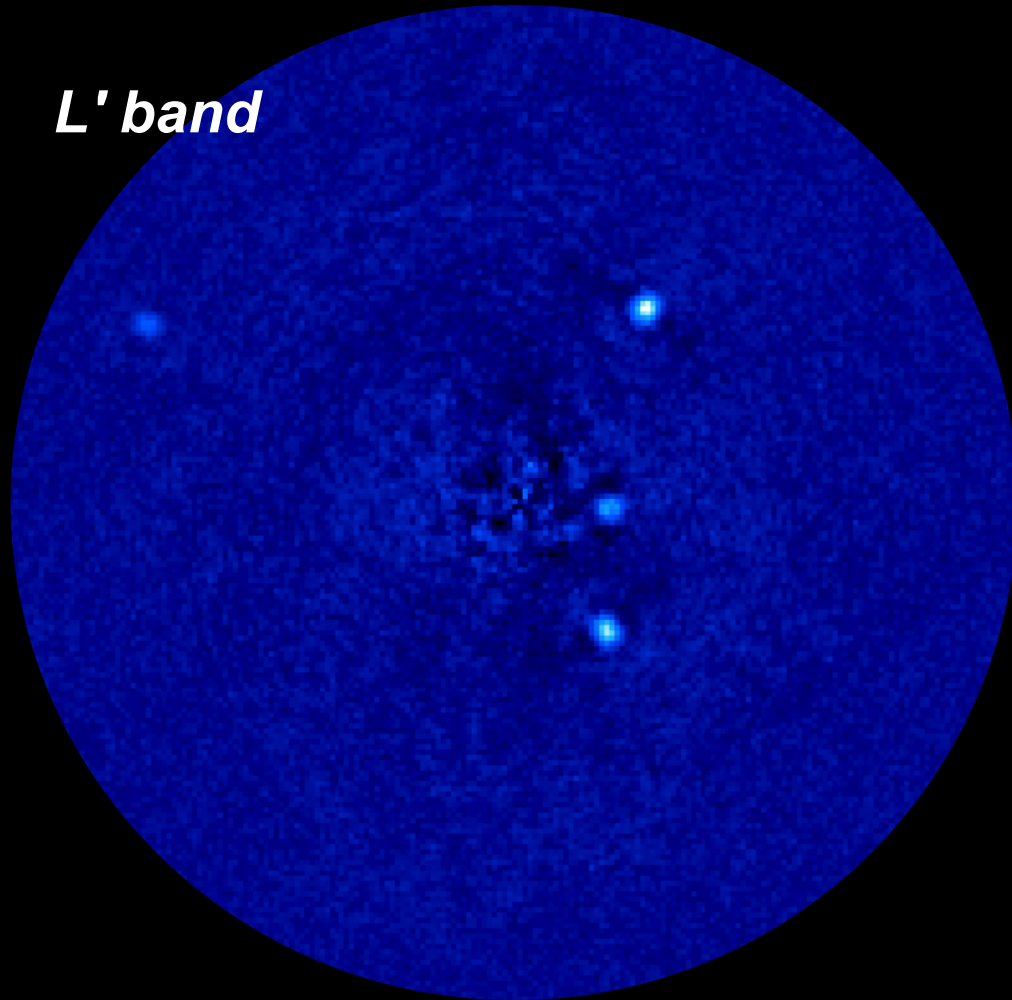


Large Binocular Telescope adaptive secondary mirror

HR8799 imaged with Large Binocular Telescope

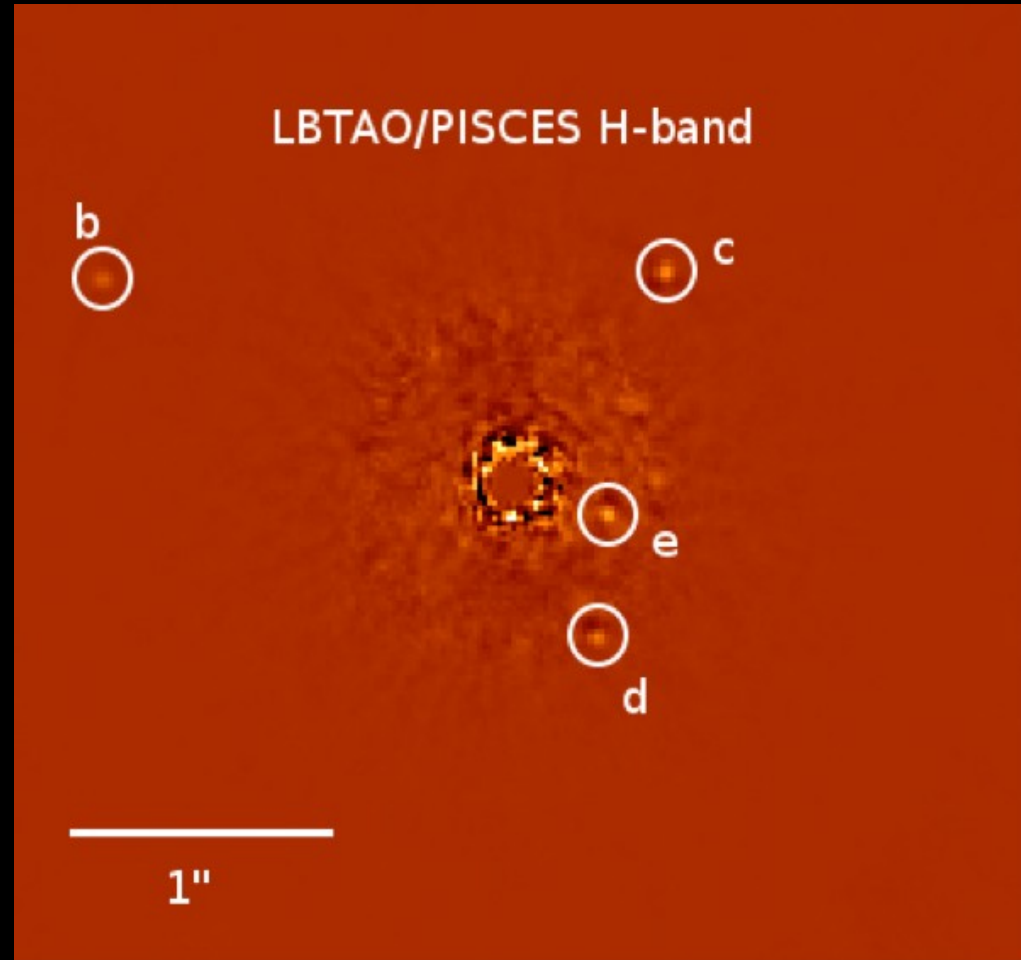
Four planets, orbital periods on the order of 100yr
Each planet 5 to 7 Jupiter Mass

L' band



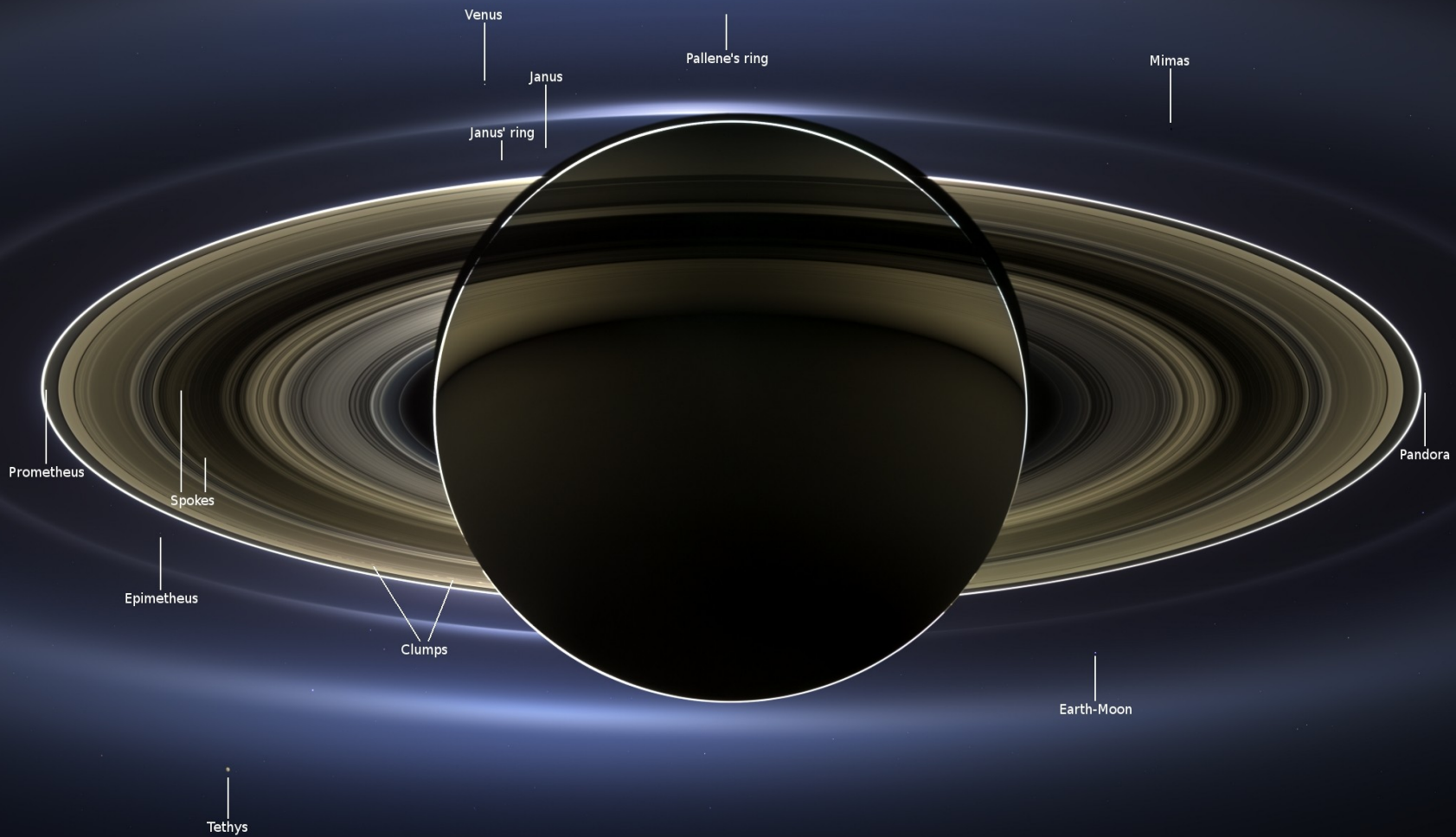
Defrere et al.

LBTAO/PISCES H-band



Skemer et al..

Taking images of habitable exoplanets: Why is it hard?



A photograph taken from space showing the planet Saturn in the upper left corner, with its characteristic yellowish-brown rings and atmosphere. A bright, glowing white ring of light, likely the planet's limb or a reflection, curves across the upper portion of the frame. Below Saturn, a thin, dark, curved line represents the horizon of the Earth. In the lower right, a small, bright blue-white point of light is identified as Earth, with a small white arrow pointing directly at it. The background is a deep, dark blue-black space.

Saturn

↑
Earth

Coronagraphy ... Using optics tricks to remove starlight (without removing planet light)



← Olivier's thumb...
the easiest coronagraph
Doesn't work well enough to
see planets around other stars

We need a better coronagraph... and a larger eye (telescope)

What is light: particle or wave ?



1807: Thomas Young publishes his double-slit experiment result ... cannot be explained by Newton's corpuscular theory of light

1818: French academy of science committee launches a competition to explain nature of light

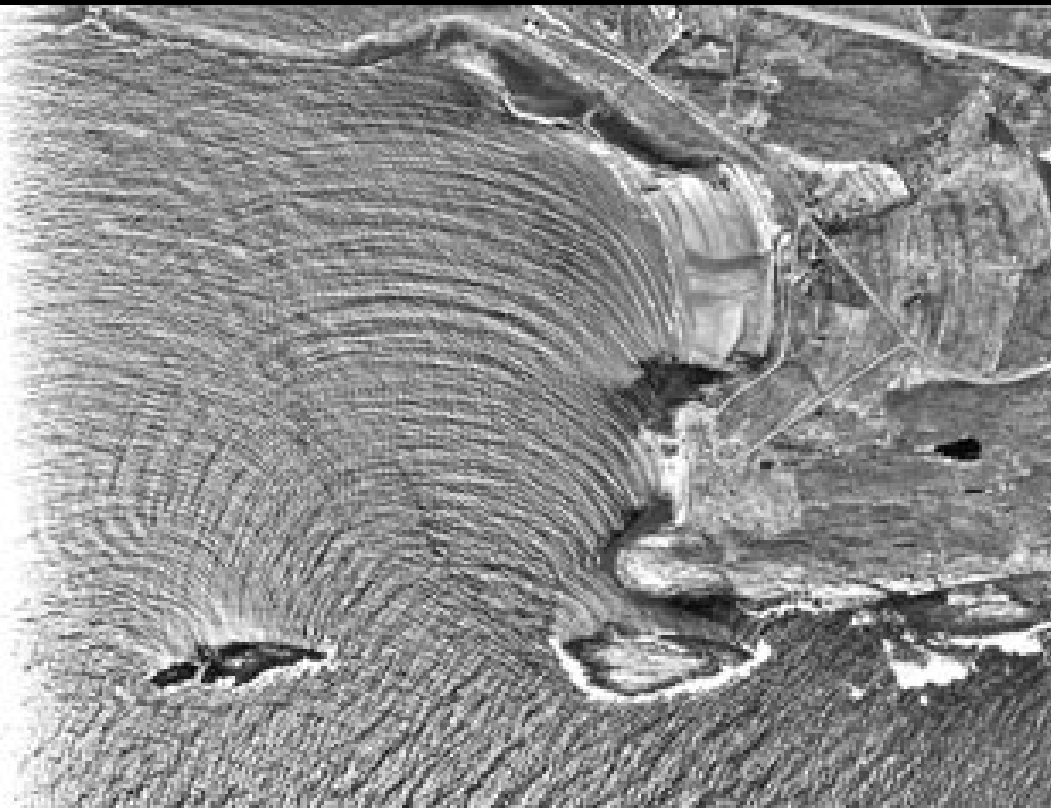


Augustin-Jean Fresnel
submits wave theory of light

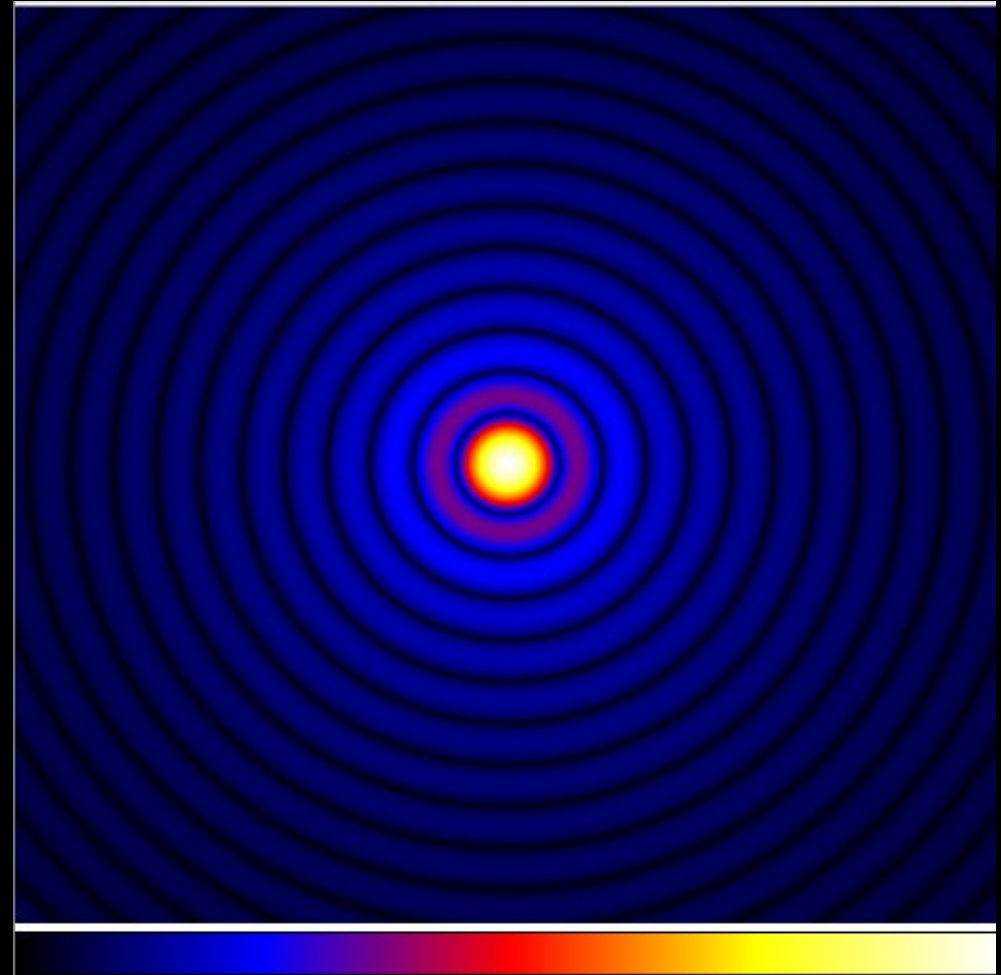
Simeon-Denis Poisson finds a flaw in Fresnel's theory:
According to Fresnel's equations, a bright spot should appear in
the shadow of a circular obstacle → this absurd result disproves Fresnel's theory

Dominique-Francois-Jean Arago, head of the committee, performs the experiment
He finds the predicted spot → Fresnel wins the competition

Water waves diffract around obstacles, edges, and so does light



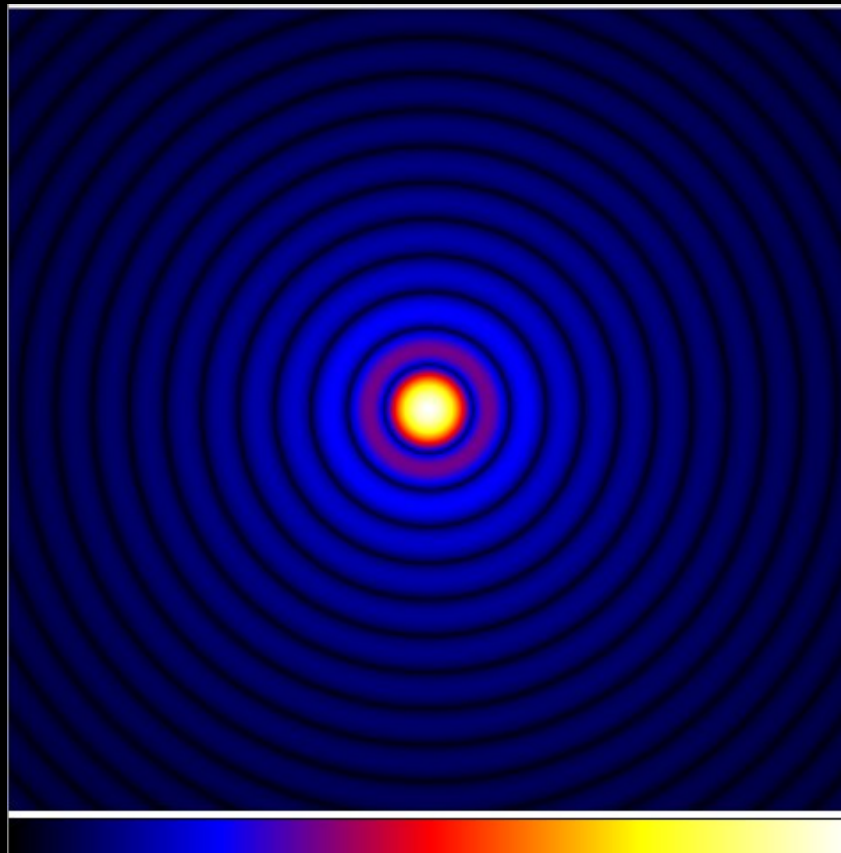
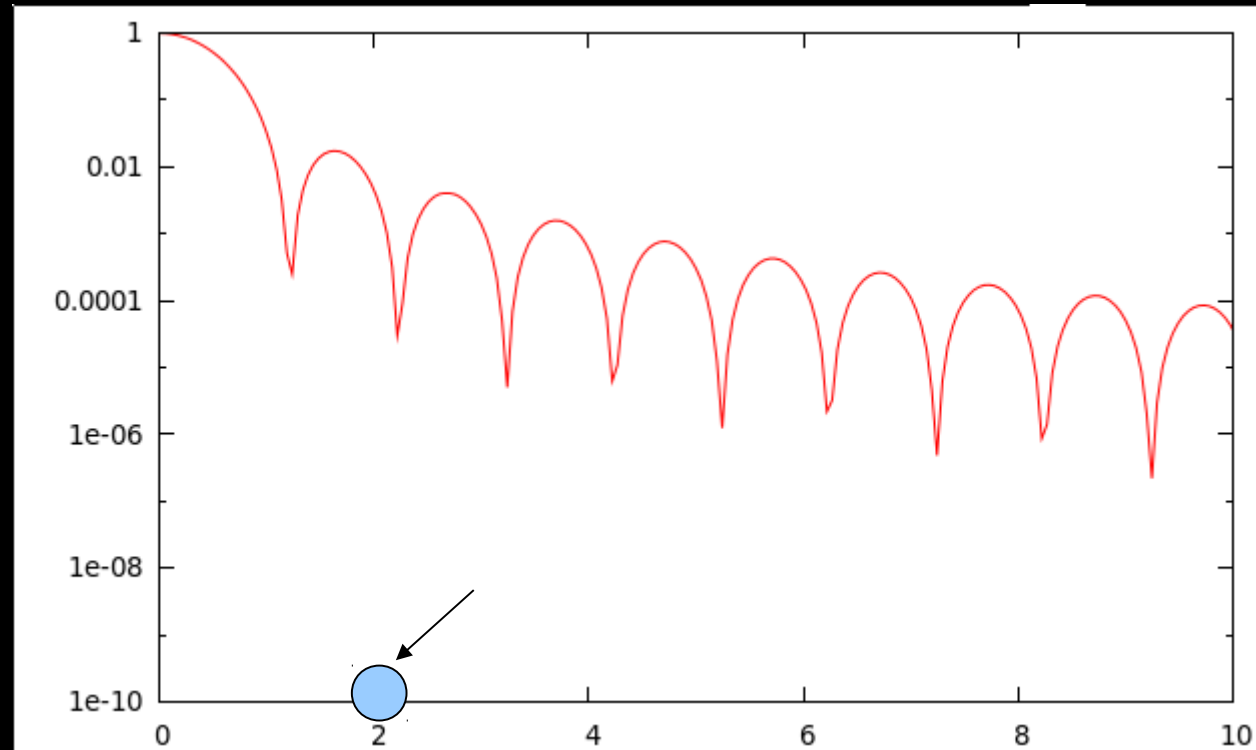
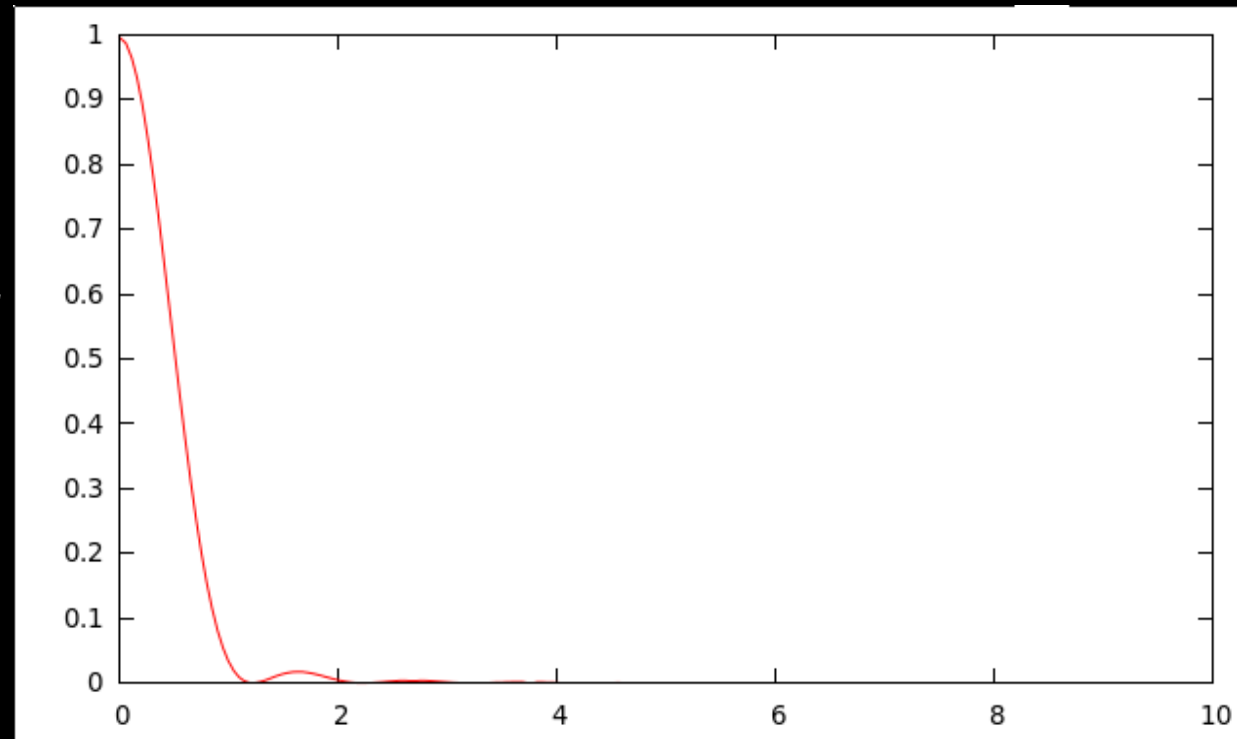
Waves diffracted by coastline and islands



Ideal image of a distant star by a telescope
Diffraction rings around the image core

Why coronagraphy ?

Conventional imaging systems are not suitable for high contrast (even if perfect) due to diffraction



Conventional Pupil Apodization (CPA)

Many pupil apodizations have been proposed.

Apodization can be continuous or binary.

- + Simple, robust, achromatic
- low efficiency for high contrast

Jacquinet & Roisin-Dossier 1964
Kasdin et al. 2003, ApJ, 582, 1147
Vanderbei et al. 2003, ApJ, 590, 593
Vanderbei et al. 2003, ApJ, 599, 686
Vanderbei et al. 2004, ApJ, 615, 555

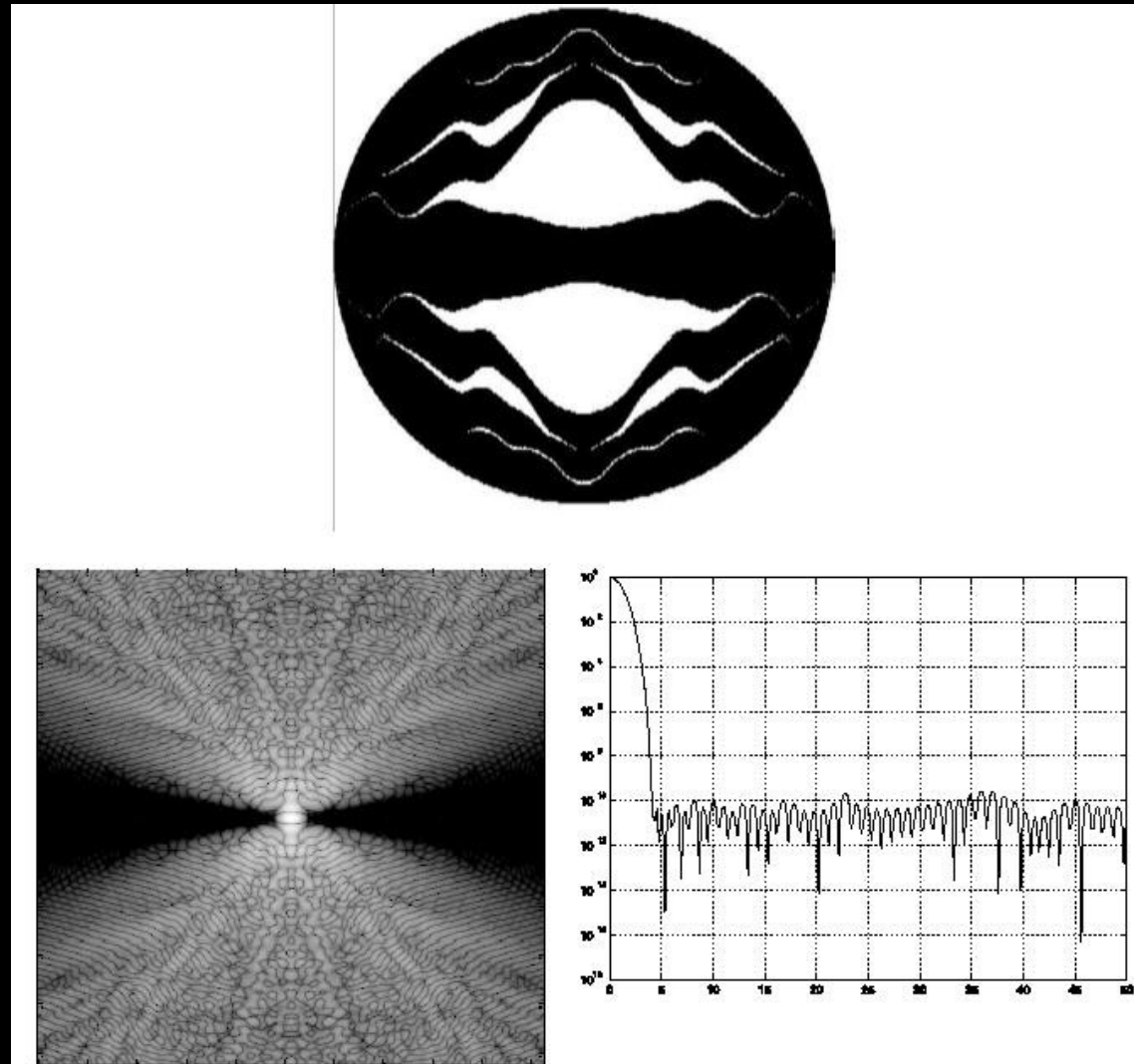
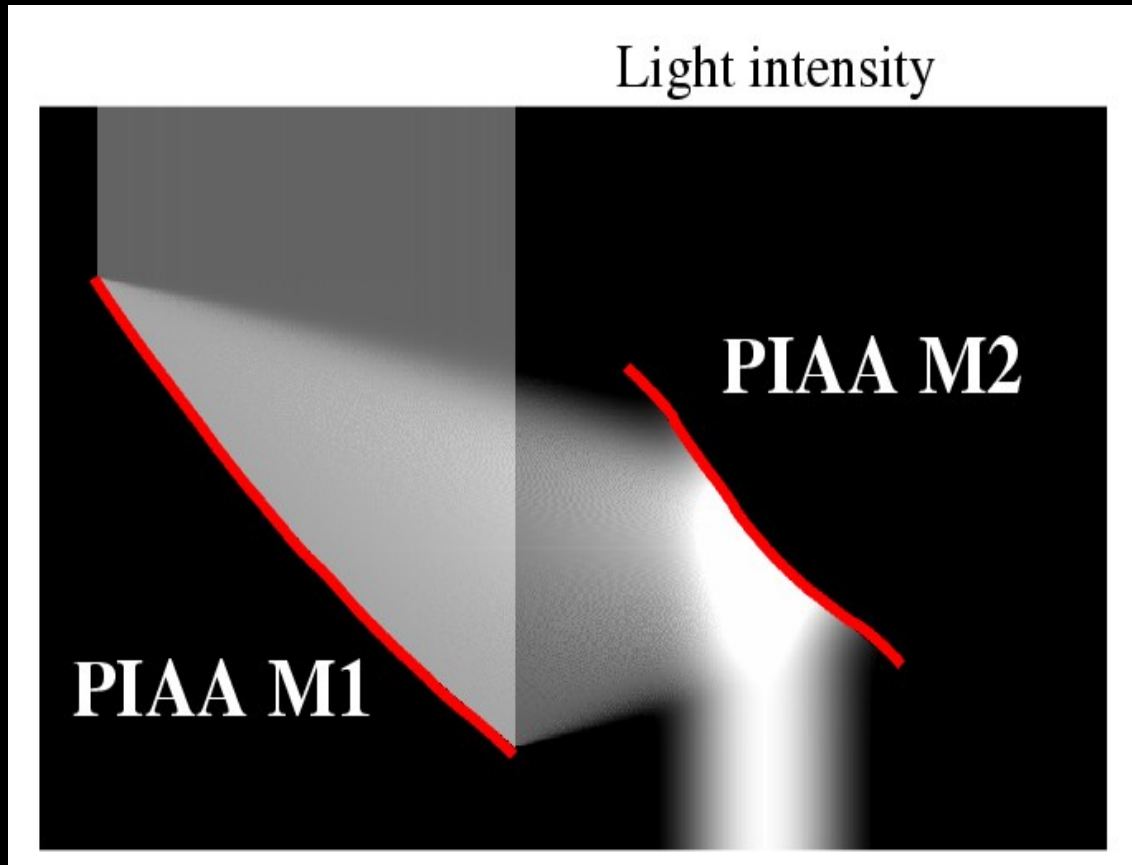


FIG. 9.—*Top:* Asymmetric multiopening mask designed to provide high-contrast, 10^{-10} , from $\lambda/D = 4$ to $\lambda/D = 100$ in two angular sectors centered on the x -axis. Ten integrations are required to cover all angles. Total throughput and pseudoarea are 24.4%. Airy throughput is 11.85%. *Bottom:* Associated PSF. (Note that this mask was originally designed for an elliptical mirror. It has been rescaled to fit a circular aperture.)

Phase-Induced Amplitude Apodization Coronagraph (PIAAC)

Lossless apodization by aspheric optics.

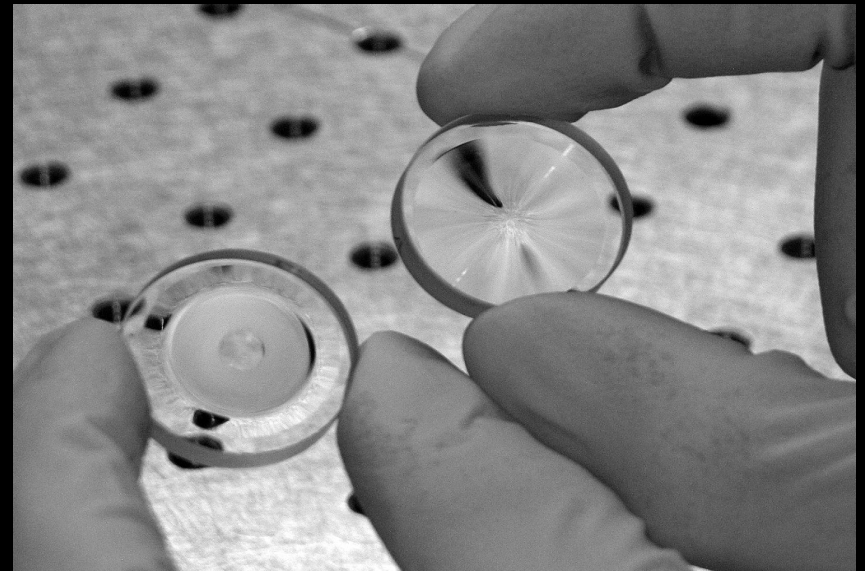
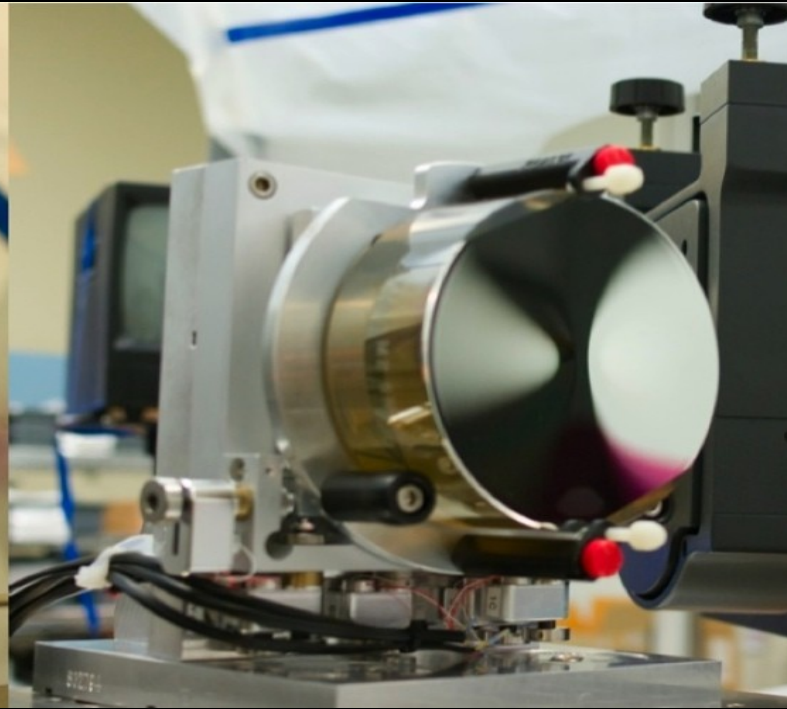
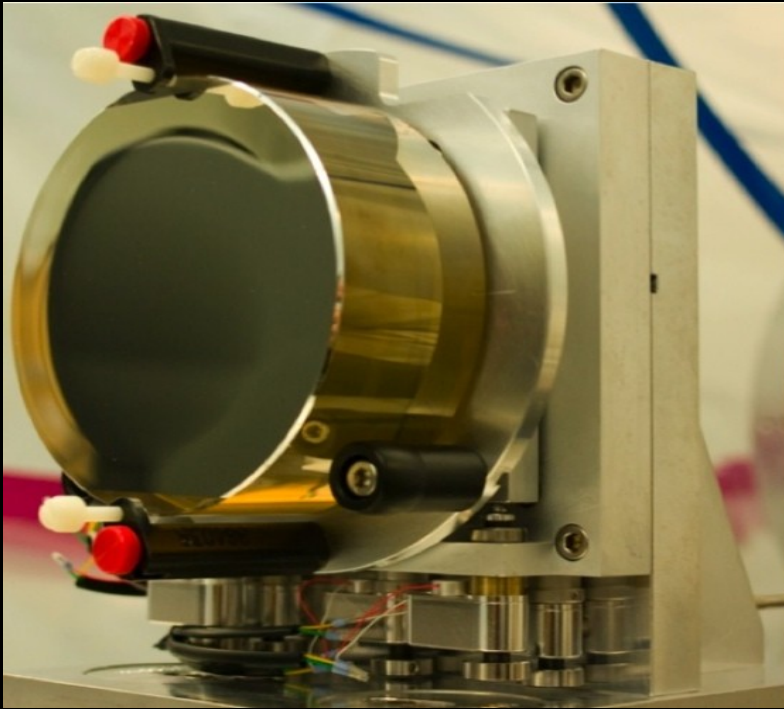


No loss in angular resolution or sensitivity
Achromatic (with mirrors)
Small inner working angle

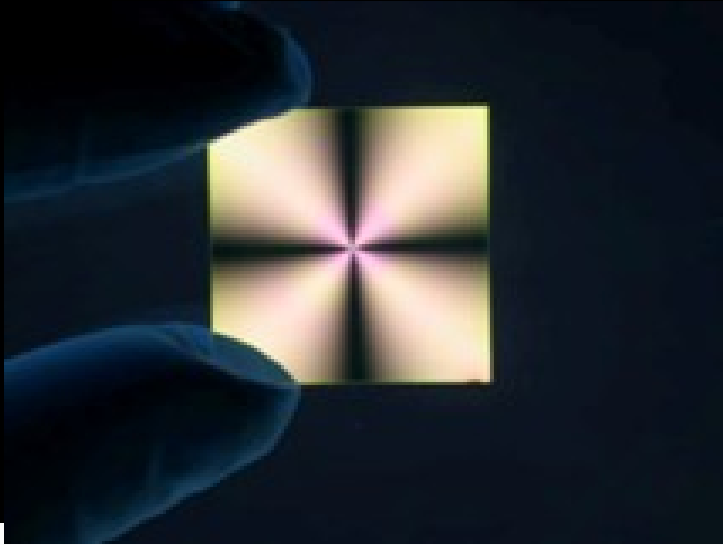
→ Gain $\sim x2$ to $x3$ in telescope diameter over previous concepts

Guyon, Belikov, Pluzhnik, Vanderbei, Traub, Martinache ... 2003-present

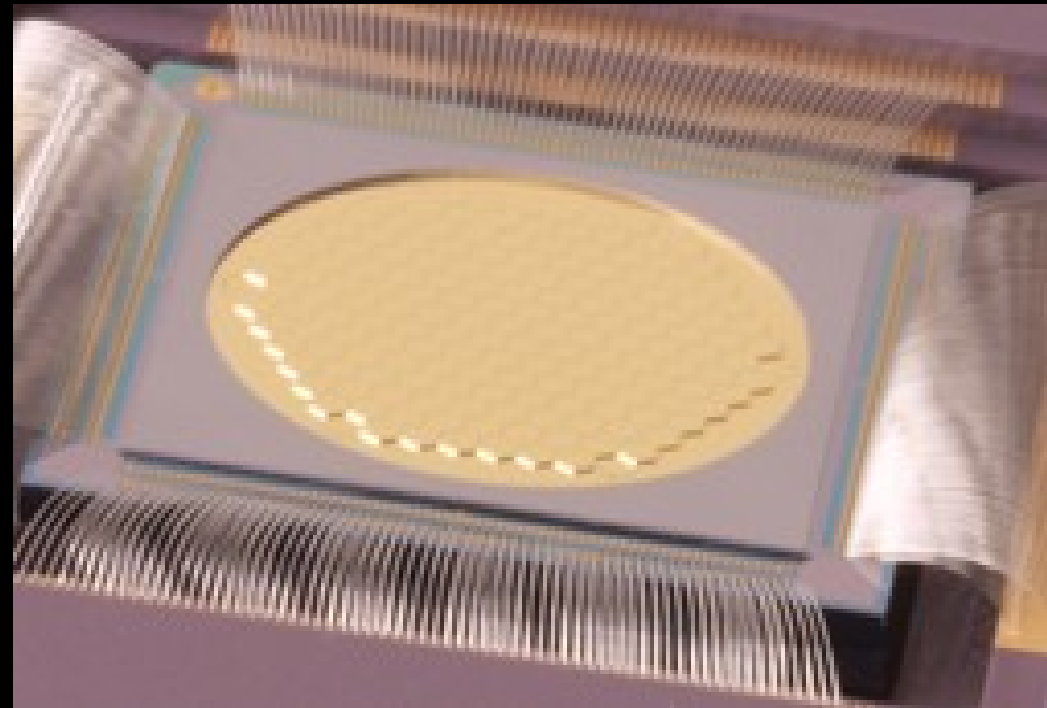
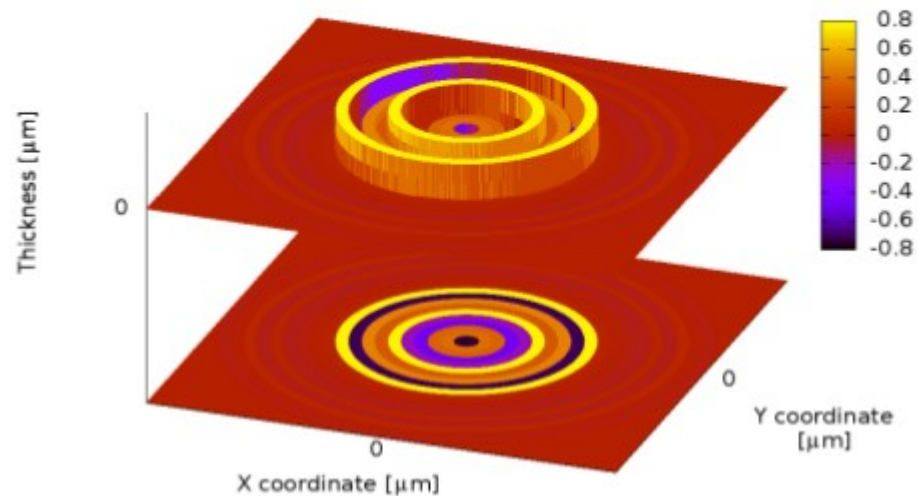
We use strangely shaped optics to reshape light



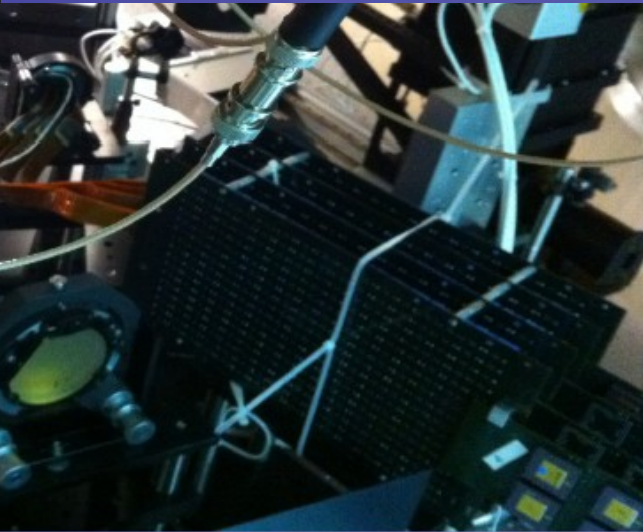
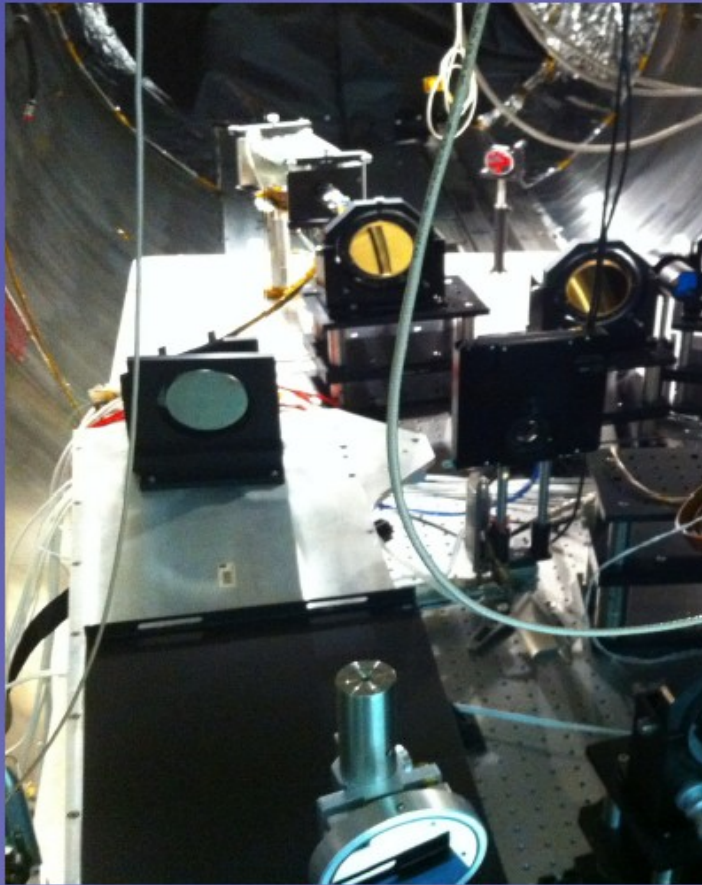
Technology: components



PIAACMC optimized focal plane mask
F/20 beam, 10% bandwidth around 0.5 μm
SiO₂, 20 zones, 4 μm max deviation

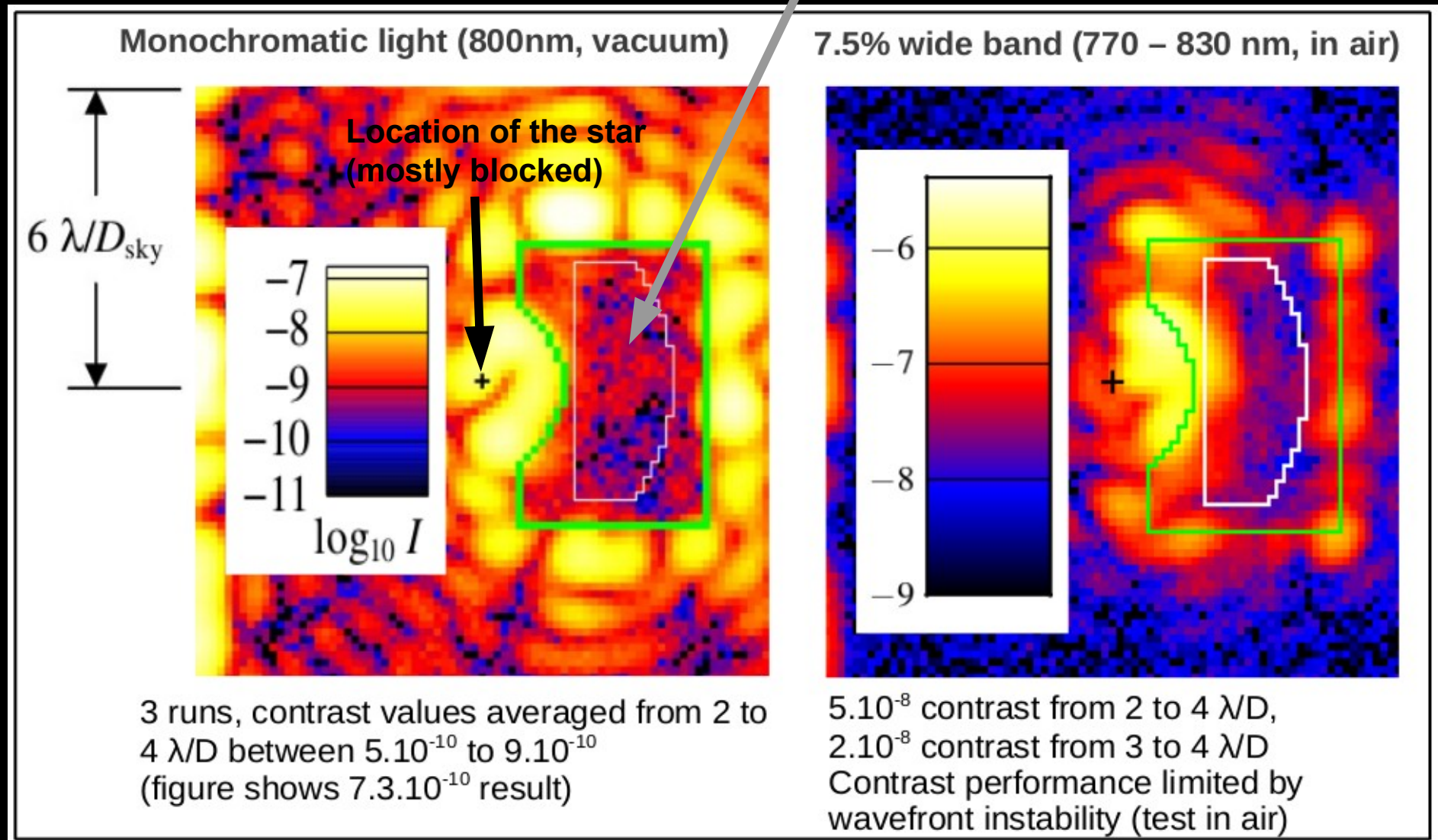


Vacuum testbeds at NASA JPL



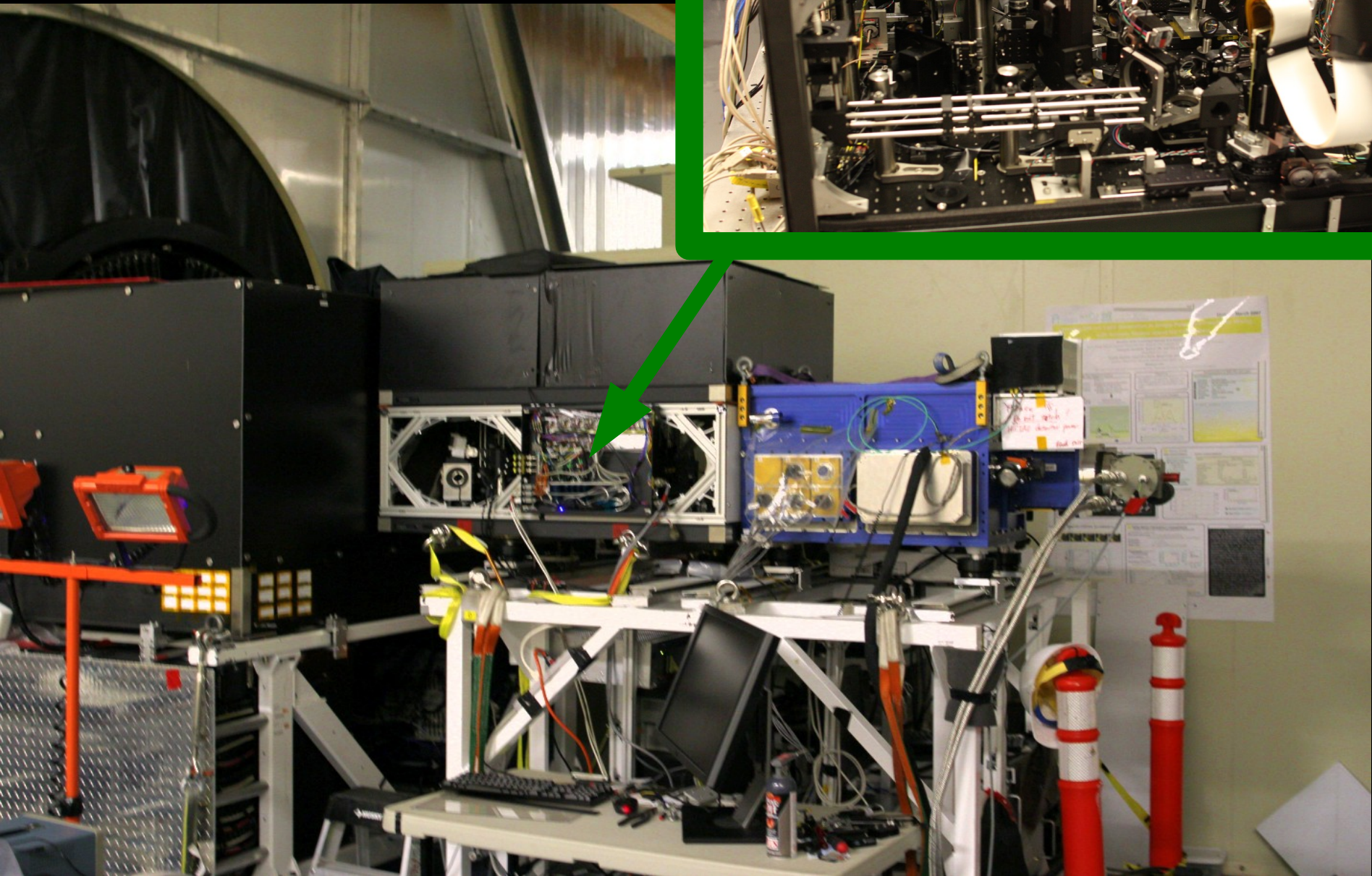
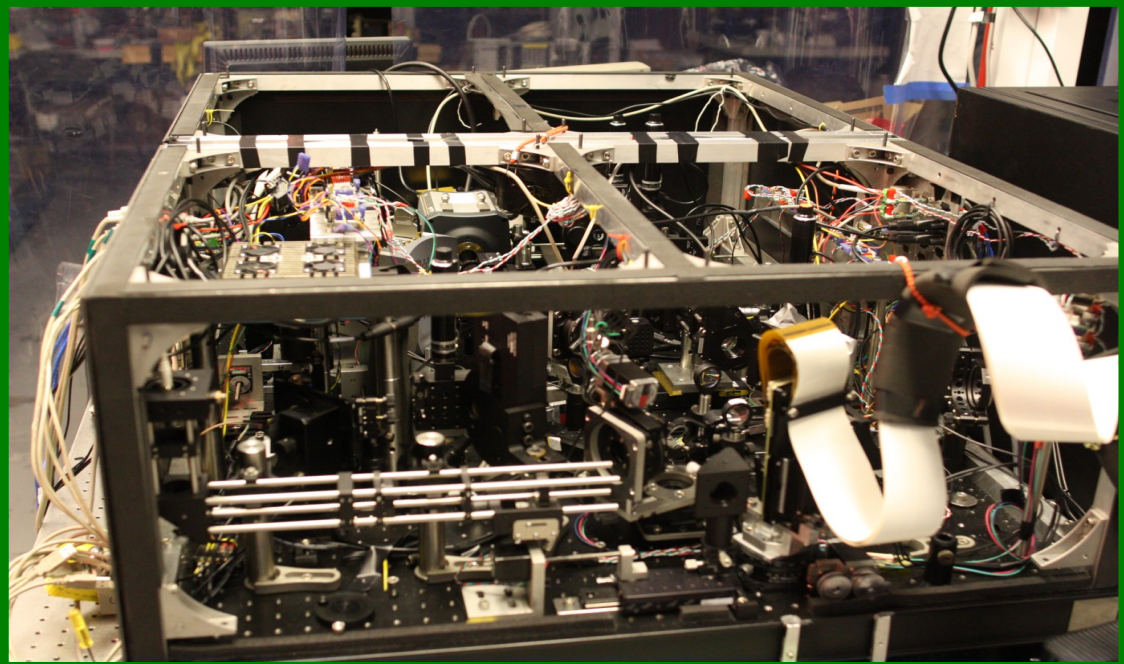
PIAA testbed at NASA JPL : lab results

An Earth-like planets could be seen !

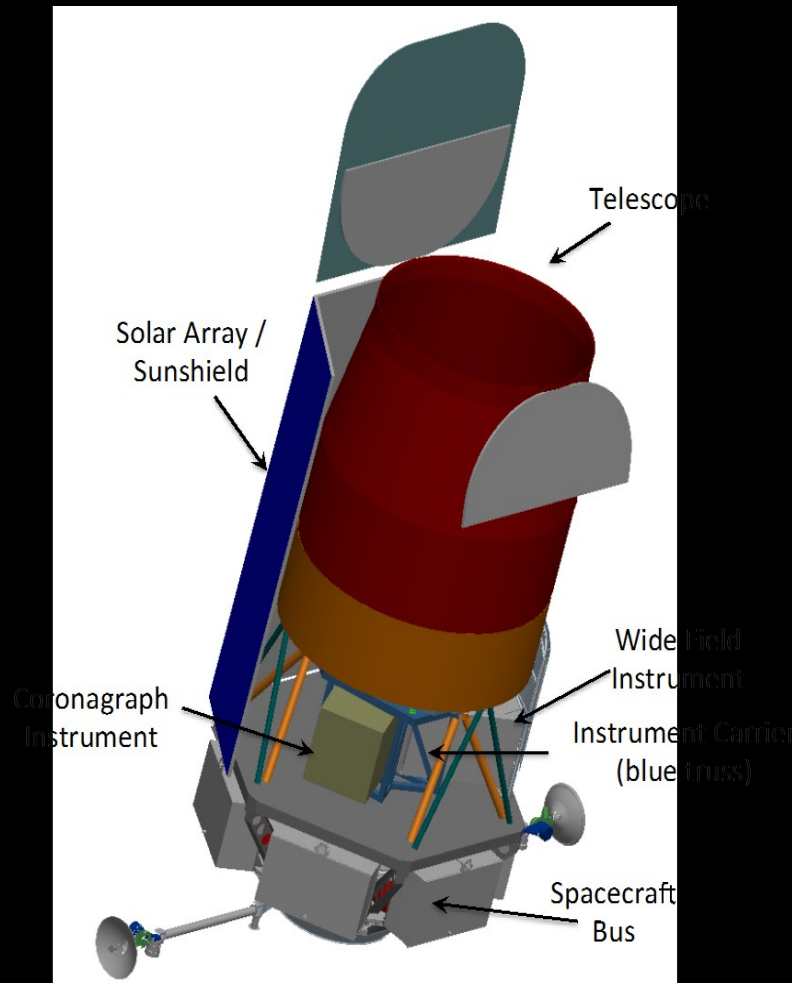




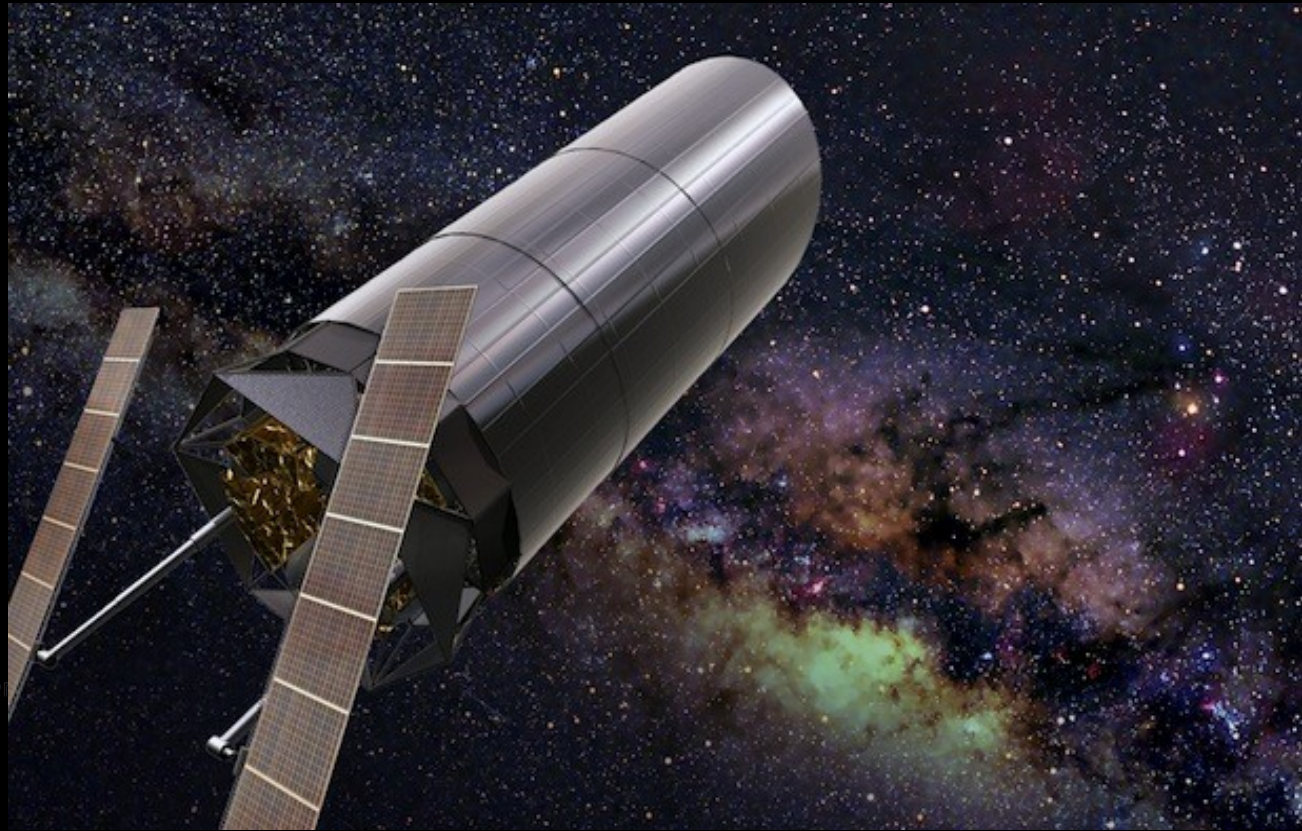
The Subaru Coronagraphic Extreme Adaptive Optics (SCExAO) system



Space telescopes



WFIRST-AFTA telescope, 2.4m diameter

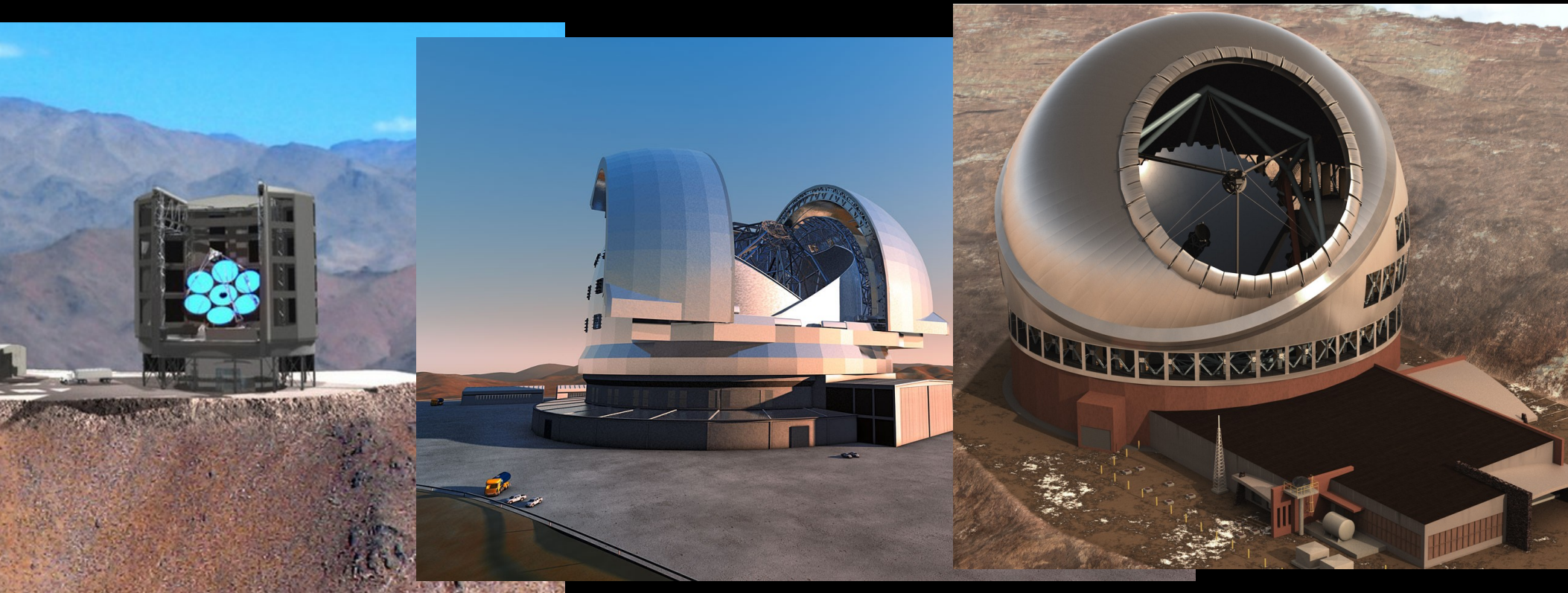


ATLAS 8m telescope project

Exciting future opportunities

Next generation of large telescopes on the ground will be able to image habitable planets around nearby low mass red stars
3 projects, ~30m diameter

Space telescopes with coronagraphs will be able to image and study Earth-like planets around sun-like stars

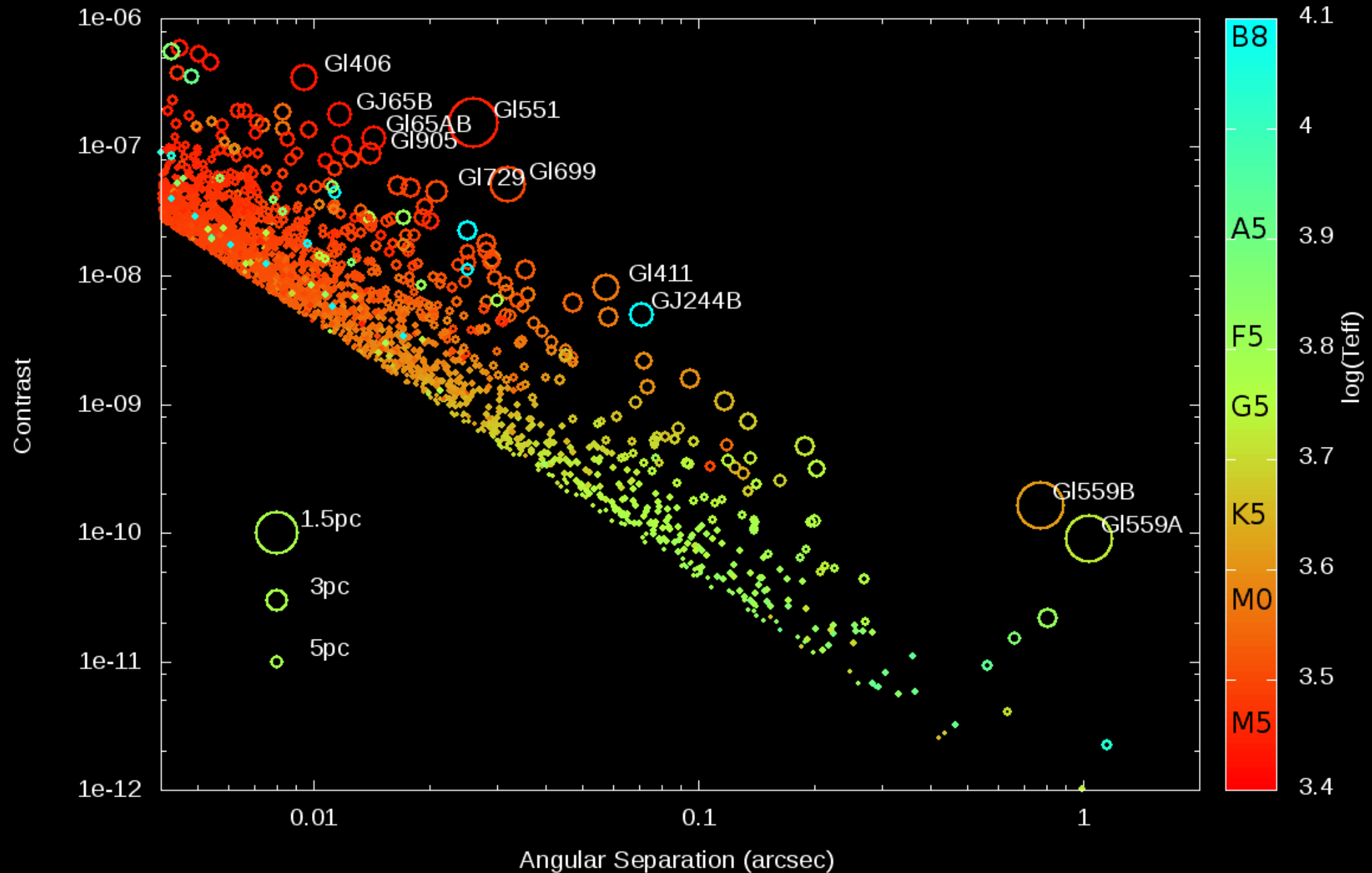


Giant Magellan Telescope

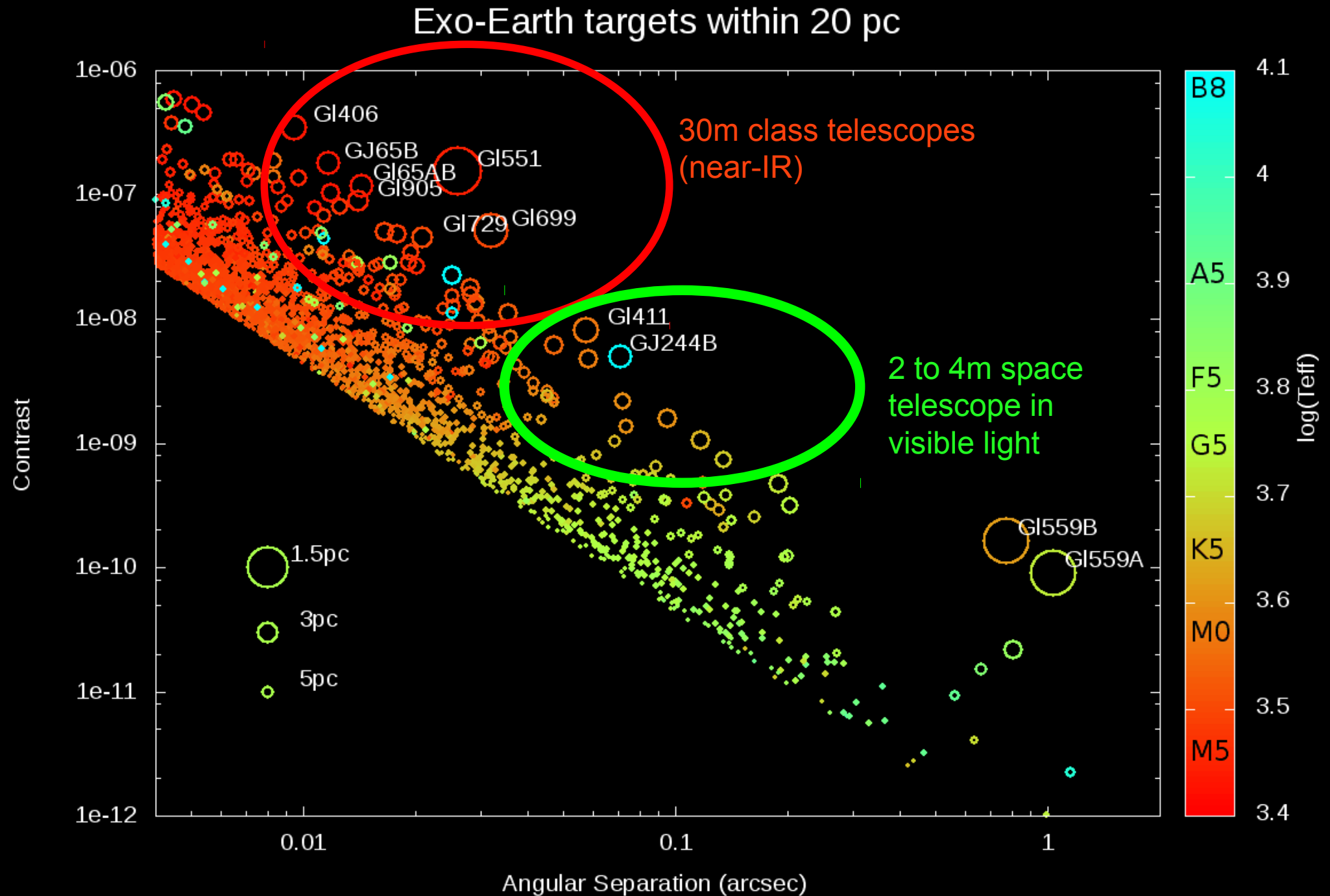


Detecting planets from space and ground

Exo-Earth targets within 20 pc



Detecting planets from space and ground



Proxima Centauri



Sun

Alpha Centauri A

Alpha Centauri B



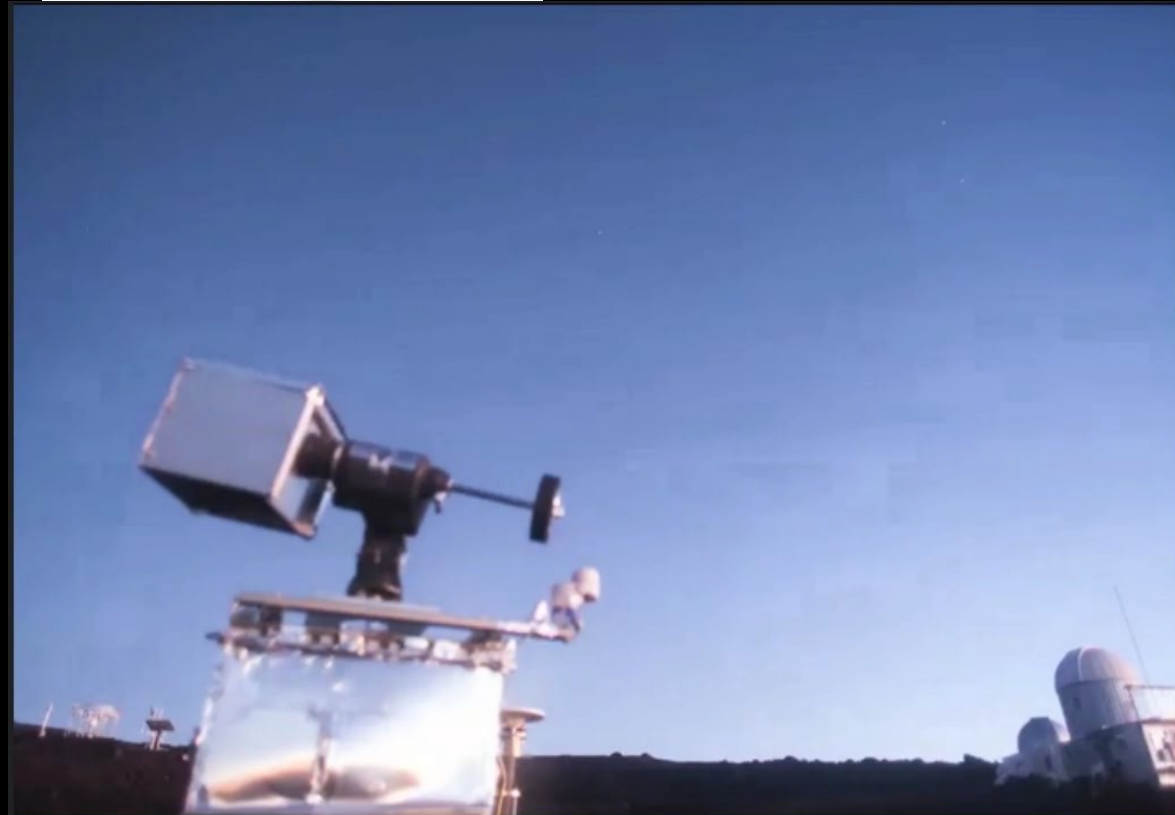
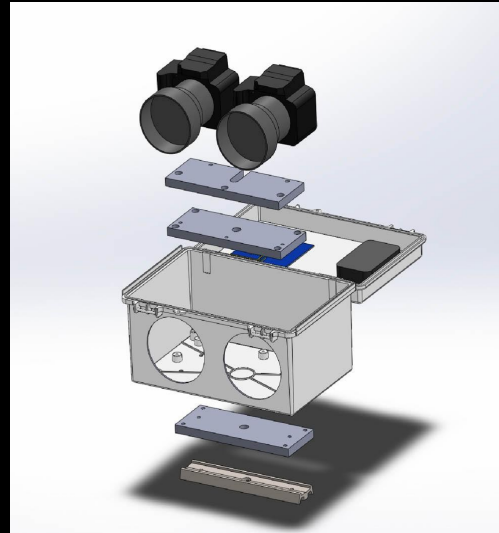
Proxima Centauri

PANOPTES Panoptic Astronomical Networked Optical observatory for Transiting Exoplanets Survey

Discovering transiting exoplanets requires monitoring large parts of the sky for long periods of time

Amateur astronomers, citizen scientists are very good at this, and schools can participate with student team projects

We are developing an easy to assemble, low cost system based on digital SLR cameras to do exoplanet transit survey work

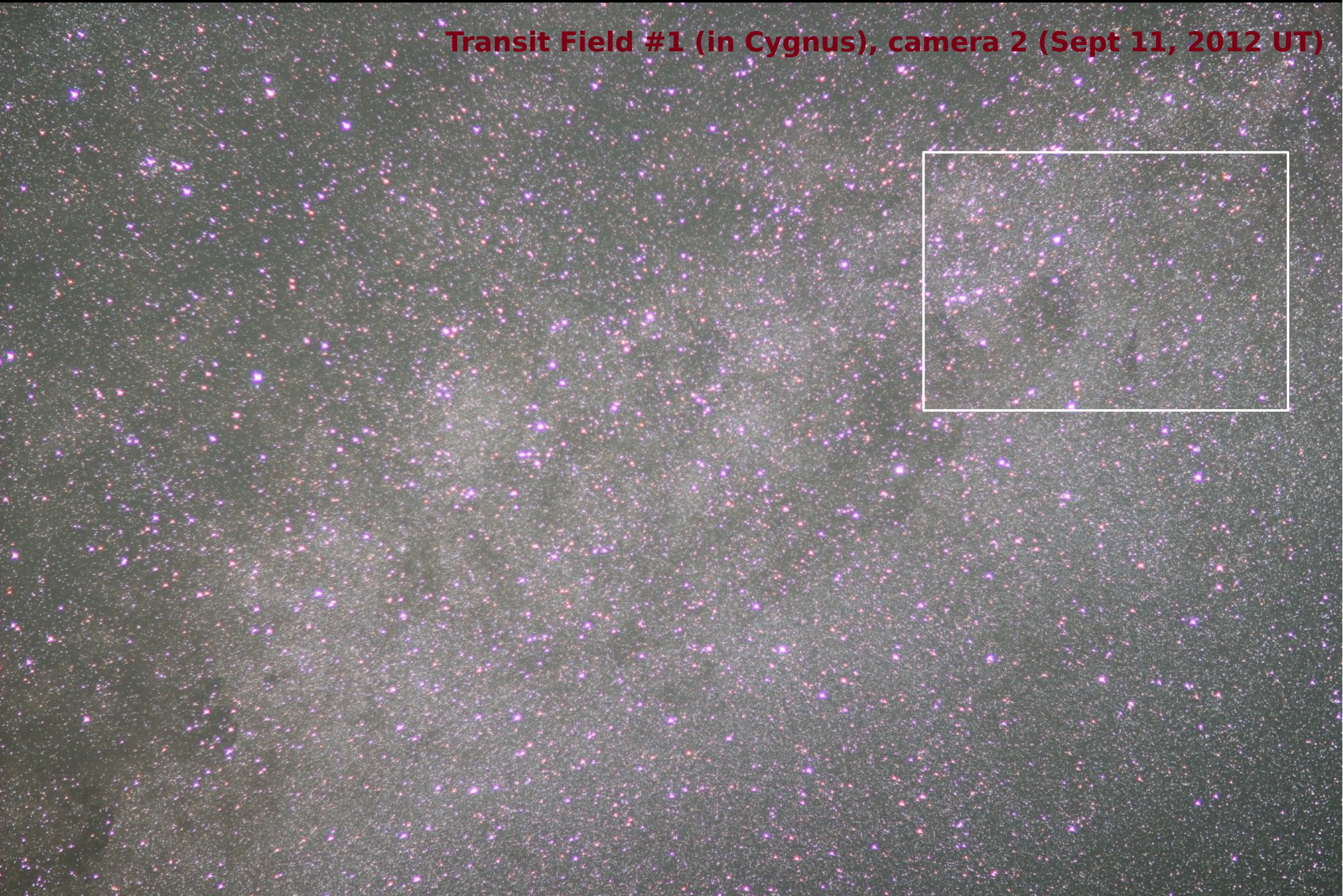


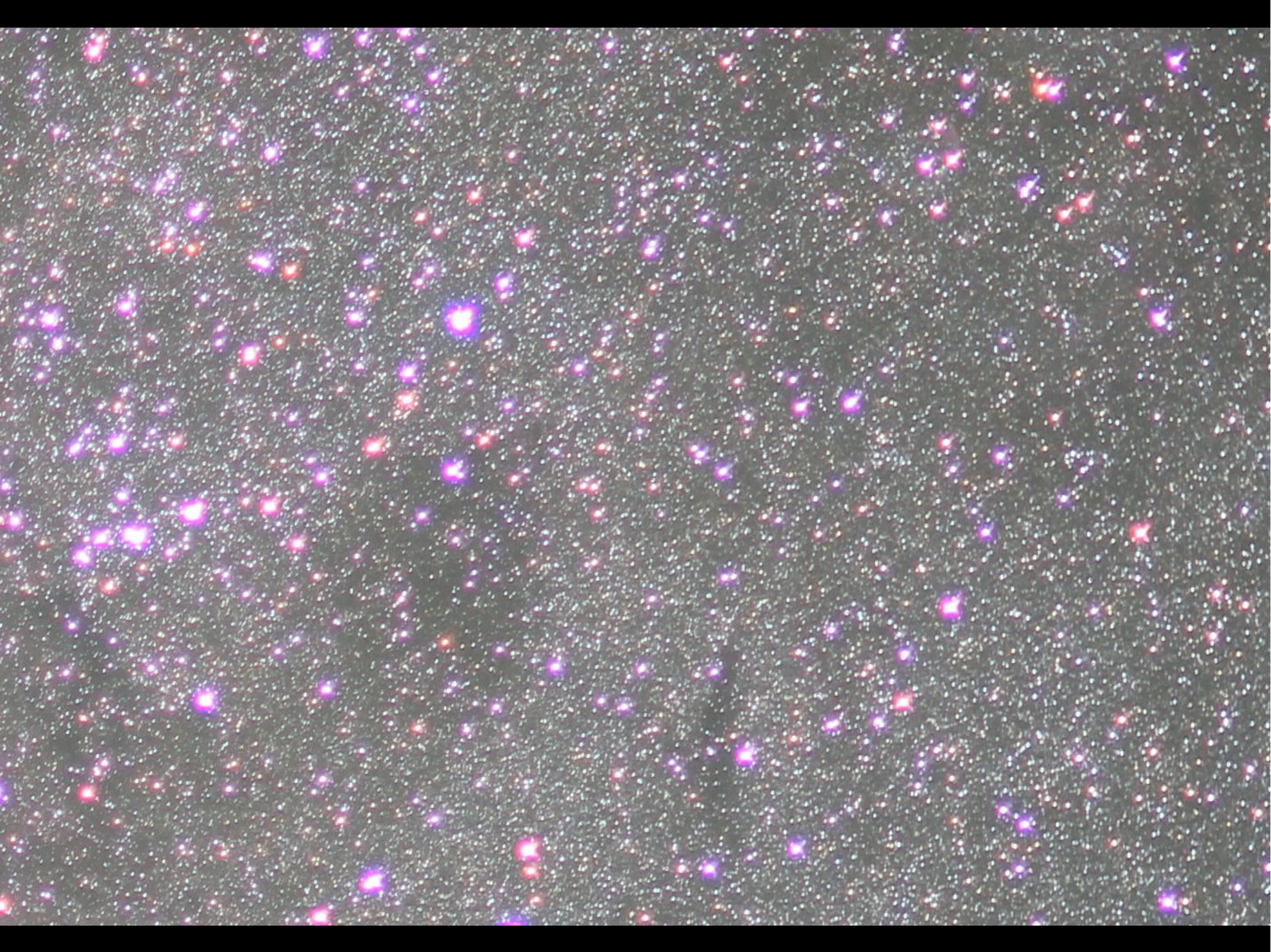
PANOPTES prototype #3 unit at Mauna Loa observatory (May 19, 2013)



**Example image (Cygnus field):
>100,000 stars in a single image**

Transit Field #1 (in Cygnus), camera 2 (Sept 11, 2012 UT)





More information, local resources

College of Optical Sciences/Steward Observatory/Lunar and Planetary Laboratory

Astrobiology at University of Arizona: <http://astrobiology.arizona.edu/>

Mirror laboratory: <http://mirrorlab.as.arizona.edu/>

Missions/Telescopes:

Large Binocular Telescope: <http://www.lbto.org/>

<http://lbt.as.arizona.edu/LBTI>

Giant Magellan Telescope: <http://www.gmto.org/>

OSIRIX-REX: <http://www.asteroidmission.org/>

NASA:

<http://exep.jpl.nasa.gov/>

Project PANOPTES: <http://projectpanoptes.org>

