

# Ly $\alpha$ Tomography in *Euclid* Deep Field North with 'Ōnohi'ula PFS

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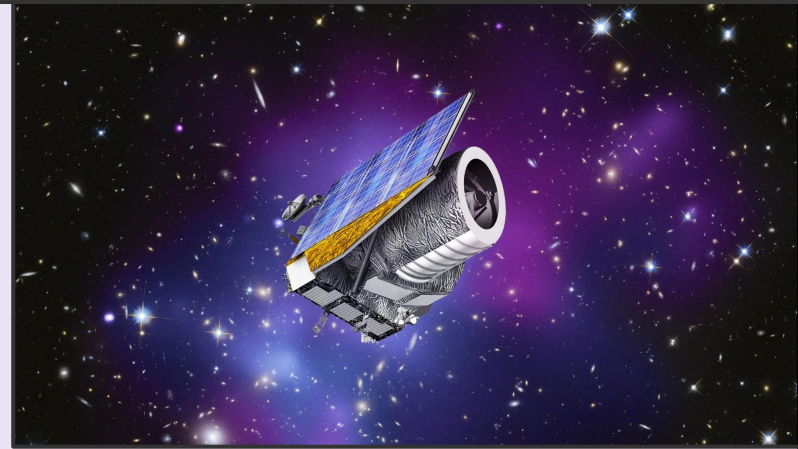
**H20 Team:** Finn Giddings (IfA), Conor McPartland (NBI), John Weaver (MIT), Istvan Szapudi (IfA), Dave Sanders (IfA), Brian Lemaux (Gemini), Sune Toft (NBI), Bahram Mobasher (UCR) et al.



Subaru Users Meeting 2026

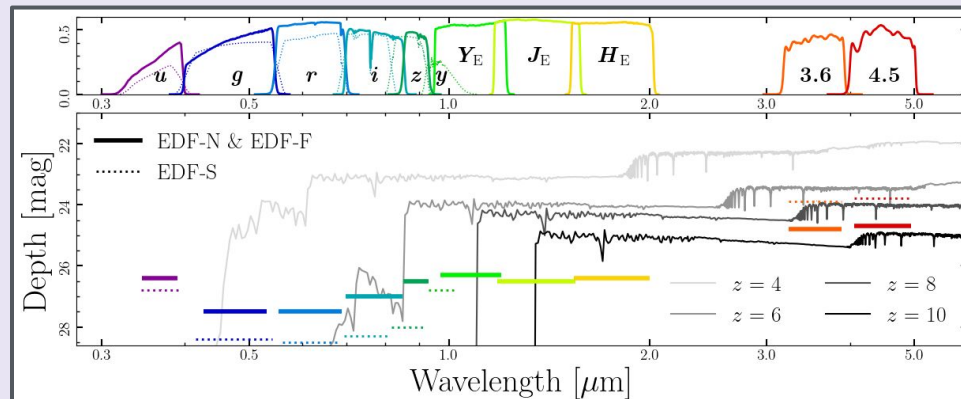
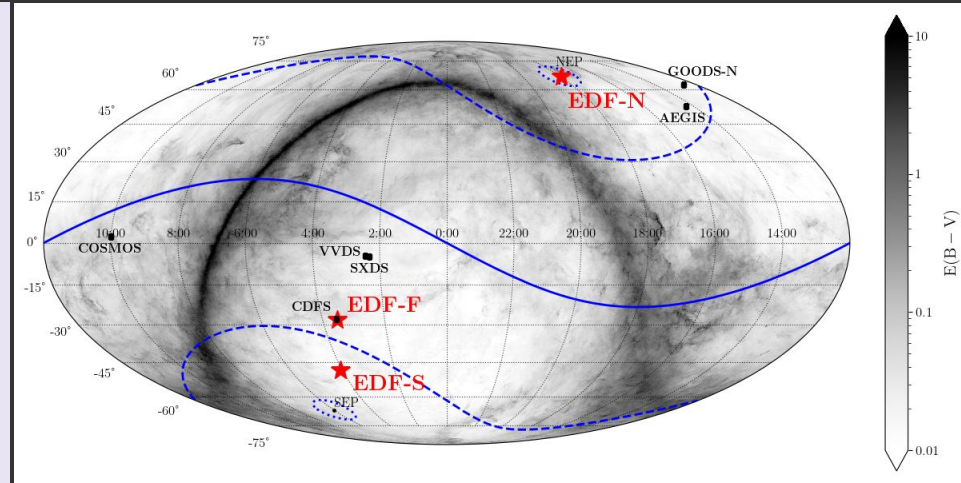
# Euclid Space Telescope

- ★ Launched July 2023!
- ★ VIS/NIR photometry & spectroscopy
  - Wide:  $\sim 14000 \text{ deg}^2$ ,  $m_{AB} < 24.5$
  - Deep:  $53 \text{ deg}^2$ ,  $m_{AB} < 26.5$
- ★ Main goal: constrain dark matter/energy with  $3 \times 2 \text{ pt}$  ( $z < 2$ )
  - Weak lensing, galaxy clustering
- ★ DR1 public release: 11/2026, mid-2027



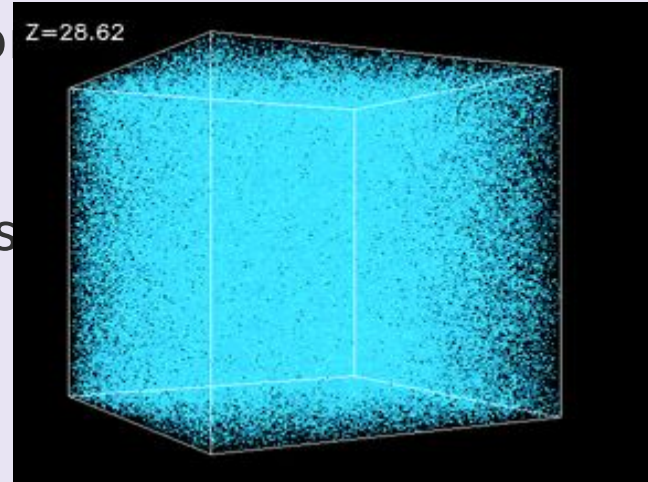
# The Hawaii Two-0 (H2O) Survey

- ★ CFHT *u* + HSC *griz* + IRAC Ch1/2 imaging in EDF-N + EDF-F
  - $m_{AB} < 26.5$  over 30 deg<sup>2</sup>
- ★ PL catalog: ~20 deg<sup>2</sup>, >4 million galaxies, **public** (EC+Zalesky et al., 2025)
- ★ DAWN survey: all EDFs, EAFs
  - Out to  $z \sim 10!$



# Mapping Large Scale Structure

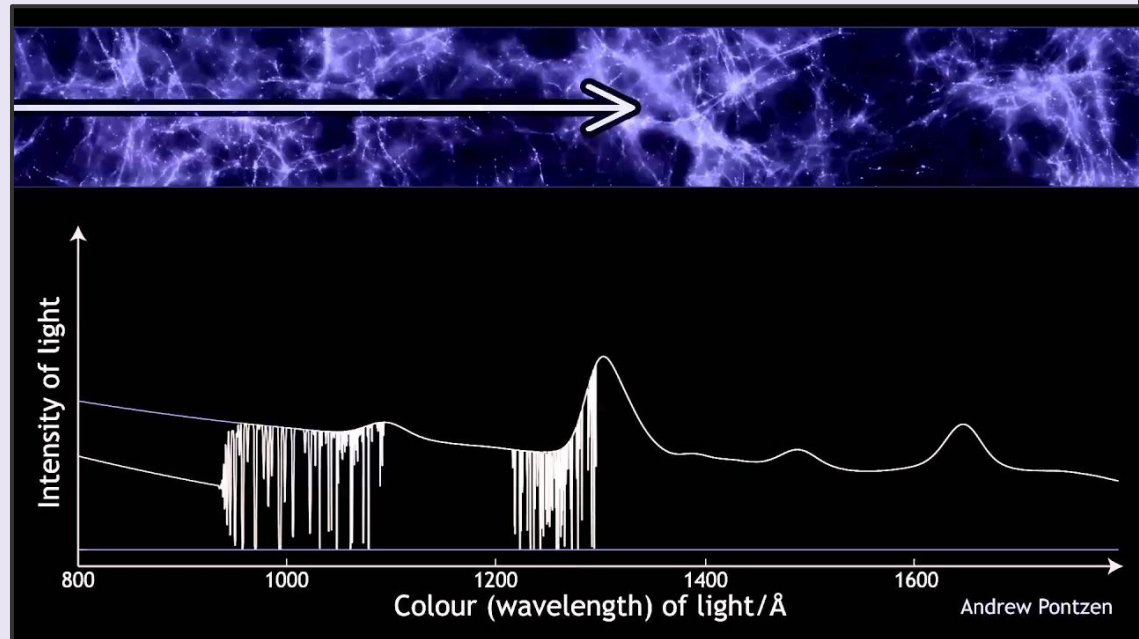
- ★ How do galaxy properties evolve as a function of environment at  $z > 2$ ?
  - Need a large sample of structures (nodes, filaments, sheets, voids)
  - Observed galaxy distribution is a biased tracer of underlying matter distribution
  - On large scales ( $>100$  cMpc), HI traces DM distribution



Simulation by National Center for Supercomputer Applications by Andrey Kravtsov (The University of Chicago) and Anatoly Klypin (New Mexico State University)

# Ly $\alpha$ Tomography

- ★ Use Ly $\alpha$  forest to reconstruct matter distribution via the IGM
  - Lyman continuum absorbed by HI
- ★ Successful with quasars (BOSS, DESI)
  - Limited by spatial sampling (< 100/deg<sup>2</sup>)

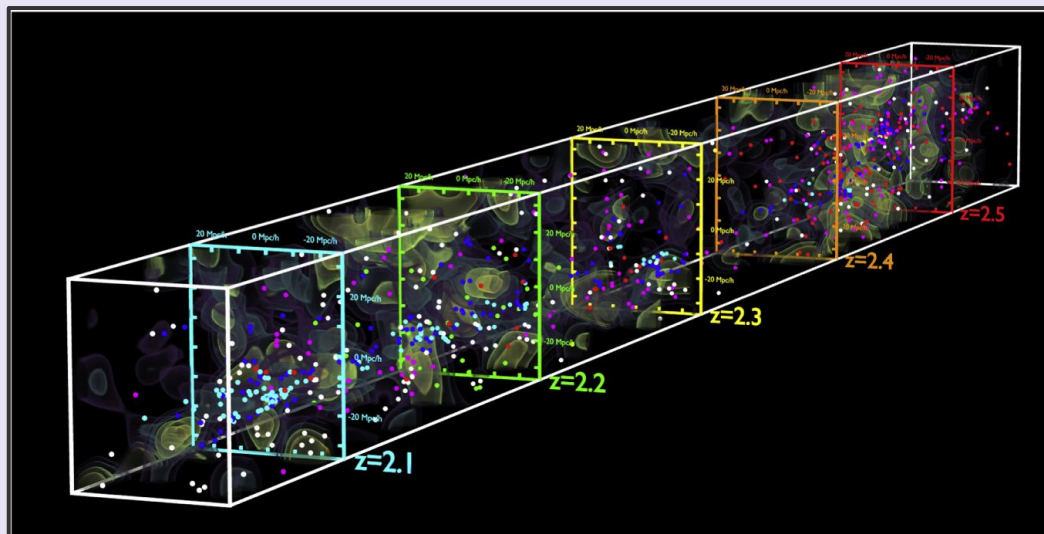


# Ly $\alpha$ Tomography

- ★ Use galaxies instead of quasars to increase spatial sampling
- ★ Main parameter:  $n$ , number of sightlines (galaxies)
  - Sets map

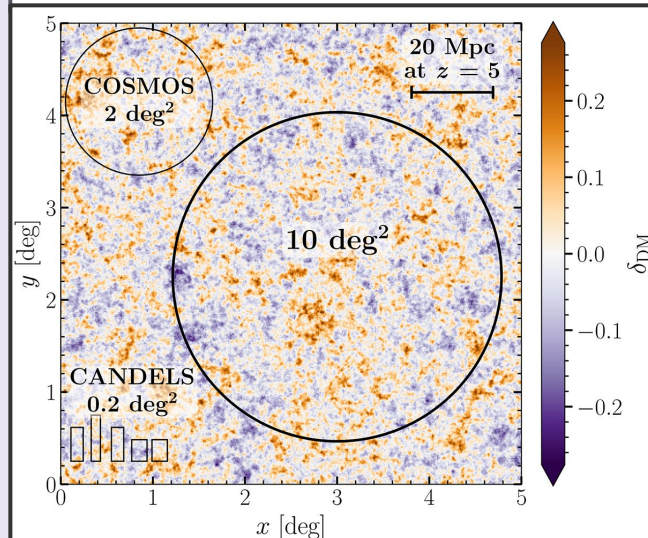
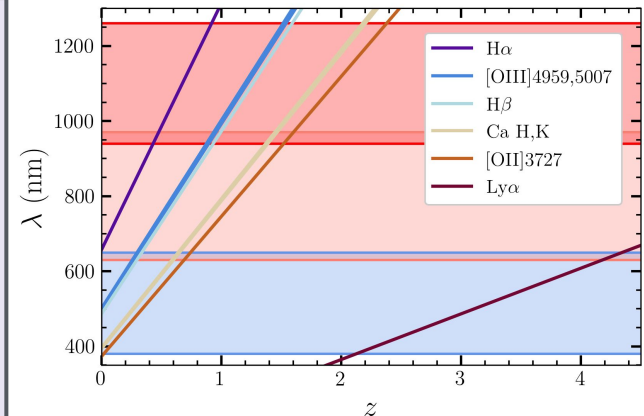
resolution  $\sim n^{-1/2}$

★ CLAMATO	LATIS
$\sim 0.2 \text{ deg}^2$	$\sim 1.7 \text{ deg}^2$
$\sim 2 h^{-1} \text{ cMpc}$	$\sim 3 h^{-1} \text{ cMpc}$
$2.05 < z < 2.55$	$2.2 < z < 2.8$



# Ly $\alpha$ Tomography with PFS

- ★ PFS: great for Ly $\alpha$  tomography!
  - High target density ( $\sim 1400/\text{deg}^2$ )
- ★ H20: great for selecting LBG targets!
  - EDF-N: central  $10 \text{ deg}^2$  complete to full depth (*ugriz* + Ch1/2)
  - Sample diverse environments
    - Rare high/low density peaks



# Ly $\alpha$ Tomography with PFS

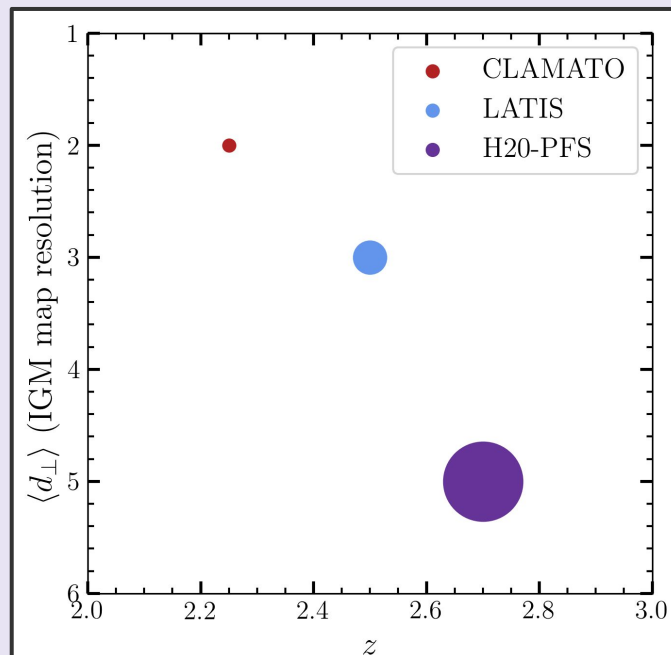
## ★ Target LBGs in EDF-N central 10 deg<sup>2</sup>

- $u$ -band dropouts:
  - $z \sim 2.5 - 3.5, 23 < m_r < 24.5$
- $2.5 < z_{\text{phot}} < 3.5, 23 < m_r < 24.5$
- In H20-PL catalog:
  - $\sim 1,000/\text{deg}^2$
  - $\rightarrow$  IGM map resolution  
 $\sim 5 h^{-1} \text{ cMpc}, 2.2 < z < 3.2$

$$g - r < 1.2$$

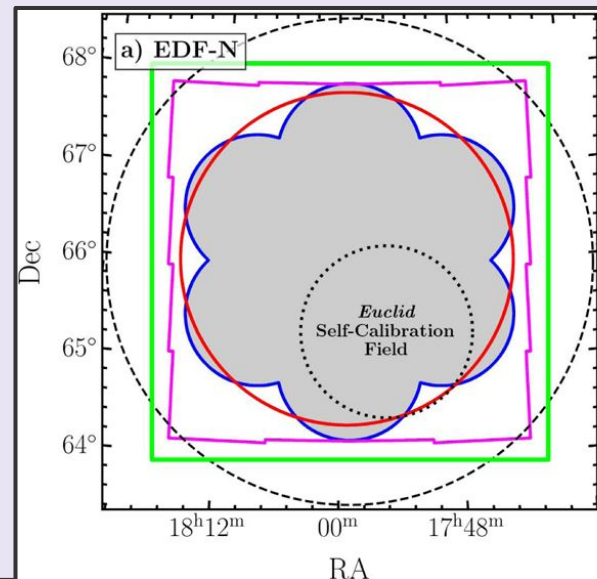
$$u - g < 0.98$$

$$u - g > 1.99 (g - r) + 0.88$$



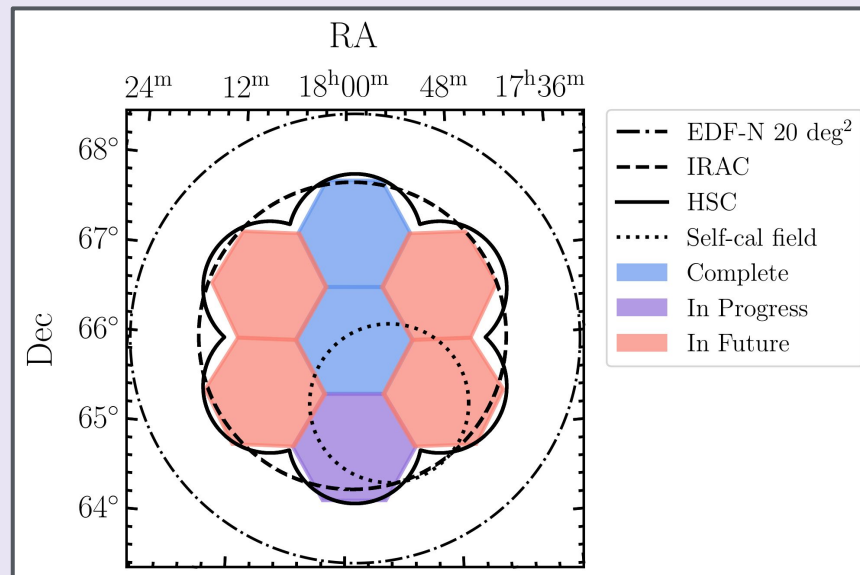
# Observing Plan

- ★ Targets: 7 pointings
  - LBGs ( $\sim 1000/\text{deg}^2$ )
    - Goal: 8hrs exposure time (SNR/pix  $\sim 2$ )
  - Fill extra fibers with photo-z sources
    - $z < 1$ , massive galaxies, strong lenses



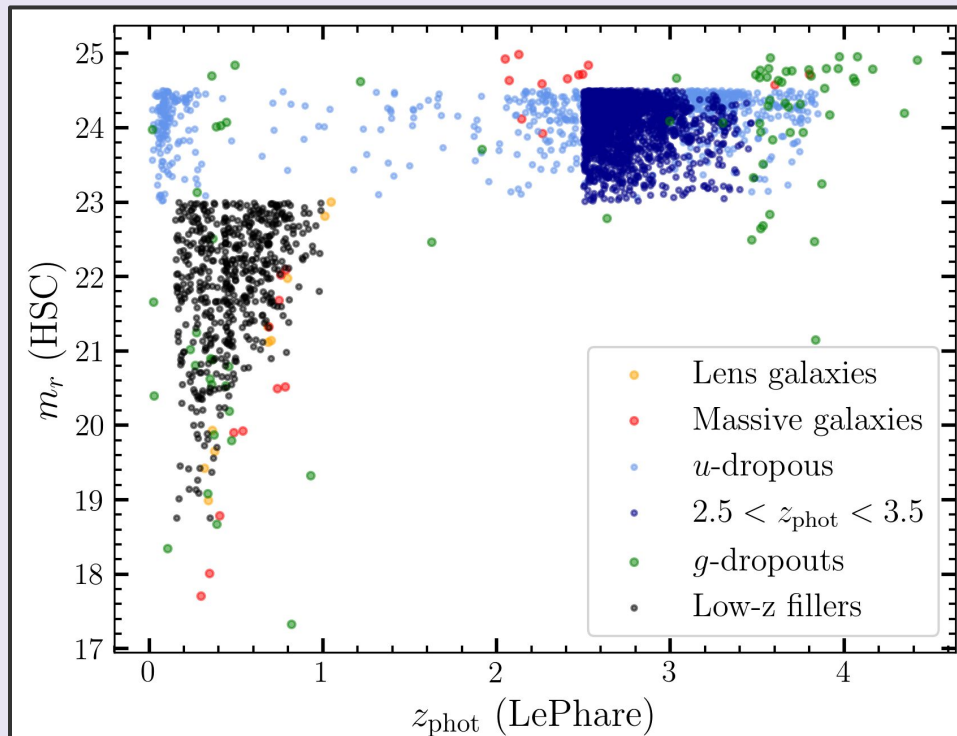
# Observations

- ★ S25A: 2.5N classical (5/23, 6/20-22)
  - 5hr on center, 6.5hr on top pointing (3hr bad weather)
  - Lost ~0.5N to weather
- ★ S26A: 2N classical (5/17-20)
  - 3-3.5hr on 3 pointings
  - Lost 0.5N to weather
- ★ S26B: 1.5N queue
  - August, lower pointing



# S25A Observations

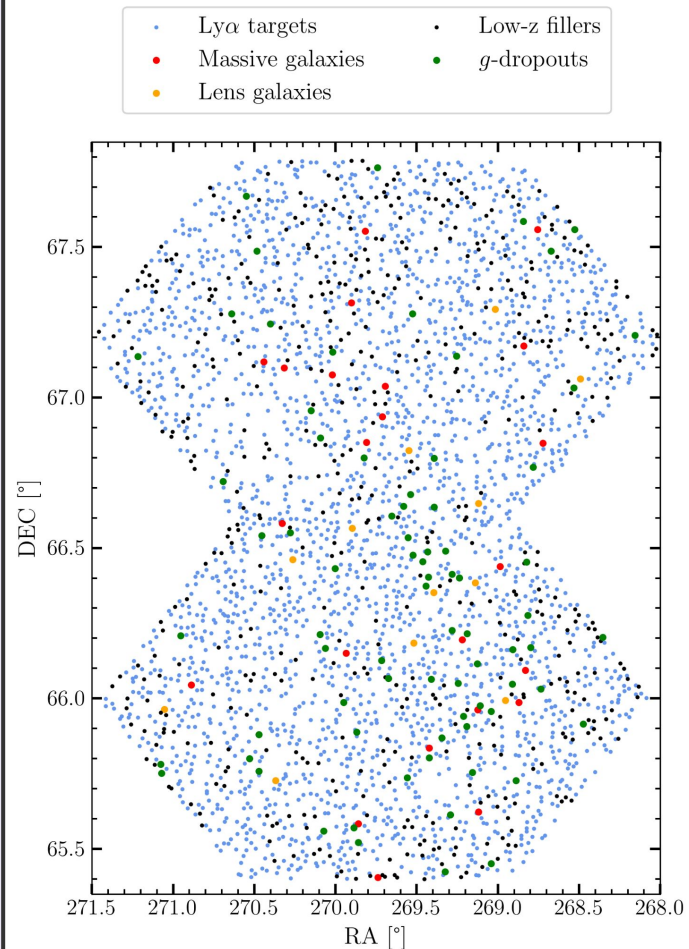
- ★ Targets: S25A\_April2026
  - Strong lenses: 12
  - Massive galaxies: 23
  - $u$ -dropouts: 663
  - $2.5 < z_{\text{phot}} < 3.5$ : 1884
  - $g$ -dropouts: 76
  - $0.1 < z_{\text{phot}} < 1$ : 535



# S25A Observations

★ Targets: S25A\_April2026

- Strong lenses: 12
- Massive galaxies: 23
- $u$ -dropouts: 663
- $2.5 < z_{\text{phot}} < 3.5$ : 1884
- $g$ -dropouts: 76
- $0.1 < z_{\text{phot}} < 1$ : 535

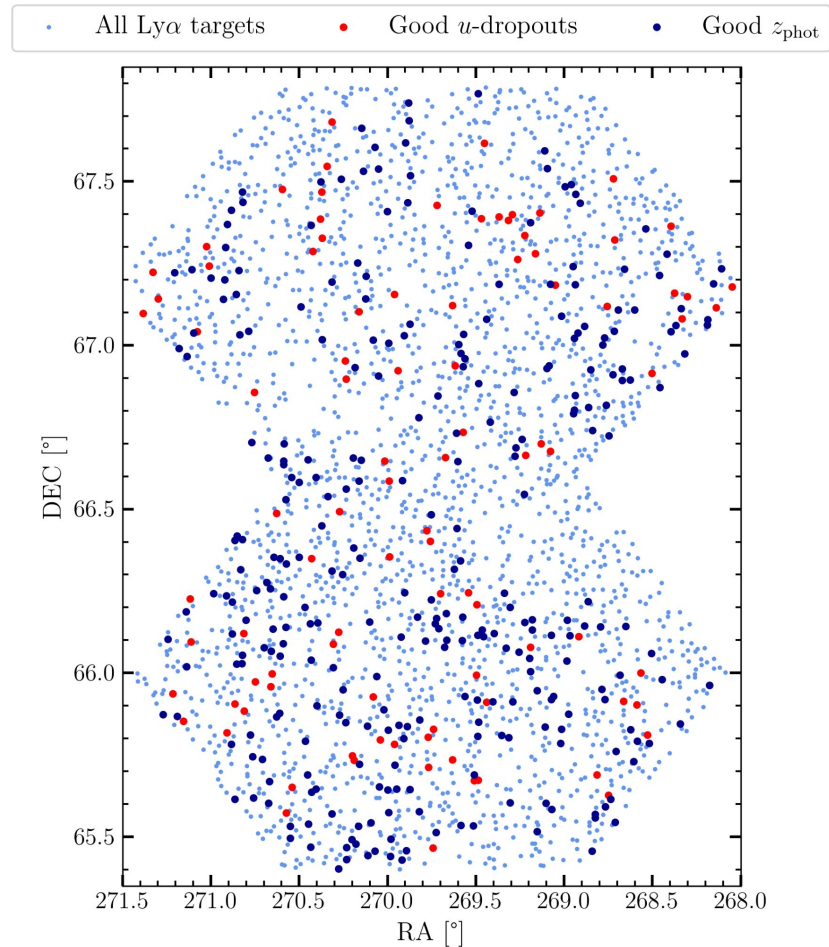


# Preliminary Results

★ Bright LAEs/LBGs confirmed

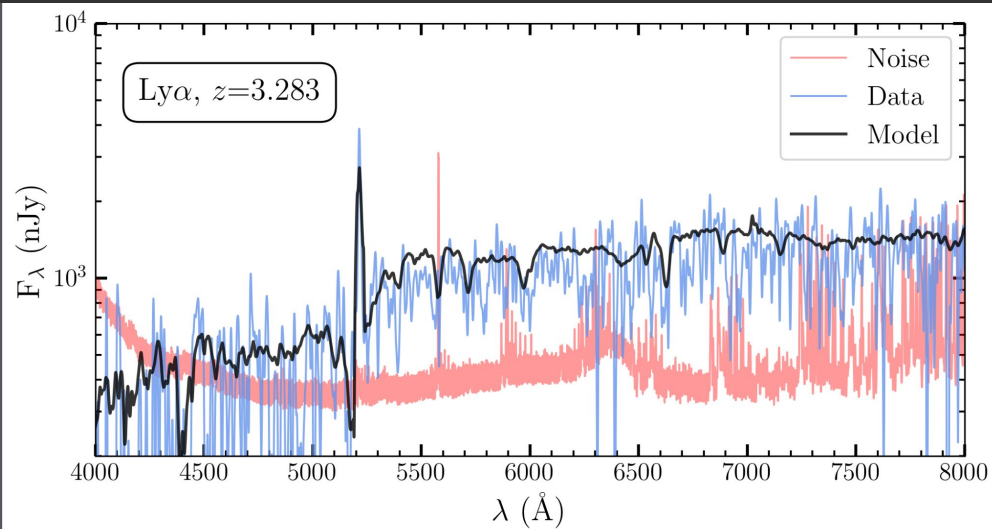
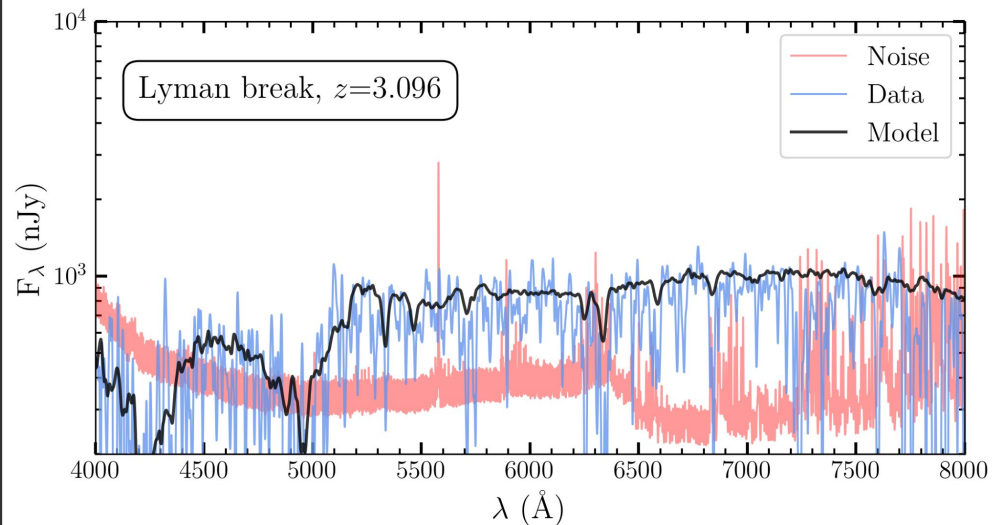
in < 6hrs!

- $r < 24$
- $u$ -dropouts: 95
- $2.5 < z_{\text{phot}} < 3.5$ : 298



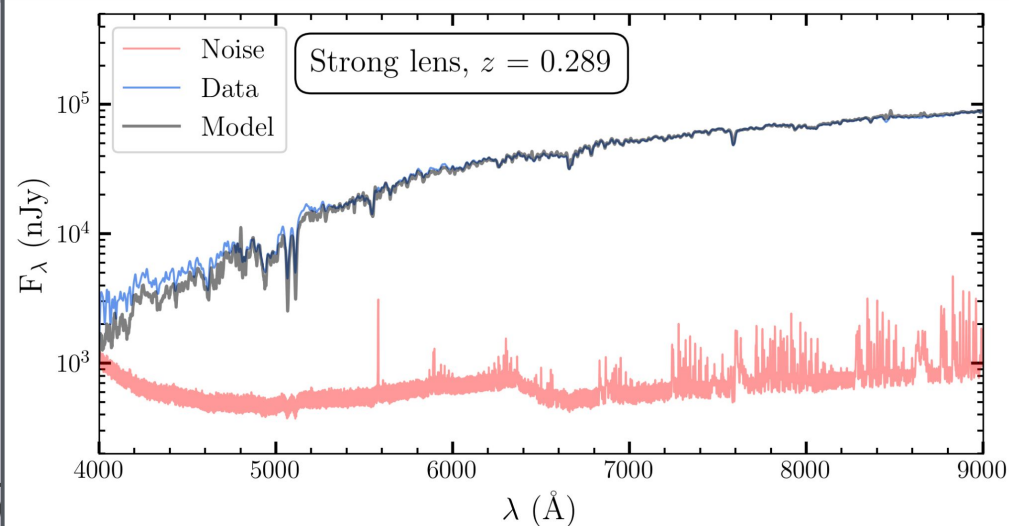
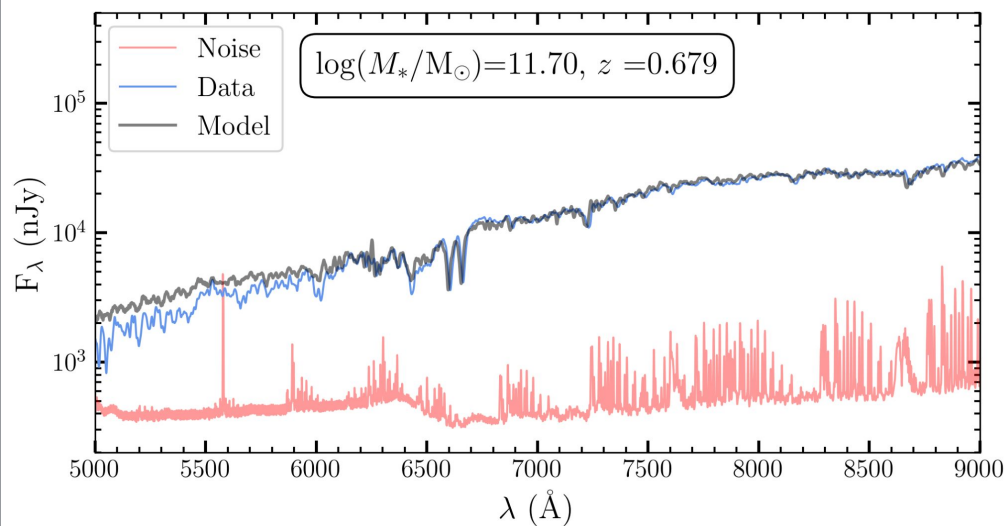
# Preliminary Results

- ★ Example LBG, LAE
  - Only 5hr exposure
  - More data on the way!



# Preliminary Results

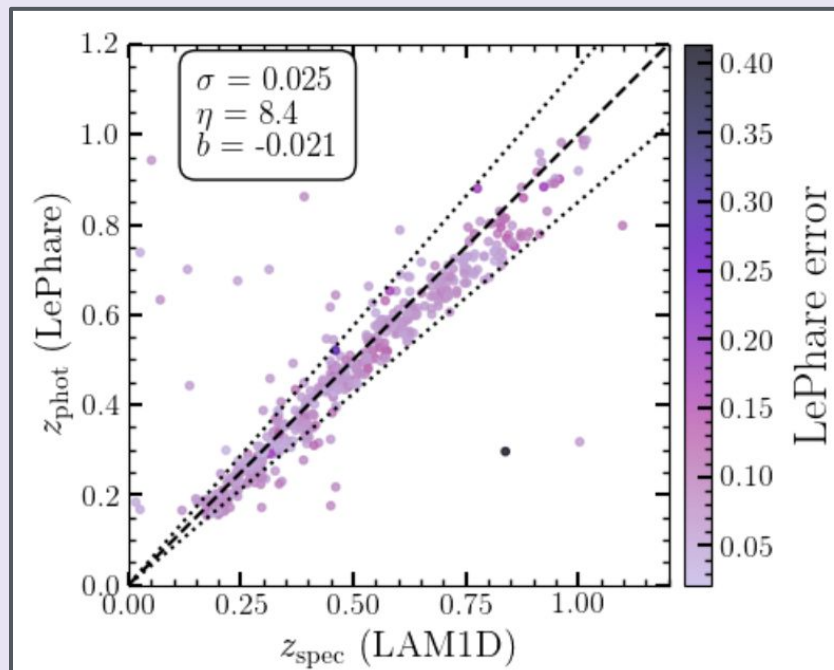
- ★ Additional targets:
  - 11 massive galaxies
    - $\log(M_*/M_\odot) > 11$
  - 12 strong lenses
    - From *Euclid* SWG
  - 535  $z < 1$  galaxies
- ★ All useful for *Euclid* team!



# LAM1D Redshifts

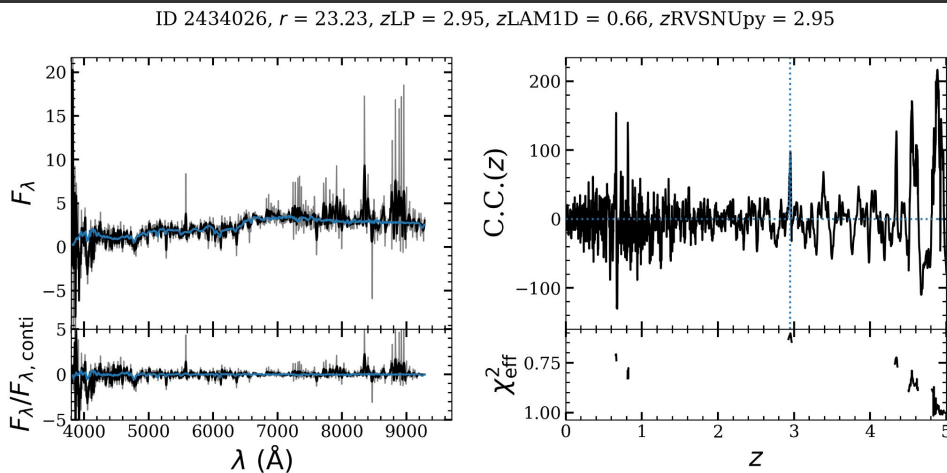
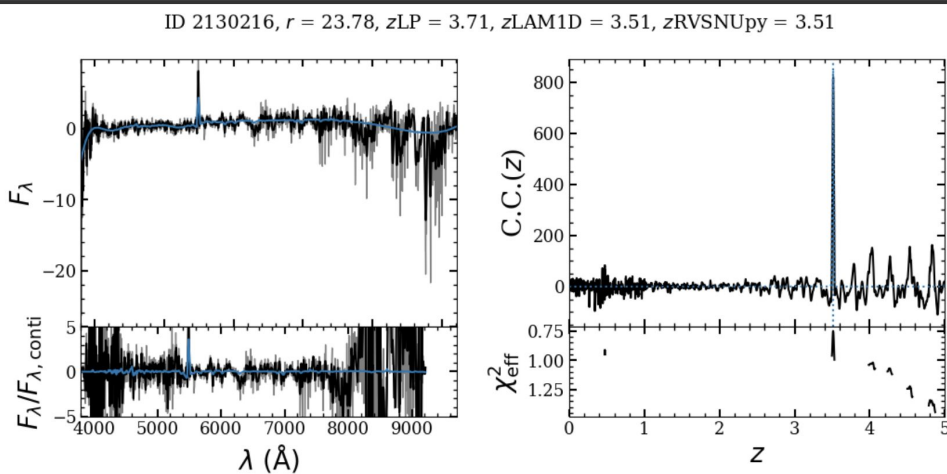
★ Great agreement with LePhare photo- $z < 1$  !

- $\sigma$  similar to COSMOS2020
- Will help calibrate H20/*Euclid* photometric redshifts



# RVSNUpy Redshifts

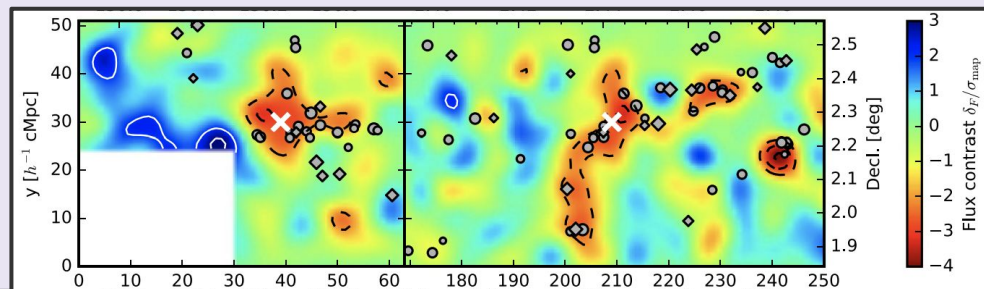
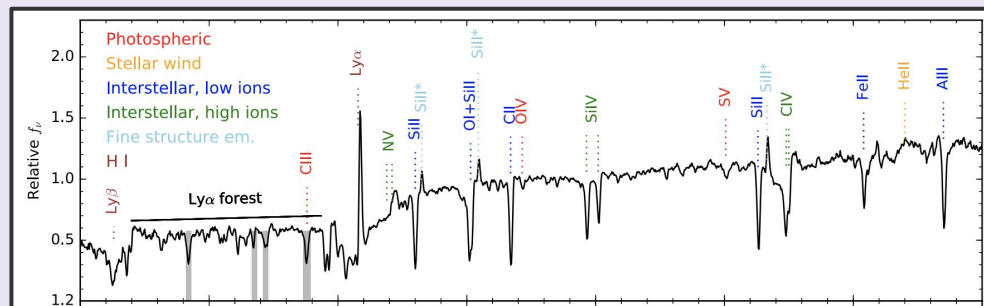
- ★ RVSNUpy: template cross-correlation (Kim+2025)
  - Goal: optimize for  $z > 2$  galaxies
  - Templates: DEEP2 (Newman+2013)
  - Fit Ly $\alpha$ , Lyman break
  - Require visual inspection
    - Finn's GUI!



# Tomographic Analysis

★ Once target depth is reached:

- Fit template to LBG spectra to extract Ly $\alpha$  forest flux
- Create tomographic maps
- Identify structures
- Analyze galaxy properties in different environments



# Summary

For more information on our catalogs,  
check out the DAWN survey website!



- ★ Ly $\alpha$  tomography observations have begun
  - Bright LAEs/LBGs confirmed
  - Successful fillers useful for *Euclid* Consortium
  - Tomographic analysis coming soon
- ★ ありがとう to the whole PFS team!
  - SAs/OAs, data reduction pipeline, science platform, instrumentation...
- ★ Happy to collaborate on tomography, fillers!



Observing from IfA Mānoa :)

# Calibrating Photo-z's

- ★ EAZY & LePhare struggle with Lyman break galaxies...
  - Best-fit SEDs are degenerate
  - CFHT *u*-band + *Euclid* YJH help
- ★ Large sample of  $0.1 < z < 3.5$  spectroscopic redshifts in EDF-N essential for H20/*Euclid* photo-z calibration

