



Life in the universe

Olivier Guyon [oliv.guyon@gmail.com]

Subaru Telescope, National Astronomical Observatory of Japan

University of Arizona

NASA Jet Propulsion Laboratory

Life on Earth

Started very early (few 100Myr after formation) and “simple”

Stayed at the bacteria level until ~2 Byr ago

Is now very abundant

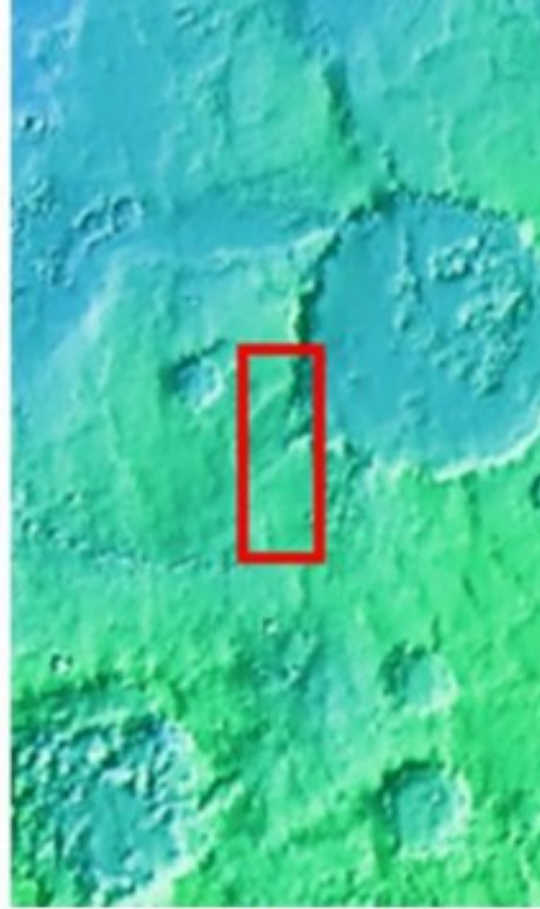
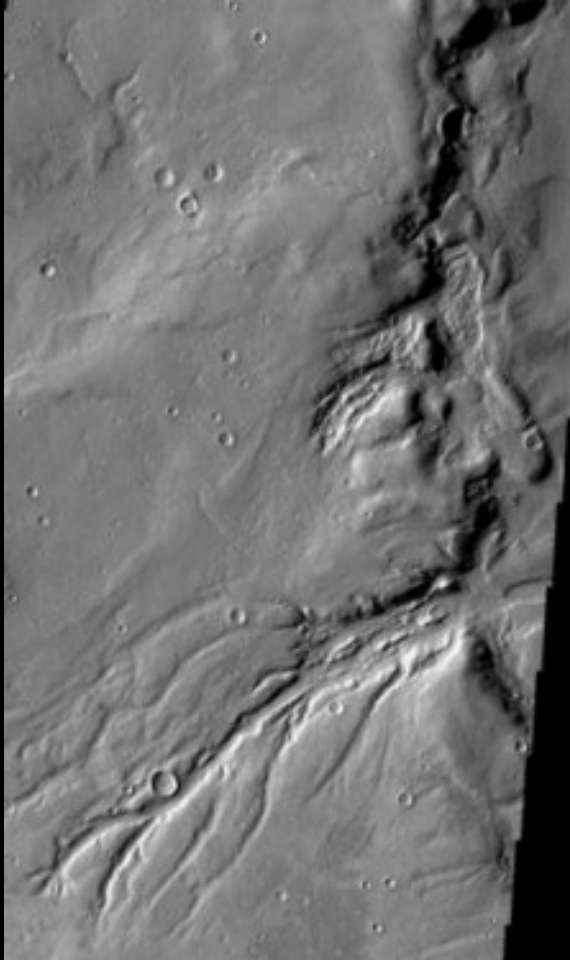
Will decline in few ~100Myr due to increase Sun luminosity

Mars

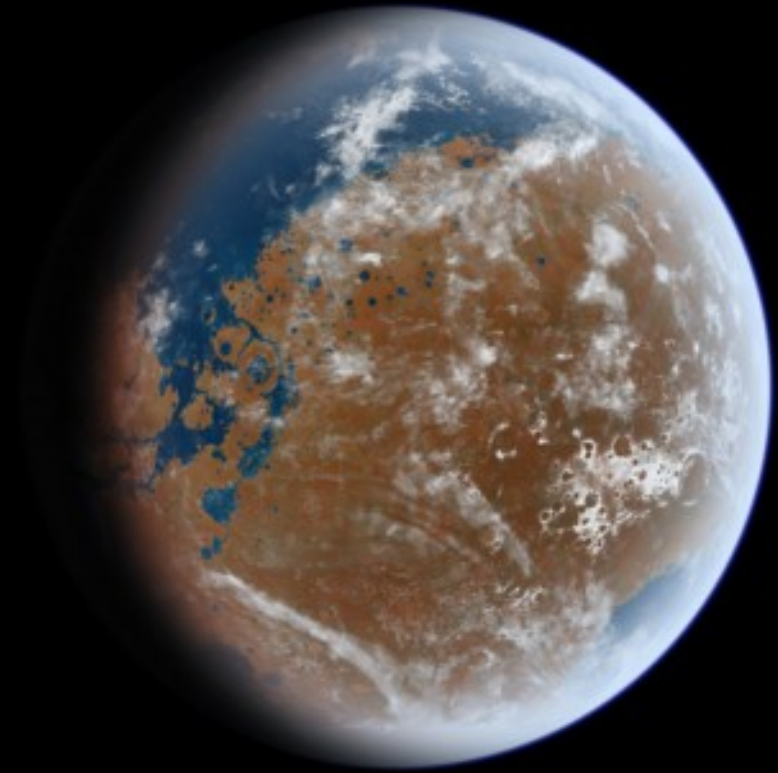


Mars is too cold, dry and its atmosphere too thin to support abundant complex life...
... but may still host bacteria

Mars



Picture on left is inside red box.



Ancient Mars ?

Mars had flowing water and more temperate atmosphere in the past

What makes planets habitable ?

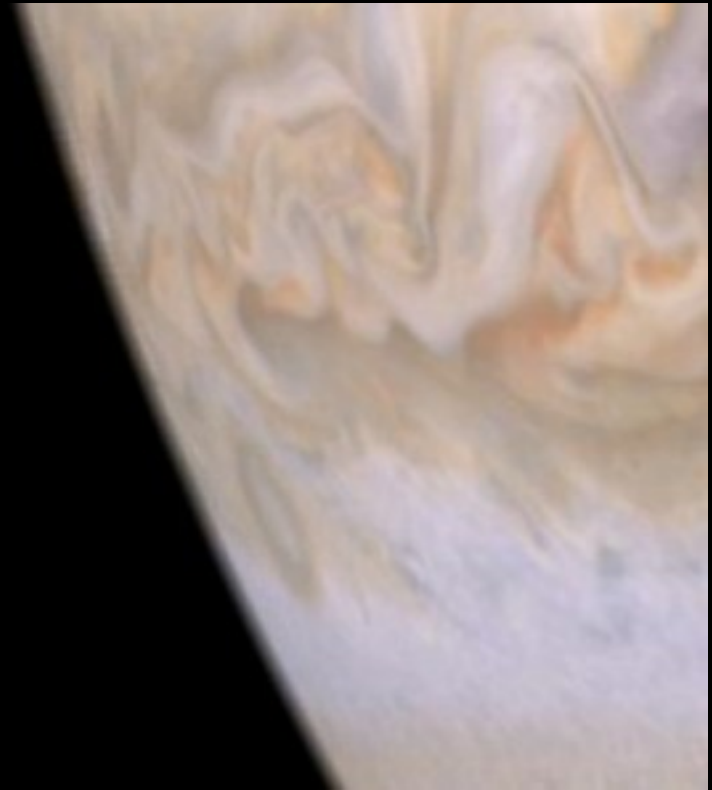
Size matters: strong gravity = thick atmosphere



Moon: too small
No atmosphere



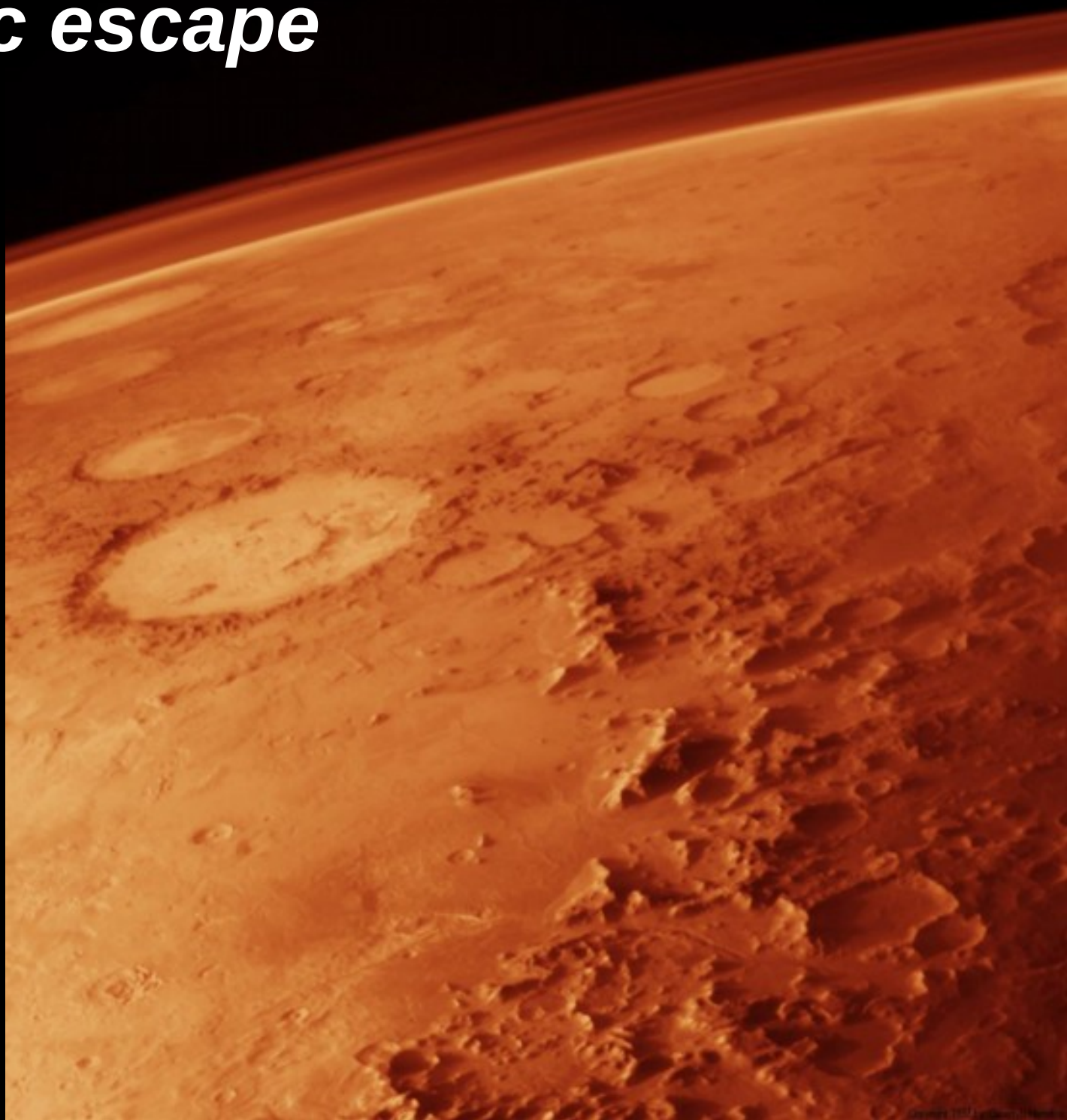
Earth



Jupiter: too big
Mostly gas

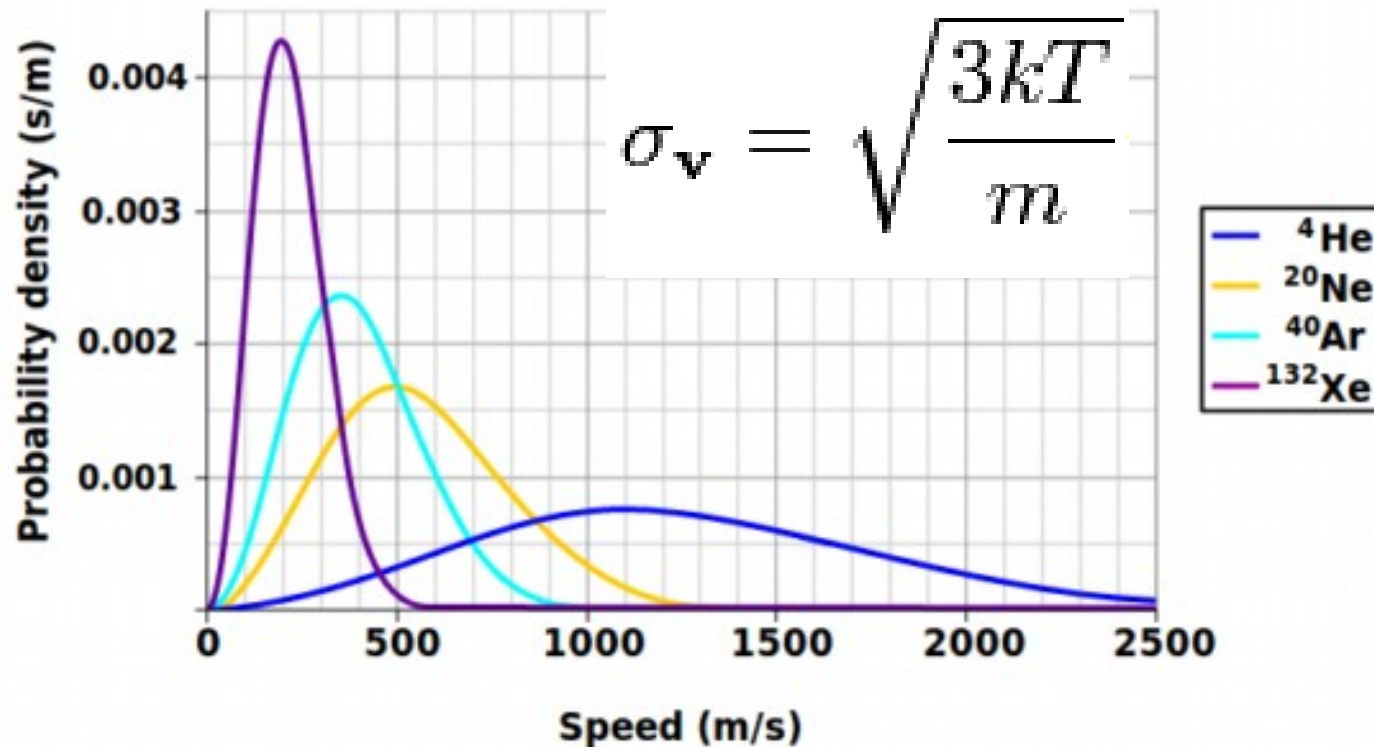
Atmospheric escape

Mars' thin atmosphere



Atmospheric escape

Maxwell-Boltzmann Molecular Speed Distribution for Noble Gases



Escape velocity

$$v_e = \sqrt{\frac{2GM}{r}}$$

60 km/s for Jupiter
11 km/s for Earth
5 km/s for Mars
2.4 km/s for Moon

Gas temperature \rightarrow molecules/atoms motion

Some molecules and gas will have speed $>$ escape velocity, and will leave the planet

What makes planets habitable ?

The planet must be in the habitable zone of its star: not be too close or too far



Venus: too close, too hot



Image credit: NASA/JPL-Caltech/MSSS

Mars: too far, too cold

Our Sun – a few facts

MASS

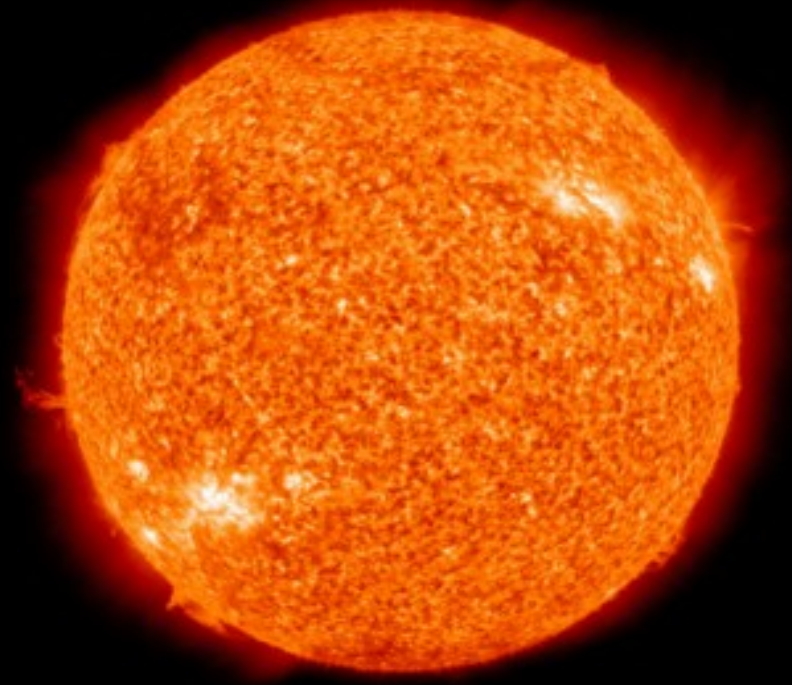
99.86% of the solar system mass
(this is why planets go around the Sun)

SIZE

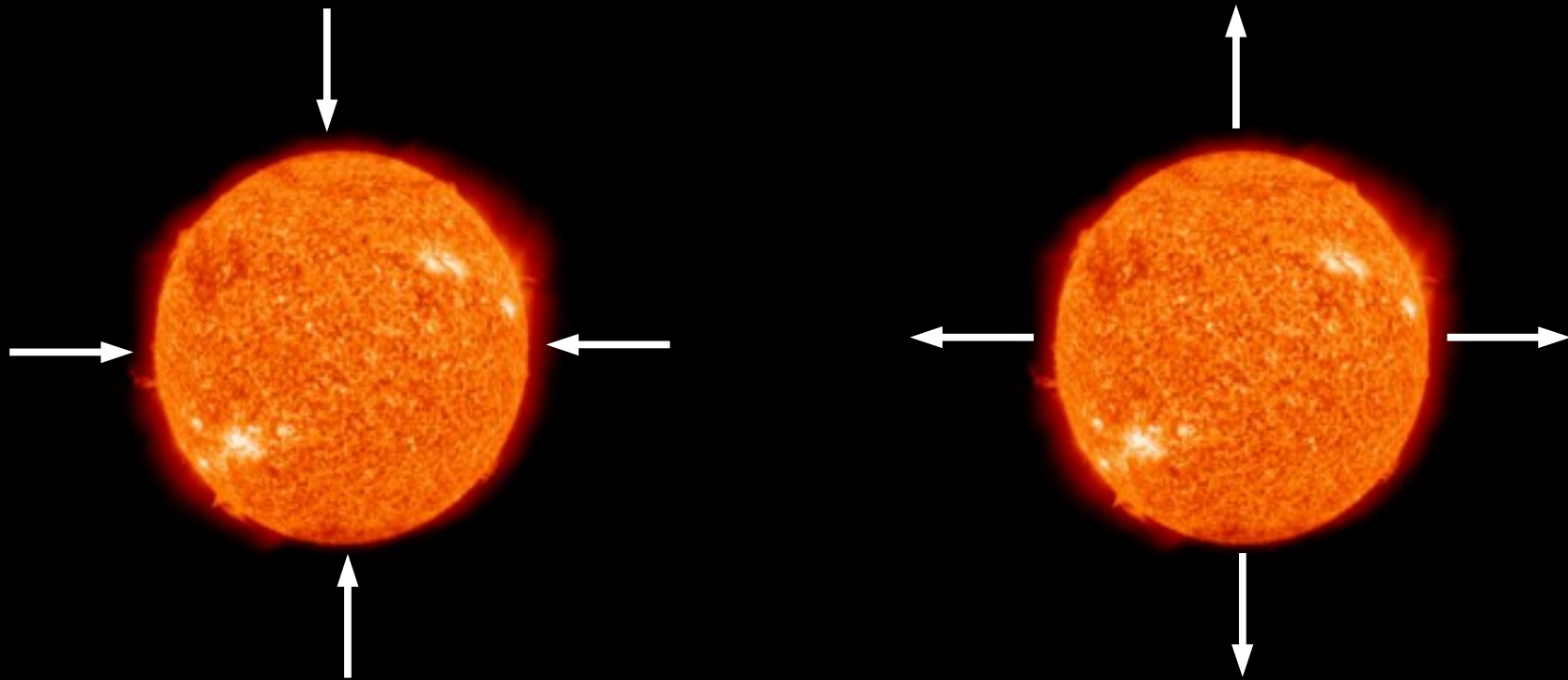
1.4 million km diameter (~100x Earth)
Jet airplane would take few months to go around Sun

Mostly Hydrogen, a bit of Helium
Surface temperature: 5505 C

Earth receives ~1.4 kW of energy (in the form of light) per square meter from the Sun



Gravity vs Thermonuclear fusion

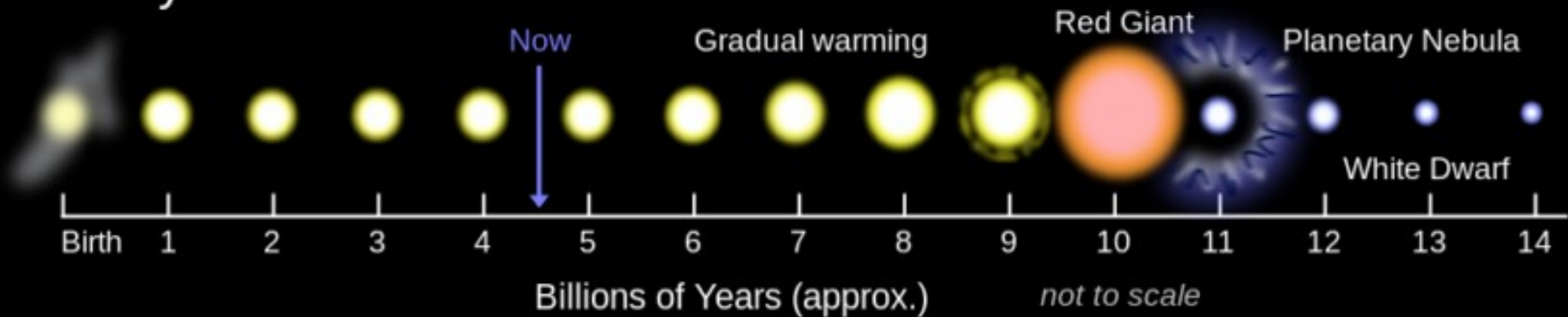


***When gravity compresses the Sun core, pressure at center increases → thermonuclear fusion gets much more efficient
→ Currently in equilibrium: neither force can win***

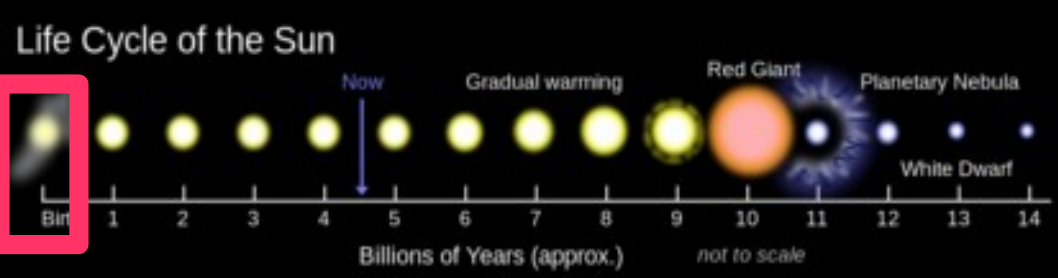
***This equilibrium lasts for ~10 billion years for the Sun
(this phase is called the “Main Sequence”)***

Sun life cycle

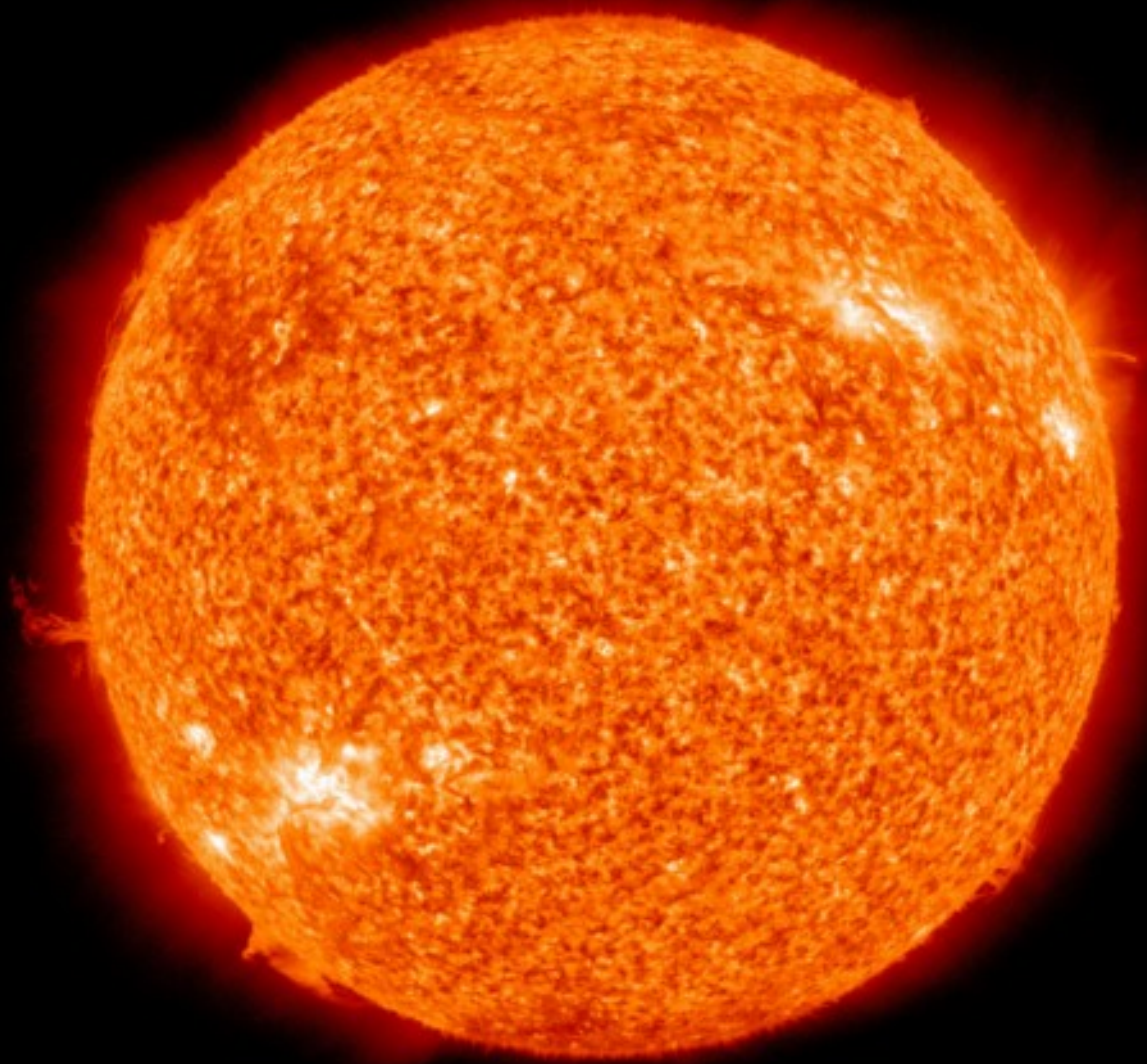
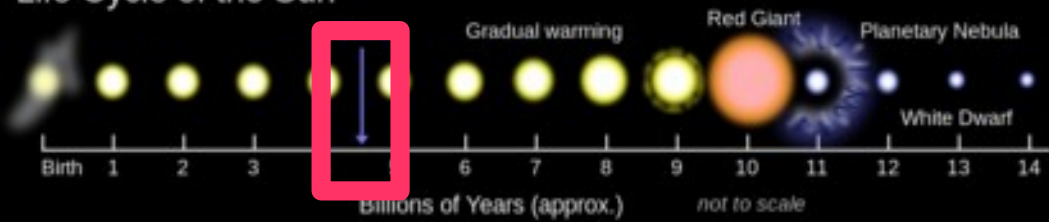
Life Cycle of the Sun

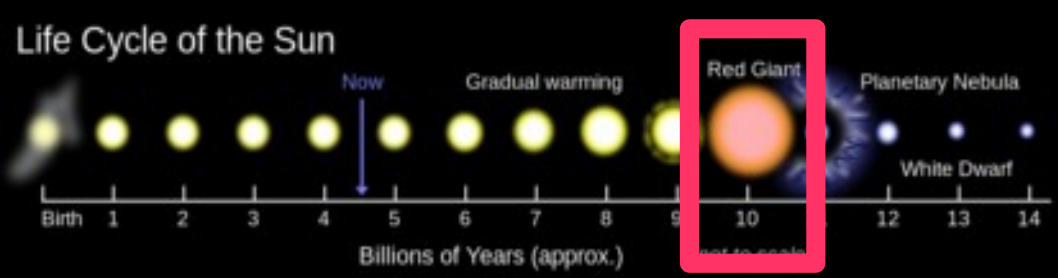


10 billion yrs of “stable life”



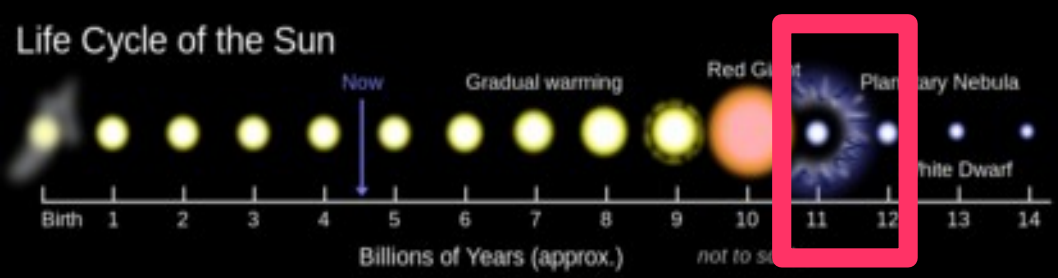
Life Cycle of the Sun



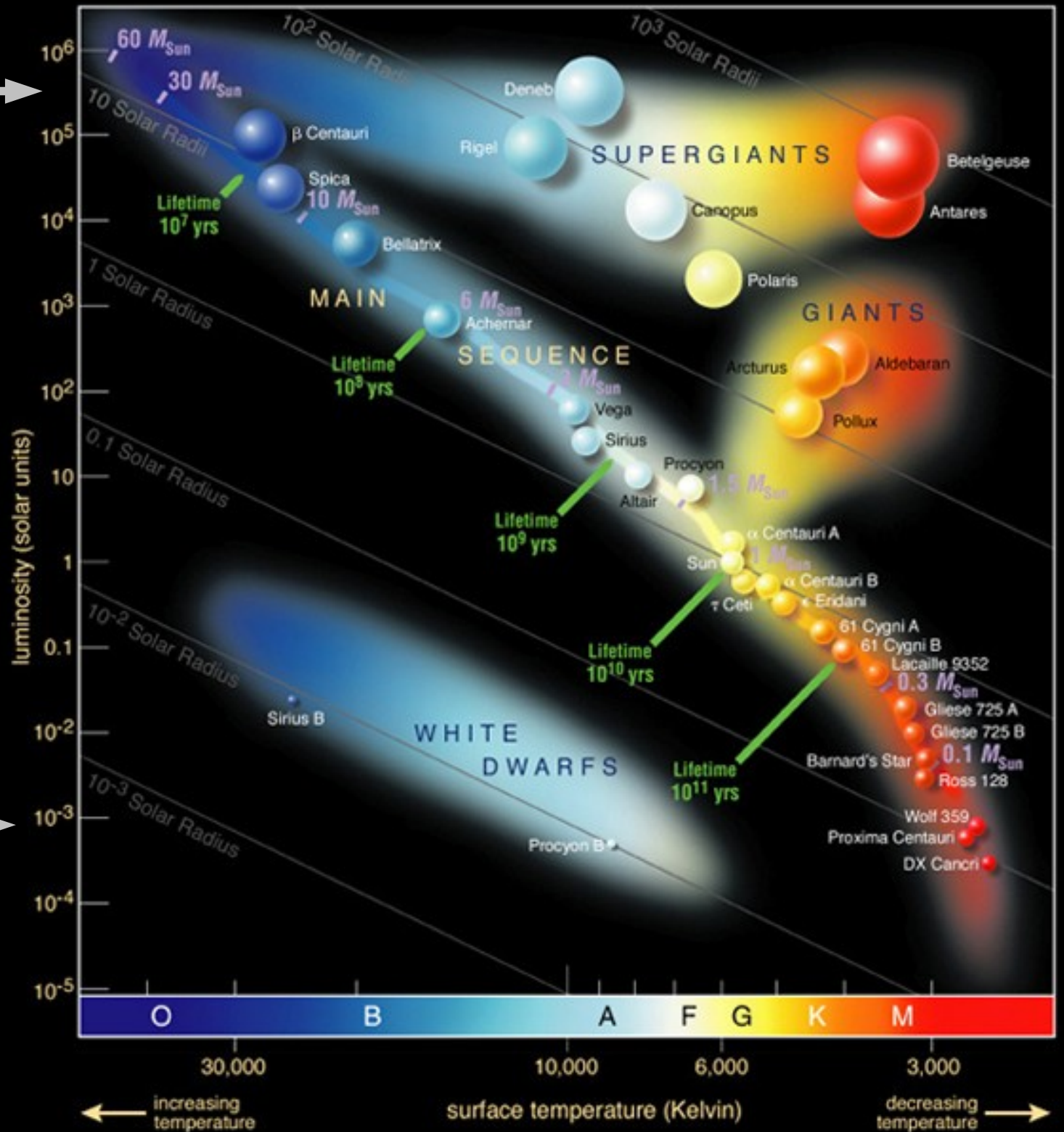


The Sun as a red giant
(diameter ≈ 2 AU)

The Sun as a main-sequence star
(diameter ≈ 0.01 AU)

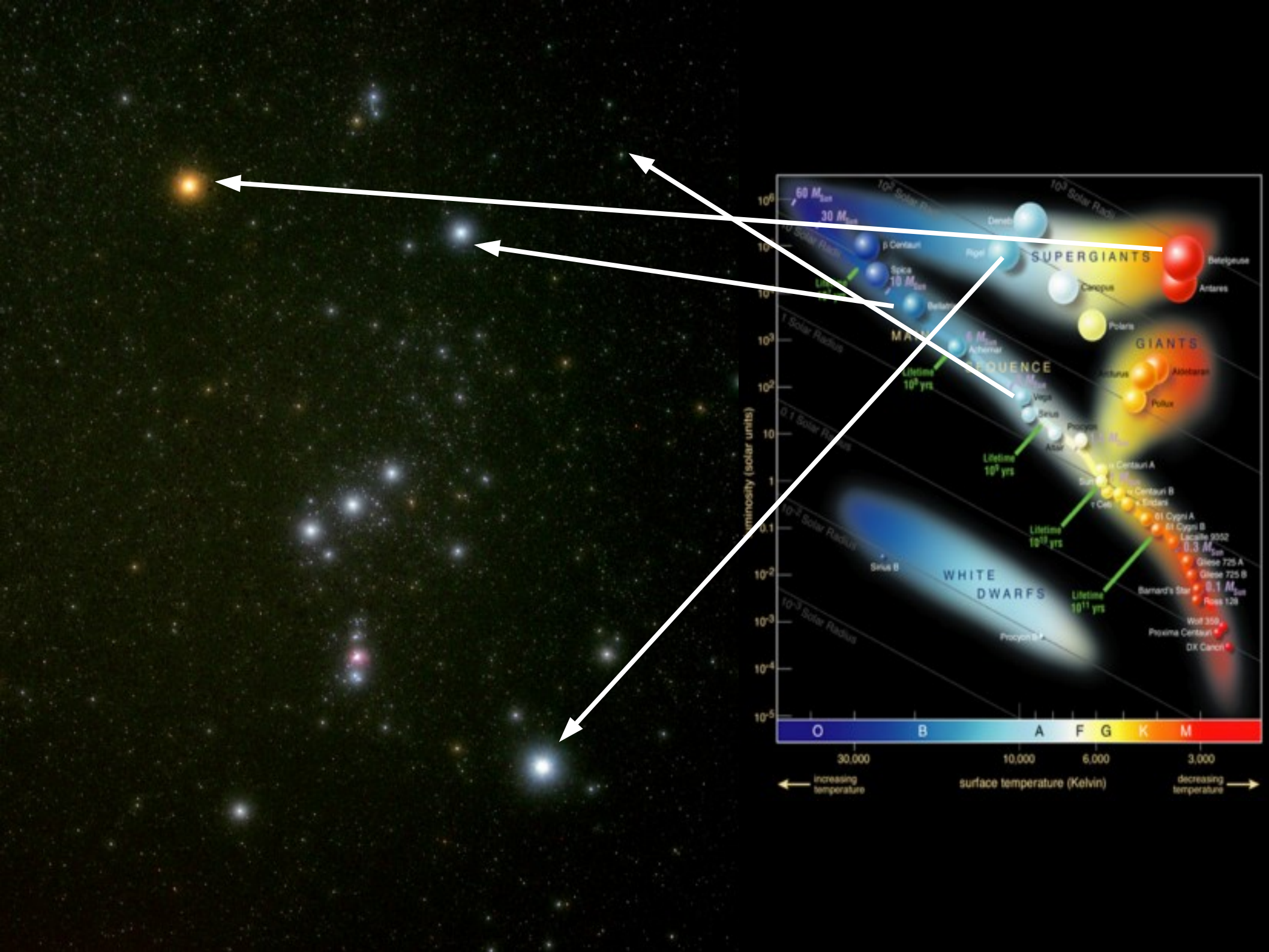


Brightest stars are 1 million times as bright as the Sun

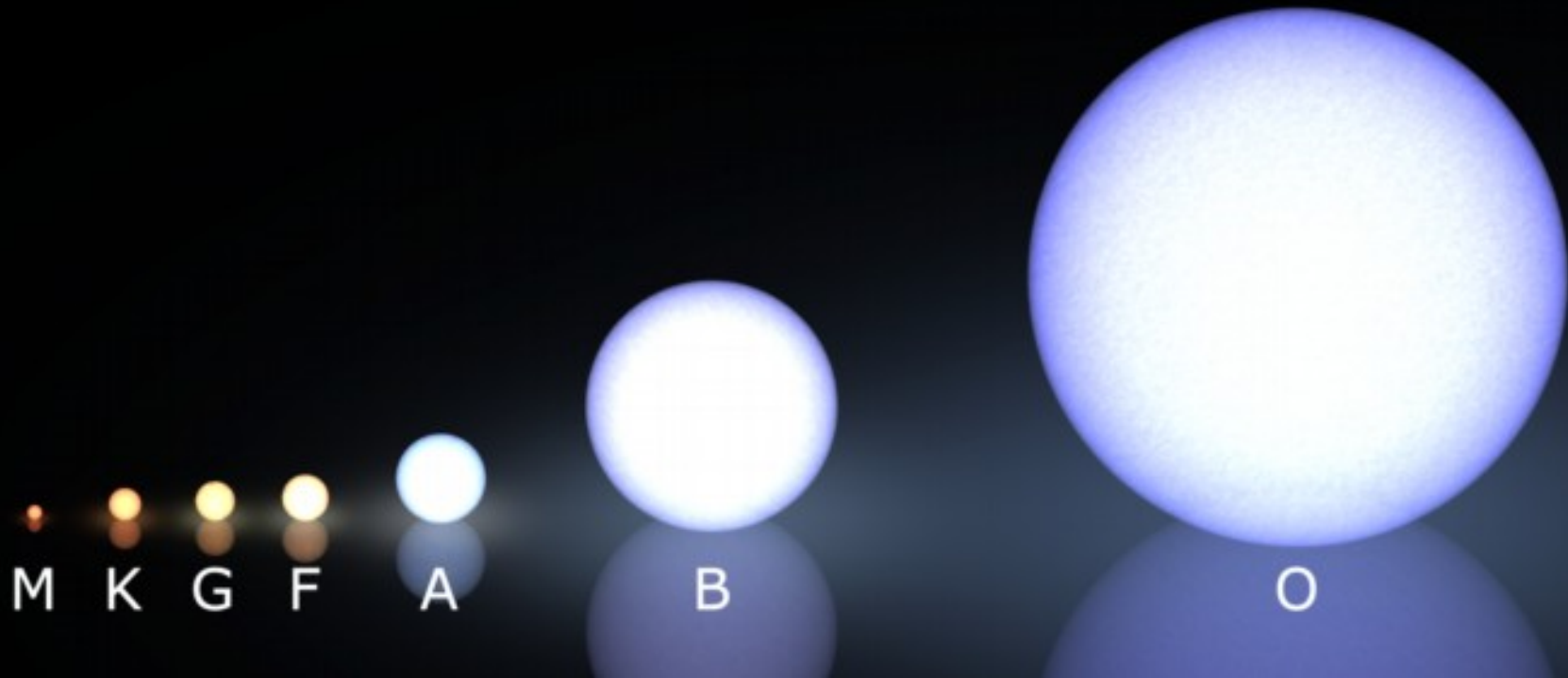


Low mass stars are 0.1% as bright as the Sun, but live for much longer

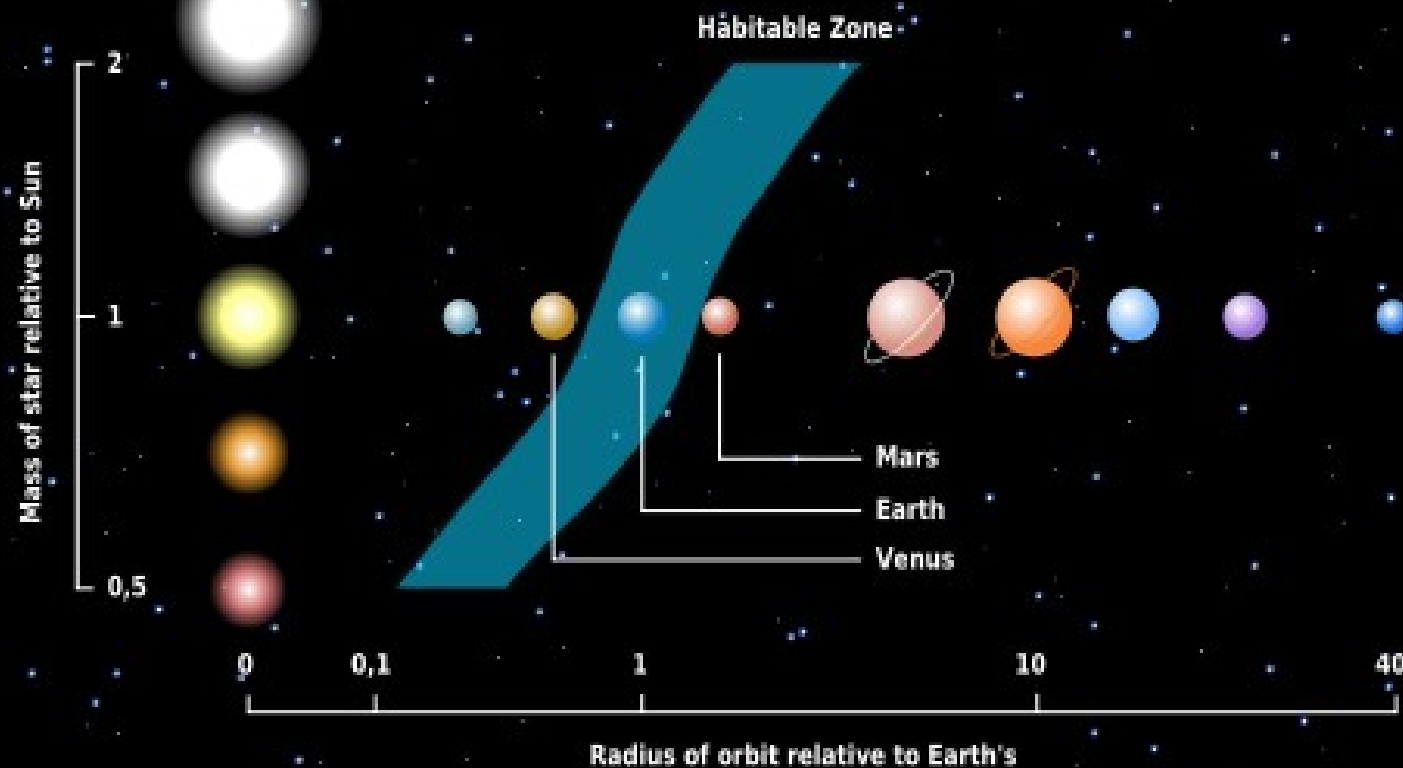




Main sequence : stellar types



Habitable zone of a star



Every star has a habitable zone

Life around... Theta1 Orionis C

Luminosity = 250,000x Sun

Distance = 1,500 light years

Age ~ 1.2 million yr

Mass = 40x Sun

Radius = 8x Sun

Surface temperature = 45,000 K
(Sun = 5,780 K)

Habitable zone =

Star apparent size from habitable zone =

Time in main sequence =



Life around...

Theta1 Orionis C

Luminosity = 250,000x Sun

Distance = 1,500 light years

Age ~ 1.2 million yr

Mass = 40x Sun

Radius = 8x Sun

Surface temperature = 45,000 K
(Sun = 5,780 K)

Habitable zone = 500 AU

Star apparent size from habitable zone = 0.5 arcmin (60x smaller than Sun)

Time in main sequence = few million yrs



Life around...

Proxima Centauri

Luminosity = $0.0017 \times \text{Sun}$

Distance = 4.2 light years

Age \sim 4.85 billion yr

Mass = $0.123 \times \text{Sun}$

Radius = $0.141 \times \text{Sun}$

Surface temperature = 3,042 K
(Sun = 5,780 K)

Habitable zone =
Star apparent size from habitable zone =
Time in main sequence
Orbital period



Life around...

Proxima Centauri

Luminosity = 0.0017x Sun

Distance = 4.2 light years

Age ~ 4.85 billion yr

Mass = 0.123x Sun

Radius = 0.141x Sun

Surface temperature = 3,042 K
(Sun = 5,780 K)

Habitable zone = 0.04 AU

Star apparent size from habitable zone = 1.5 deg (3x larger than Sun)

Time in main sequence > age of universe

Orbital period ~ 1 week → tidal locking



Other things that matter ...

Ingredients: what is the planet made of ? How much water ?

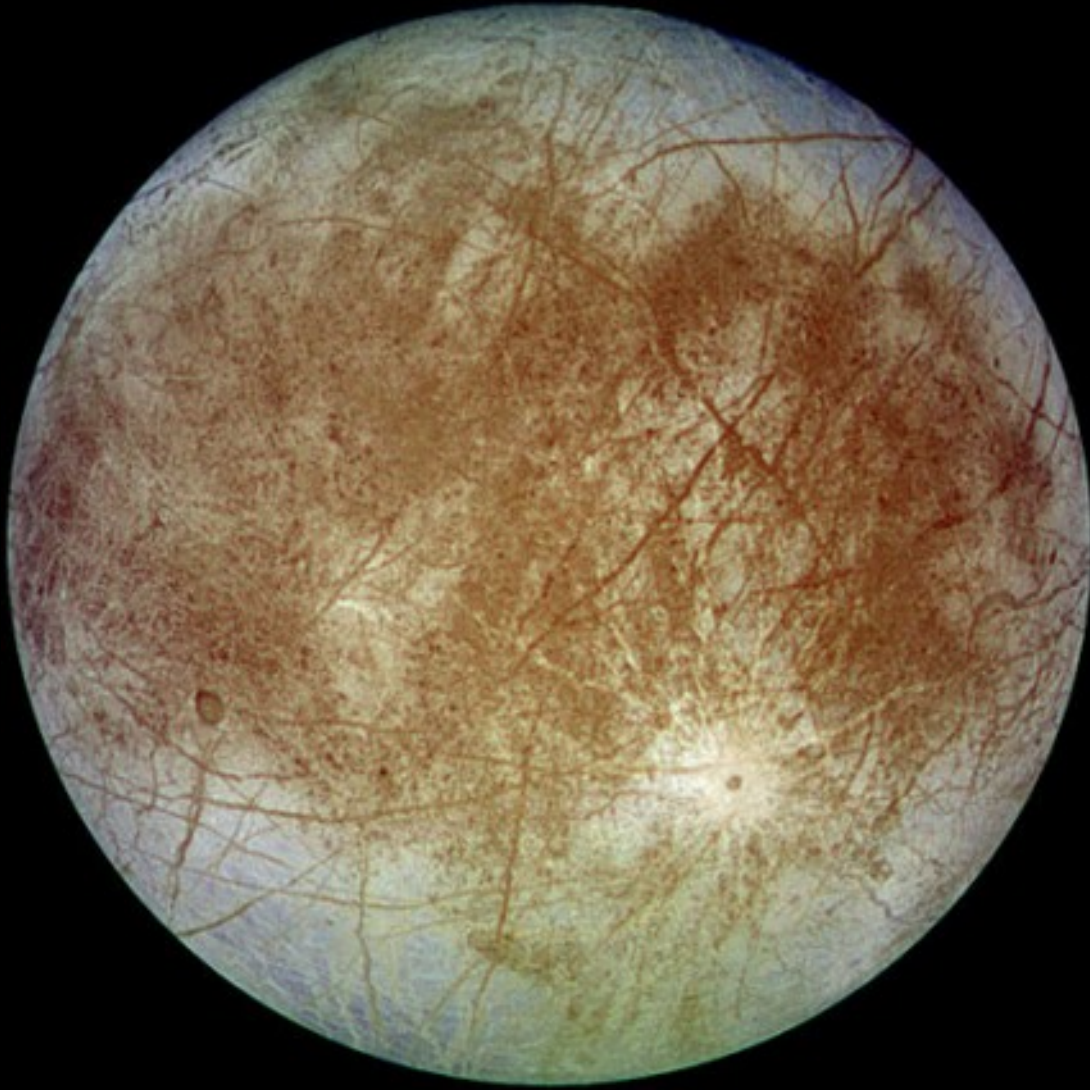
Carbon cycle on Earth regulates temperature

On Earth, carbon cycle enabled by plate tectonics, itself enabled by oceans & continents

Orbit stability (large climate shifts triggered by orbital instabilities)

Large impacts

Other possibilities ?



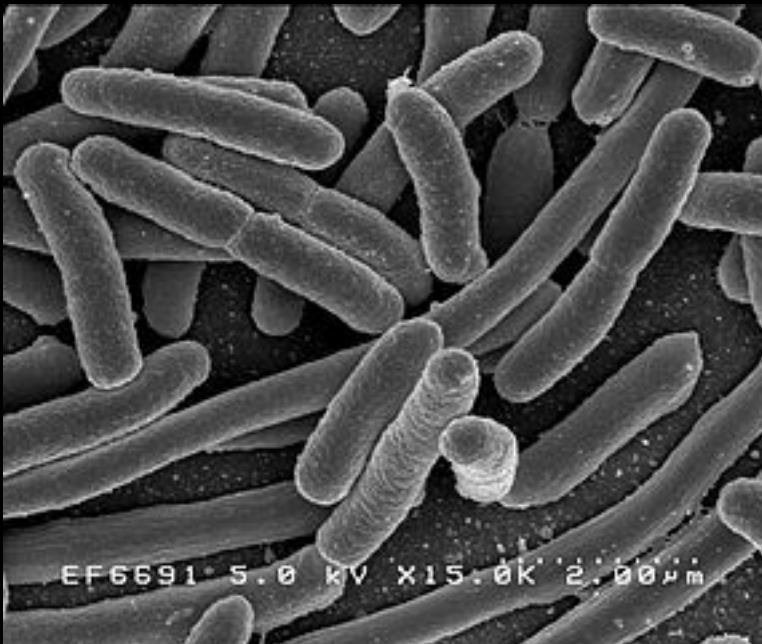
Europa (moon of Jupiter) may have liquid ocean under ice surface

What kind of life ?

“simple” bacteria, or “complex” intelligent life ?

On Earth, life did not evolve regularly

Large steps forward (from bacteria to multicellular organisms 2B yr ago, Cambrian explosion more recently) and also setbacks (snowball Earth episodes, mass extinctions)



Almost endless possibilities for life in the universe...

The universe has MANY galaxies, each composed of MANY stars with planets

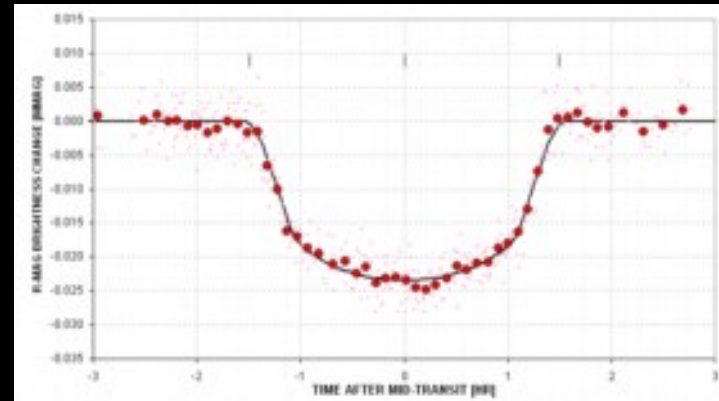
Even if only 1 out of 10 stars has an habitable planet ...

... If you could visit 1 habitable planet per second, it would take 600 years to visit all habitable planets in our galaxy alone

Almost all known planets are discovered indirectly (no image)

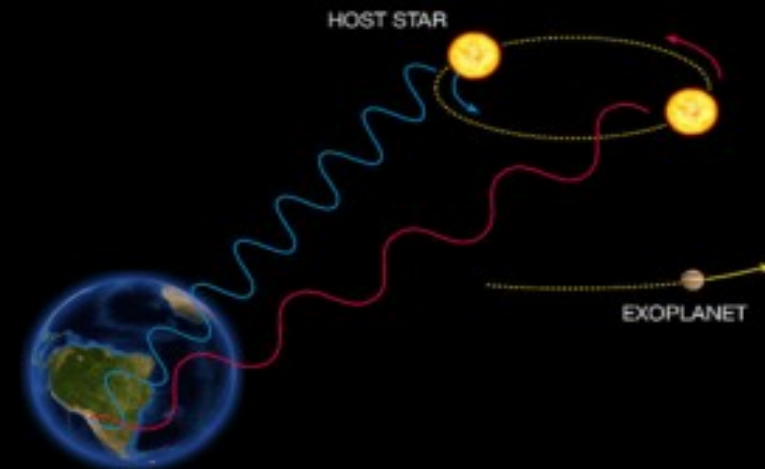
Transit

planet passes in front of star
→ star appears to dim

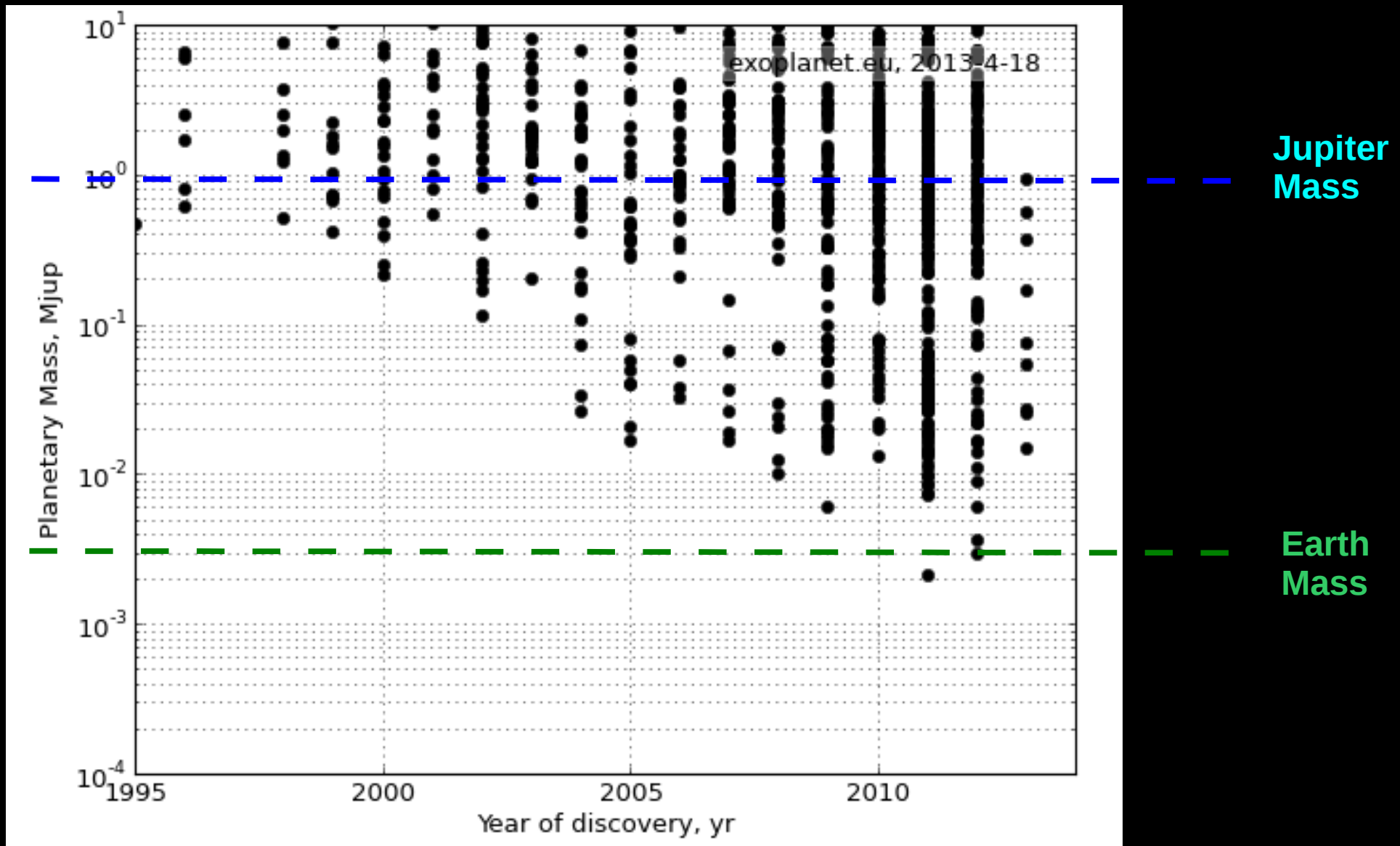


Radial velocity

planet pulls star in small periodic back and forth motion
→ detected thanks to Doppler shift, which makes light bluer as star moves toward us and redder as it moves away



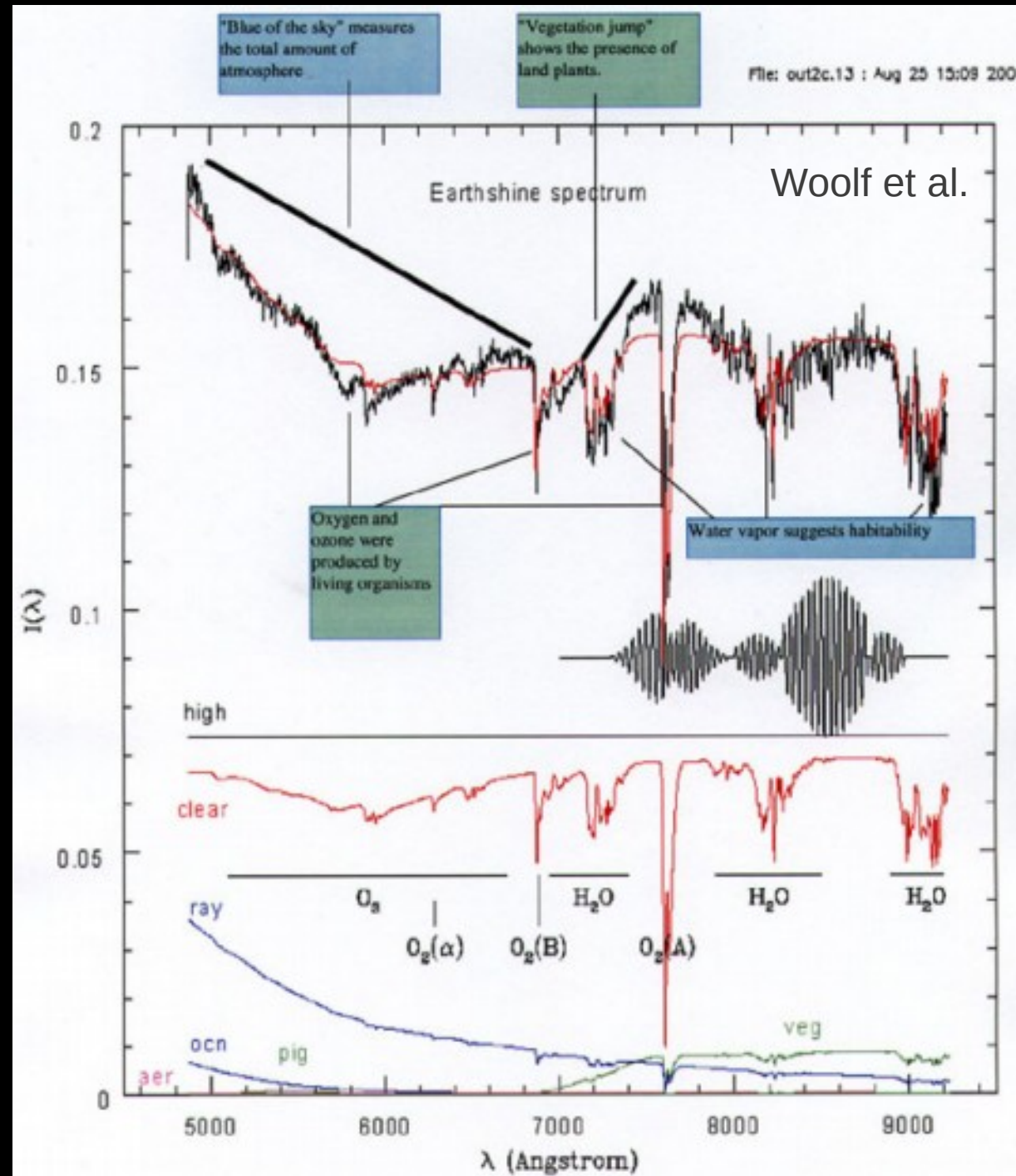
Planets identified – small planets are harder to find



Directly imaging planet is necessary to find life

We need to take spectra of habitable planets

Spectra of Earth (taken by looking at Earthshine) shows evidence for life and plants



Taking images of exoplanets: Why is it hard ?



Earth



