

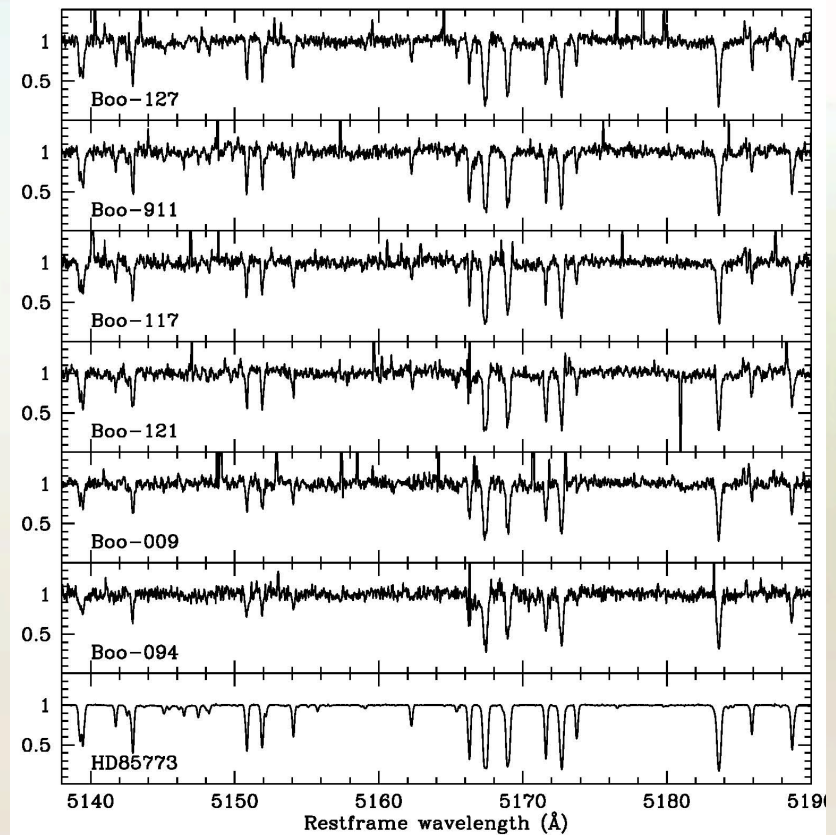
Chemical compositions of a faint Milky Way dwarf satellite galaxy Boötes I

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For more detail:

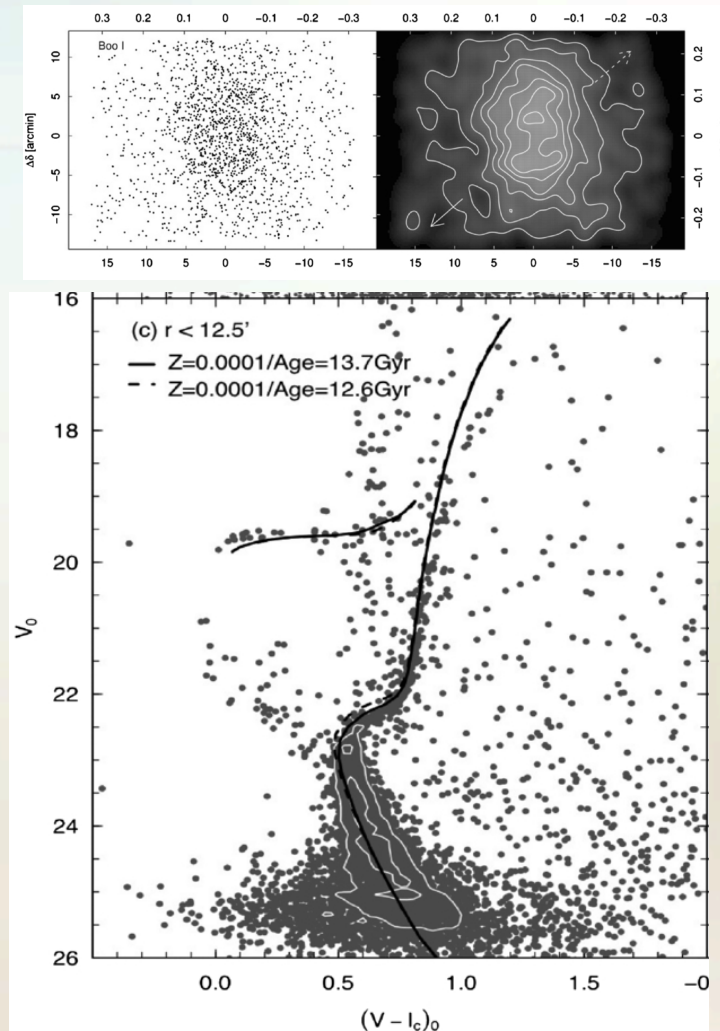
Ishigaki et al. 2014, A&A,
accepted, arXiv:1401.1265



Subaru/HDS spectra of the six Boötes I sample stars

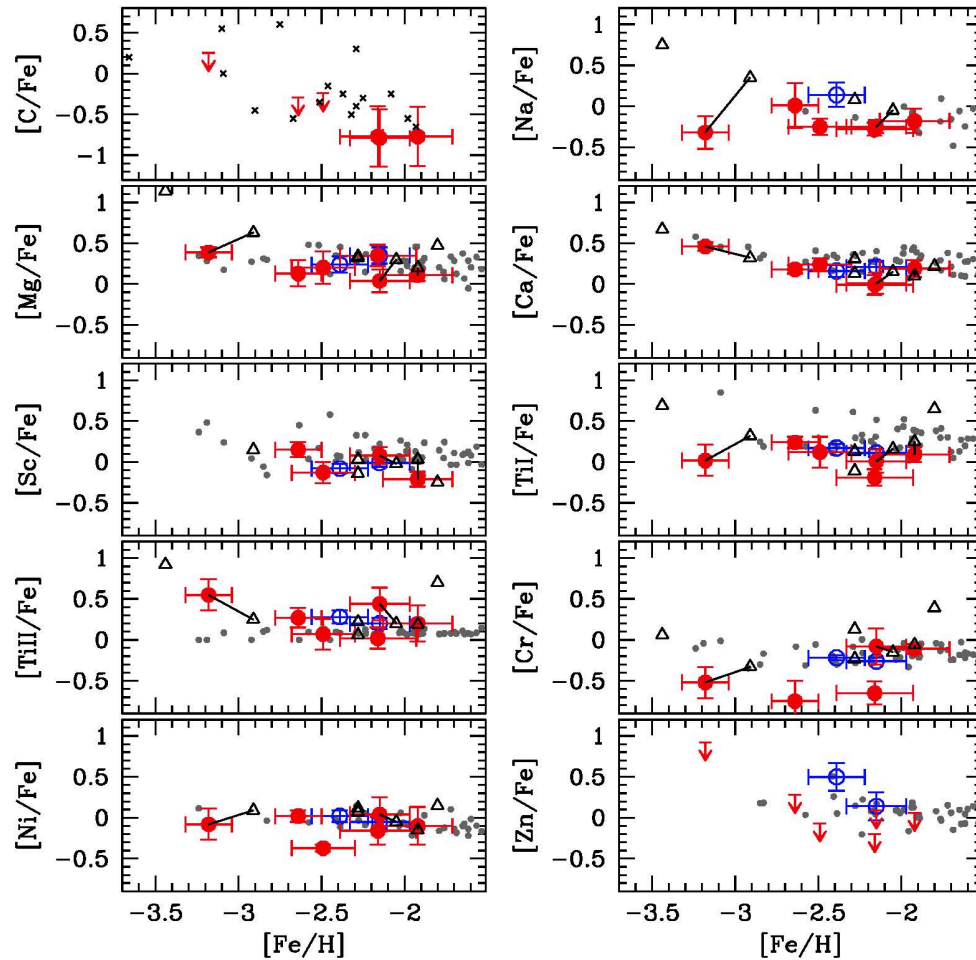
A ultra-faint dwarf galaxy: Boötes I

- Ultra-faint dwarf galaxies in the Milky Way: *fossil record of the galaxy formation*
 - The building blocks of larger galaxies
 - Nucleosynthesis in the first generation of stars
- Boötes I
 - $L \approx 10^4 L_{\odot}$, $M \approx 10^7 M_{\odot}$, $D_{\odot} \approx 66 \text{ kpc}$
 - An exclusively old stellar population
 - Very metal-poor ($[\text{Fe}/\text{H}]_{\text{mean}} \sim -2.5$)
- **High-resolution spectroscopy with Subaru/HDS (PI: S. Okamoto) for six giant stars in Boötes I**

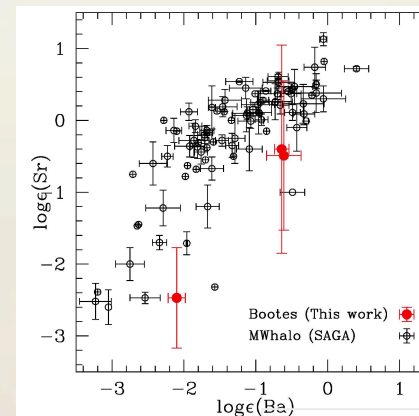


(Top) Spatial distribution of Boötes I stars. (Bottom) Color magnitude diagram of Boötes I (From Okamoto et al. 2012).

Chemical abundances



- The $[X/Fe]$ abundance ratios are mostly homogeneous for the elements lighter than zinc.
- The $[Sr/Ba]$ abundance ratios are lower than bulk of the Milky Way halo stars



Ishigaki et al. 2014