

すばる望遠鏡 装置の現在と近未来 概観

Instruments for Subaru: Overview

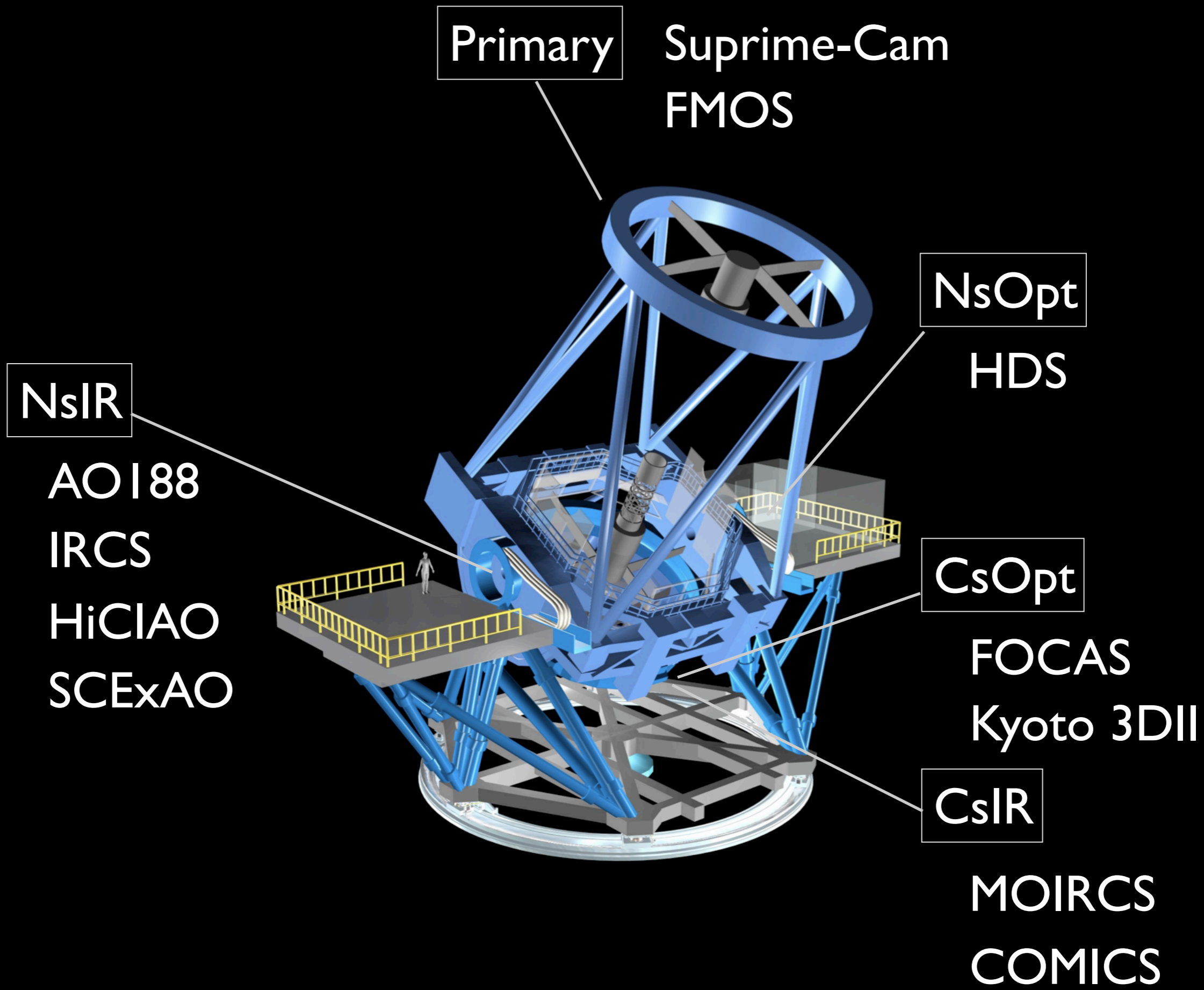
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Ikuru Iwata

Outline

- Current Subaru Instruments
- New and Proposed Instruments
- Subaru Instrumentation Future Plan
 - Next-generation Instruments for Other Telescopes
 - Future Instrumentation Workshop
- Objective of This Workshop

Current Subaru Instruments



Primary

Suprime-Cam
FMOS

NsIR

AOI88
IRCS
HiCIAO
SCEXAO

NsOpt

HDS

CsOpt

FOCAS
Kyoto 3DII

CsIR

MOIRCS
COMICS

Optical Instruments

- **FOCAS**: 撮像、多天体分光、偏光撮像分光
 - FOV: $6'\Phi$, $0.10''/\text{pix}$, $R=250 - 7,500$ ($0.4''$ slit)
- **HDS**: 高分散分光
 - $R=100,000$ ($0.38''$ slit), $0.14''/\text{pix}$
- **Suprime-Cam**: 撮像
 - FOV: $34' \times 27'$, $0.20''/\text{pix}$
- **Kyoto 3D II**: 3次元分光
 - Fabry-Perot: FOV: $1.9' \times 1.9'$, $0.056''/\text{pix}$, IFS: FOV: $3.4'' \times 3.4''$, $0.094''/\text{pix}$
- **Hyper Suprime-Cam**: 撮像
 - FOV: $1.5^\circ\Phi$, 116 CCDs (4 for guiding), $0.2''/\text{pix}$

Infrared Instruments

- **COMICS**: 中間赤外線撮像、分光
 - $\lambda=8 - 25\mu\text{m}$, FOV: 42"x32", 0.13"/pix
- **FMOS**: 近赤外線ファイバー多天体分光
 - 400 fibers, $\lambda=0.9 - 1.8\mu\text{m}$, FOV: 30'φ, R= 500 & 2,200
- **IRCS**: 近赤外線撮像、分光 (w/ AO)
 - $\lambda=0.9-5.5\mu\text{m}$, FOV: 21" (20mas/pix), 54" (52mas/pix), R: 100-2,000 (grism), ~20,000 (echelle)
- **MOIRCS**: 近赤外線撮像、多天体分光 (w/o AO)
 - $\lambda=0.9-2.5\mu\text{m}$, FOV: 4'x7', 0.12"/pix
- **HiCIAO**: 高コントラスト近赤外線撮像
 - $\lambda=0.85-2.5\mu\text{m}$, FOV: 20"x10" (DI, PDI), 5"x5" (SDI), $1e-5.5$ contrast at $r=1''$

New and Proposed Instruments

On-going and Proposed New Instruments

- SCExAO

- IR Doppler
- MIMIZUKU
- SWIMS
- SuFAIFS
- GIGMICS
- RAVEN

- IRCS-HDU
- SIRMOS
- MOIRCS IFU-Unit

- Hyper Suprime-Cam

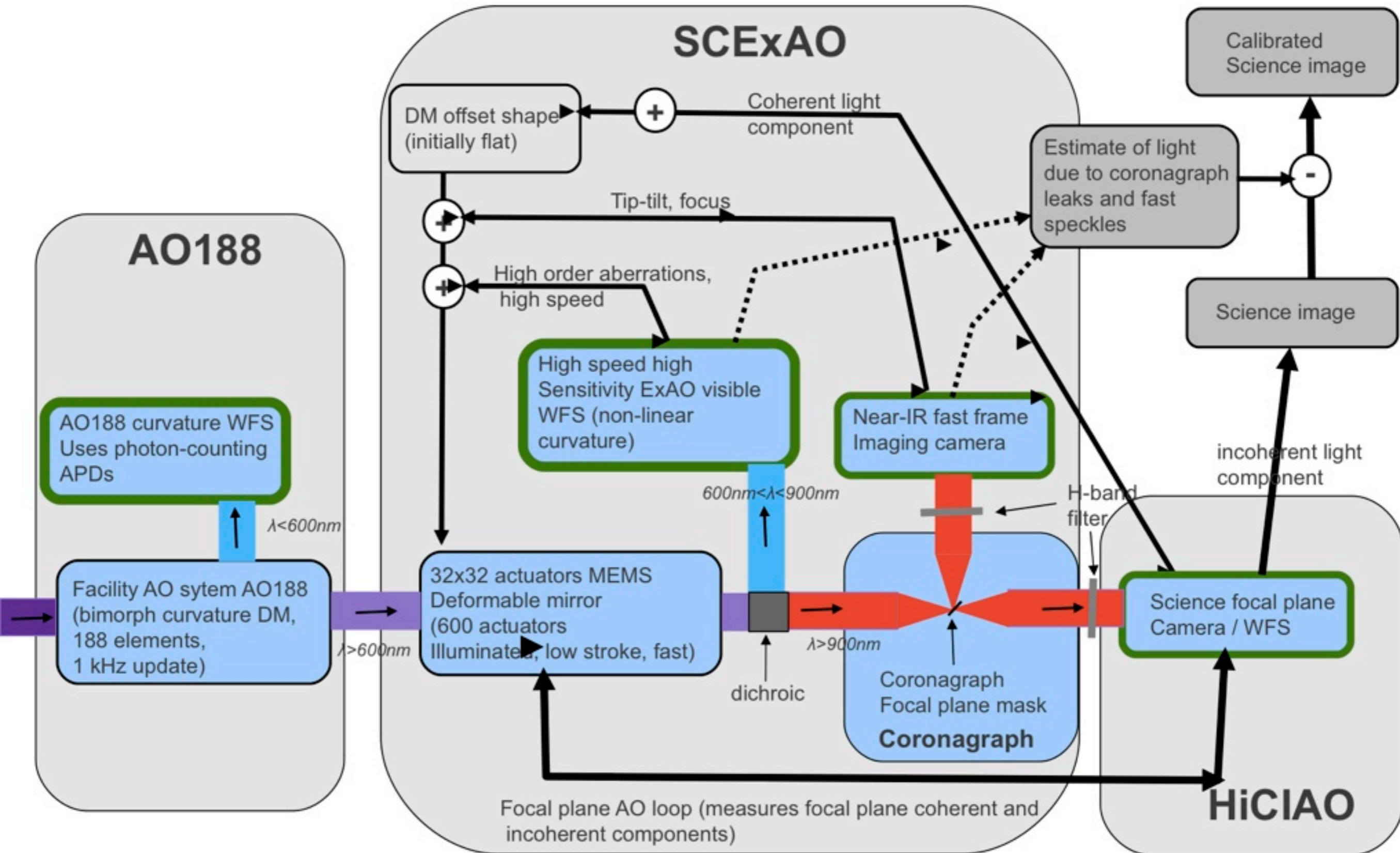
- Kyoto 3D II + AO188
- FOCAS Image Slicer

- Prime Focus Spectrograph (PFS)

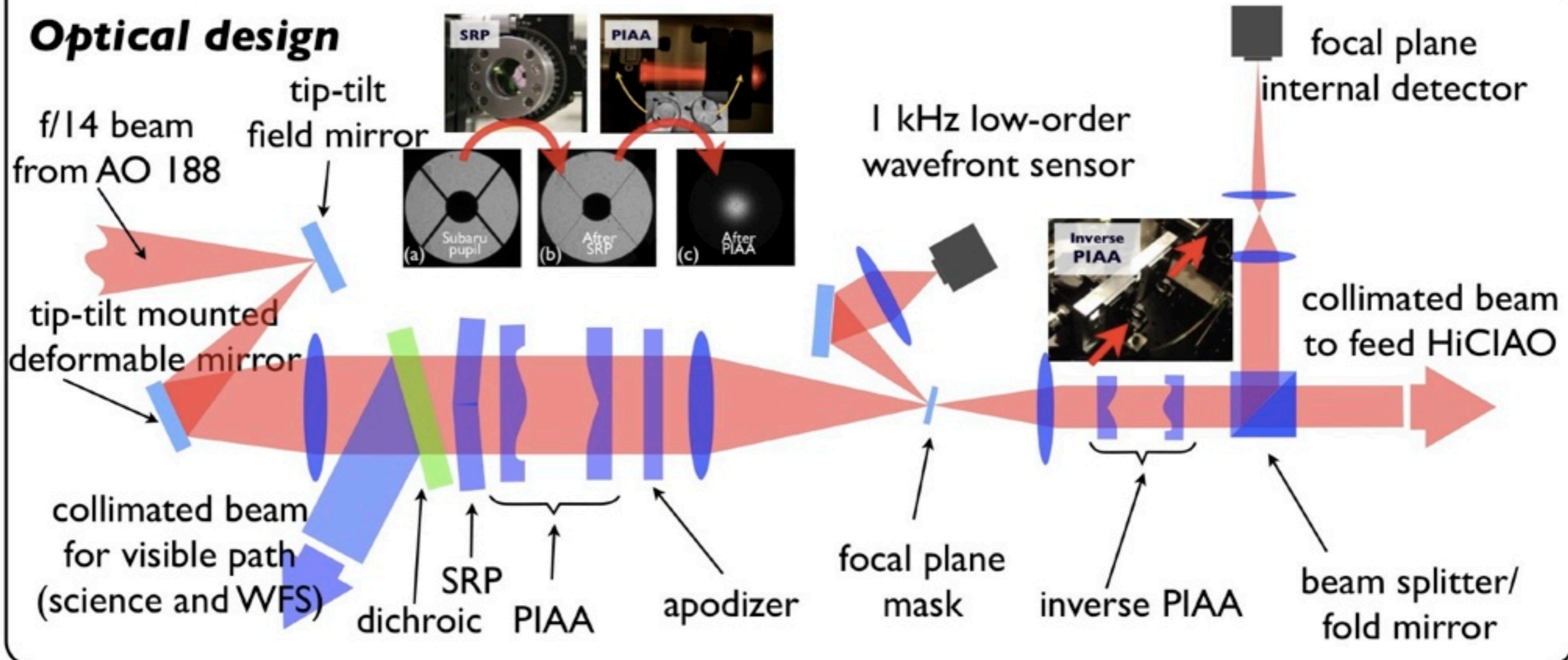
SCExAO: Subaru Coronagraphic Extreme AO

- 1k MEMS, PIAA Coronagraph to Reach $<1e-6$ at $R<0.5''$
- Upgrade to AO188 + HiCIAO
- PI-Type Instrument, PI: Oliver Guyon (Subaru / Univ. of Arizona)
- Day-time Engineering at the Summit (Aug. and Dec., 2010)
- Engineering Acceptance Test in the End of Jan., 2011
- Engineering Observations Scheduled in Feb. and March, 2011

SCExAO Wavefront Control architecture and speckle calibration



SCE_{Ex}AO Optical Design

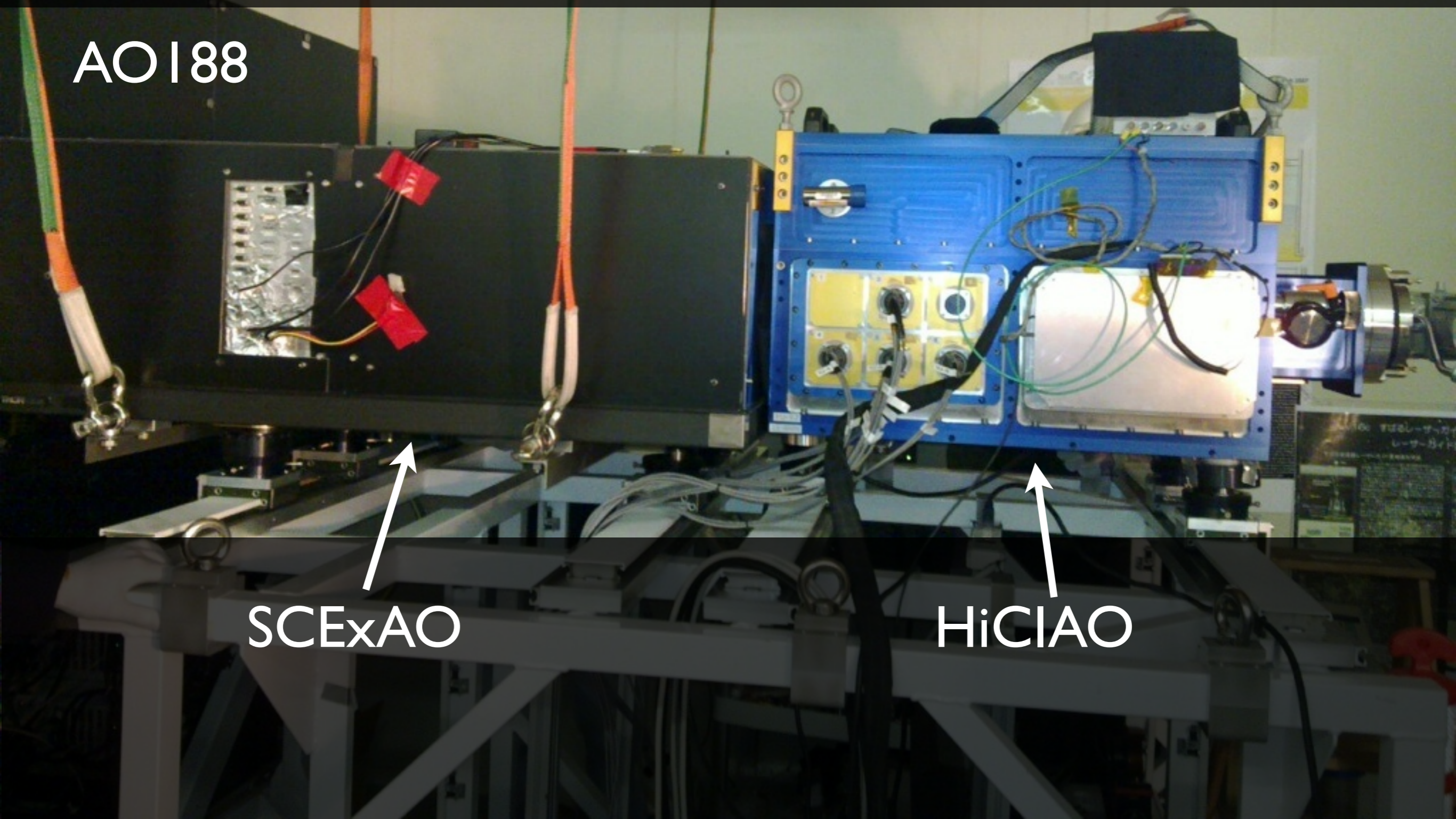


SCExAO Day-Time Engineering, Dec. 2010

AO 188

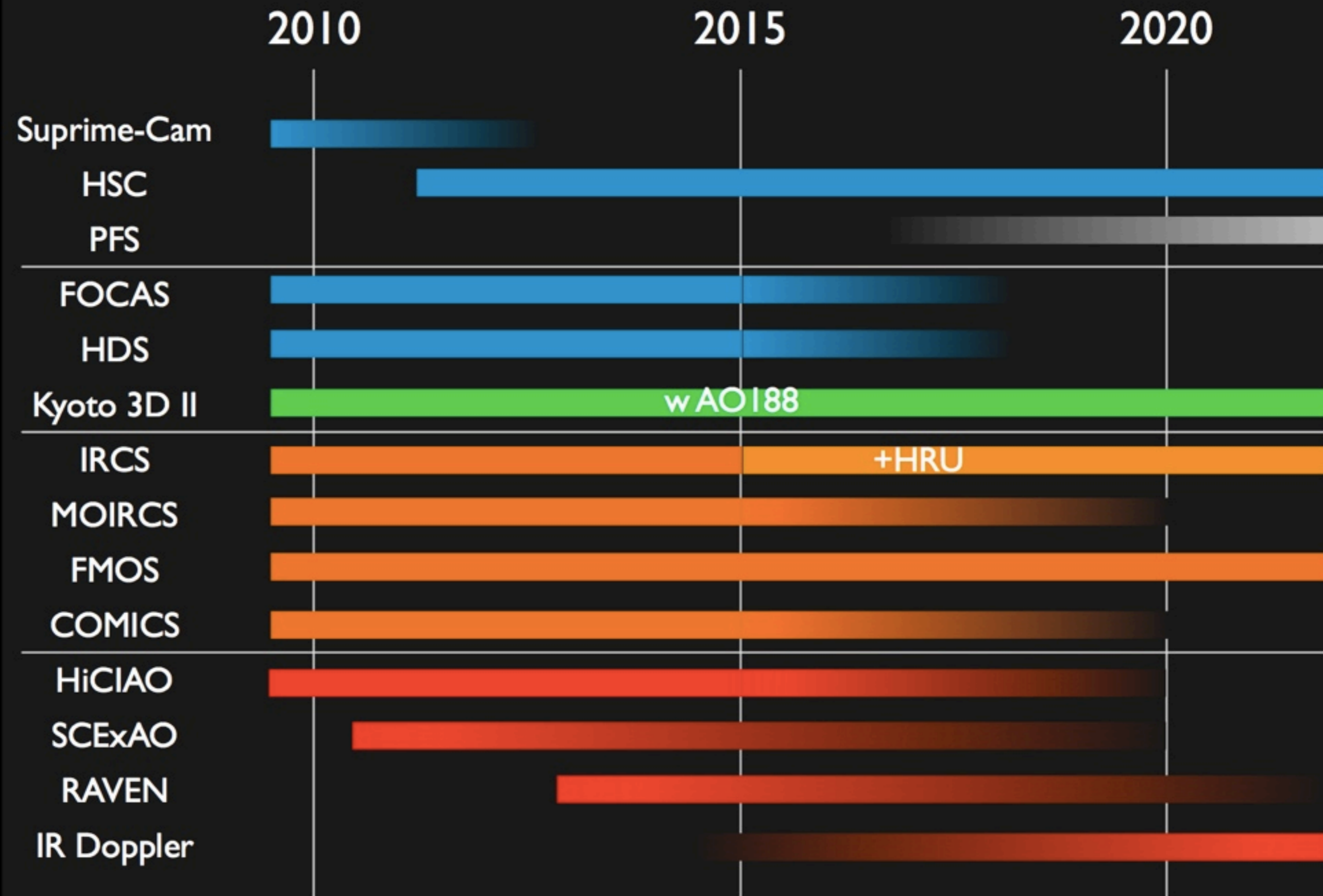
SCExAO

HiCIAO



Subaru Future Instruments

Current Instrument Plan



Strategy for Subaru Next-Gen Instruments

- Future Instrument Candidates (2009年SACレポートなど)
 - 主焦点可視多天体分光器 Wide-Field Multi-Object Spectrograph
 - 広視野近赤外線カメラ Wide-Field Near-IR Camera
 - AO多天体+面分光装置 Multi-Object + Integral-Field Spectrograph assisted by AO
 - 新中間赤外線装置 New Mid-infrared Instrument

What We are Missing

- 可視多天体分光装置

- → Prime Focus Spectrograph (PFS)

- 面分光装置

- 可視域: Kyoto 3D II
 - FOCAS Image Slicer Unit (Ozaki san)
- 近赤外線
 - MOIRCS 面分光ユニット (2011 科研費 基盤S 申請中)

New Generation Large Instruments for Other Telescopes

- MOSFIRE, NIRES
- GPI, FLAMINGOS-2
- KMOS

- MUSE

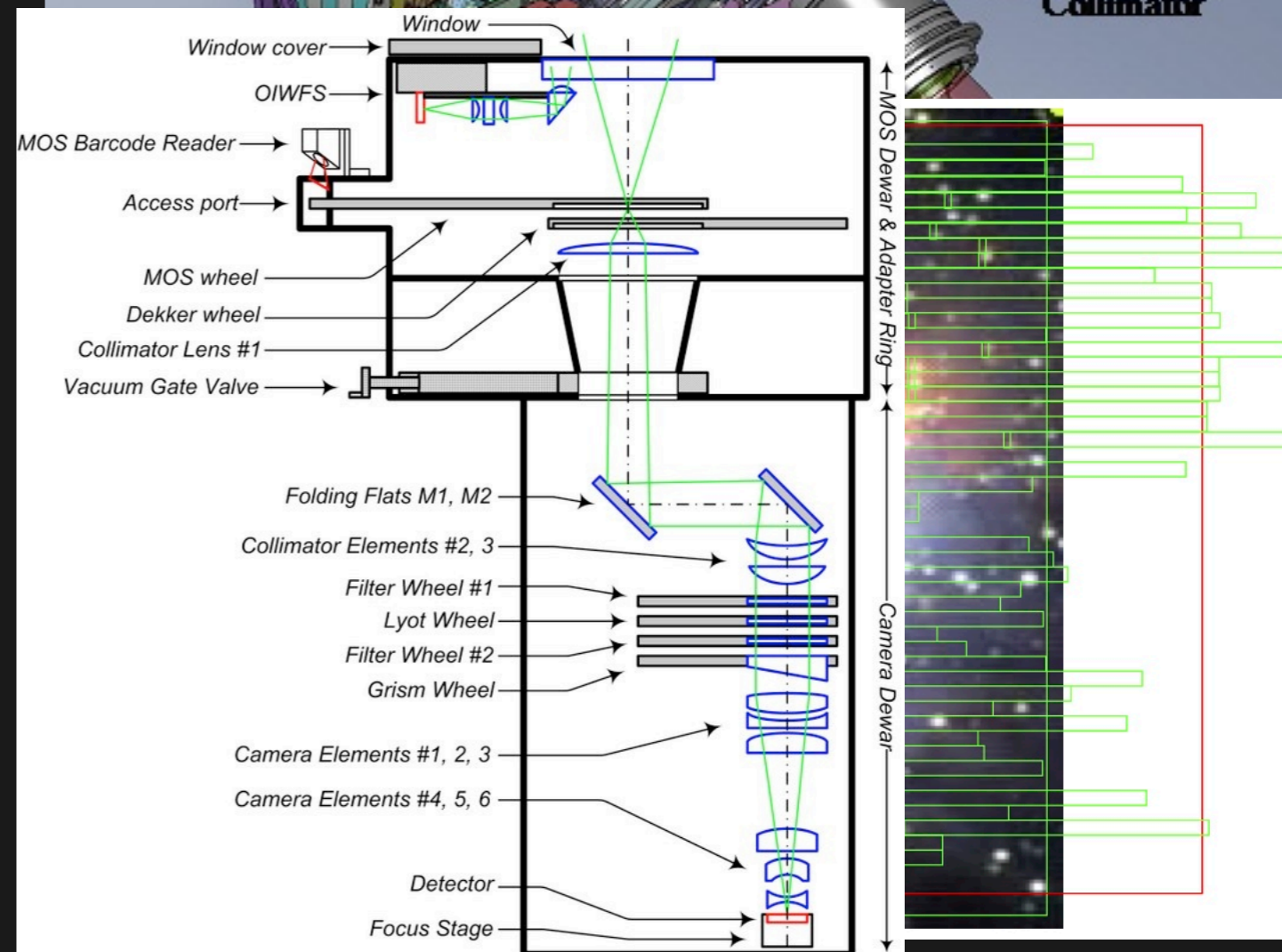
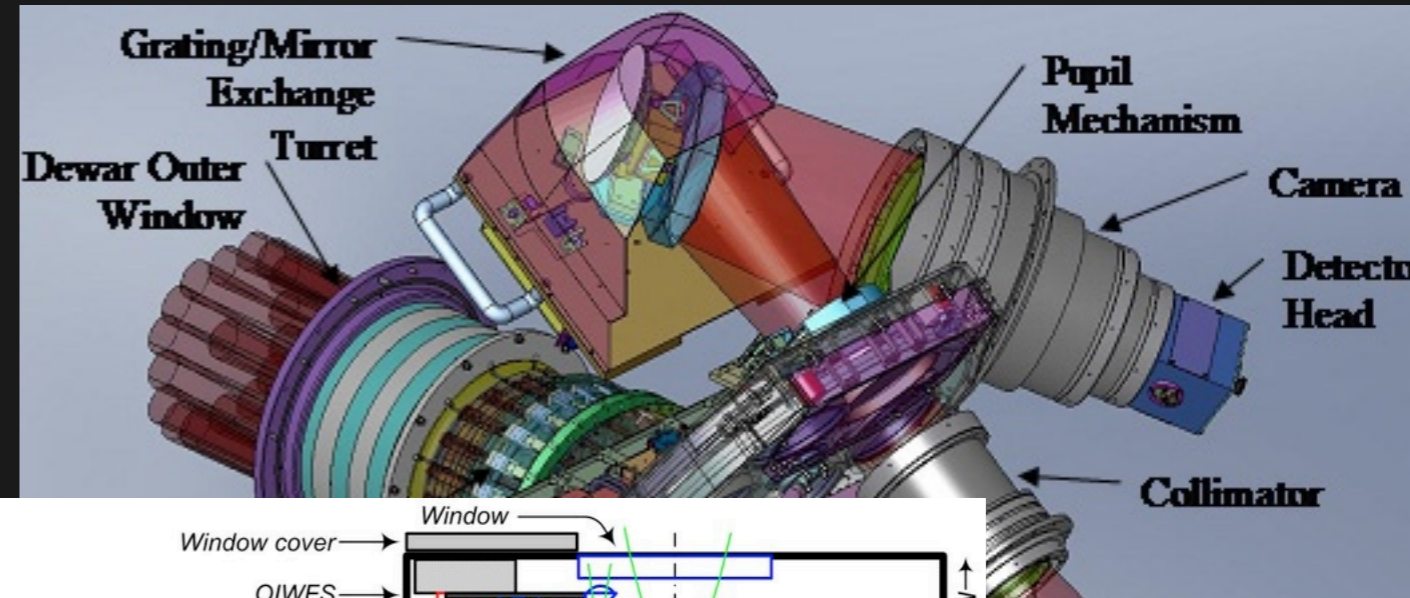
MOSFIRE for Keck I, FLAMINGOS-2 for Gemini-S

- MOSFIRE:

- 46 Cryogenic Configurable Slit Unit
 - Each 7.3" long
- FOV 6.1'
- $R=3270$ (w/ 0.7" slit), $0.975 - 2.40\mu\text{m}$

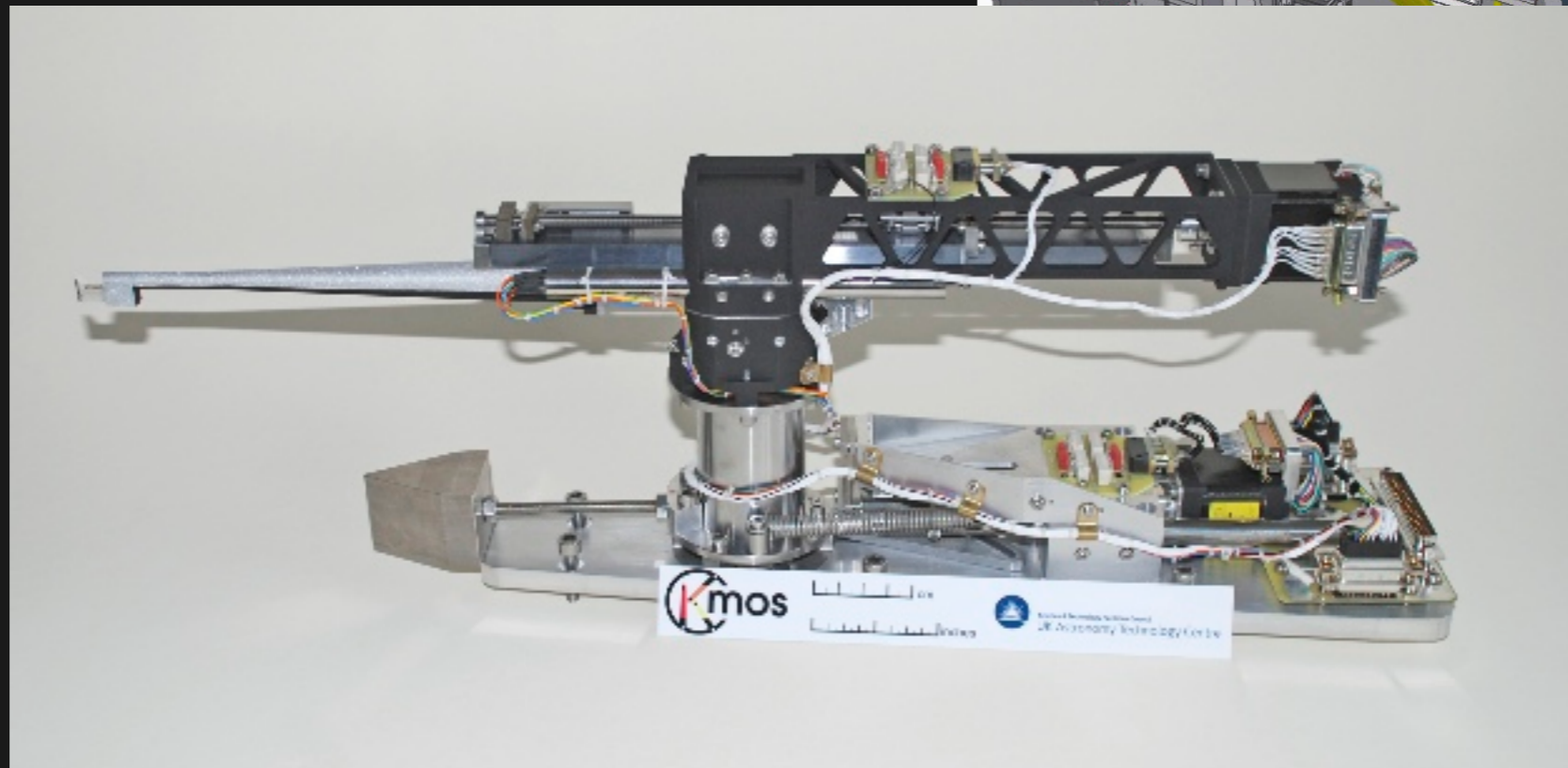
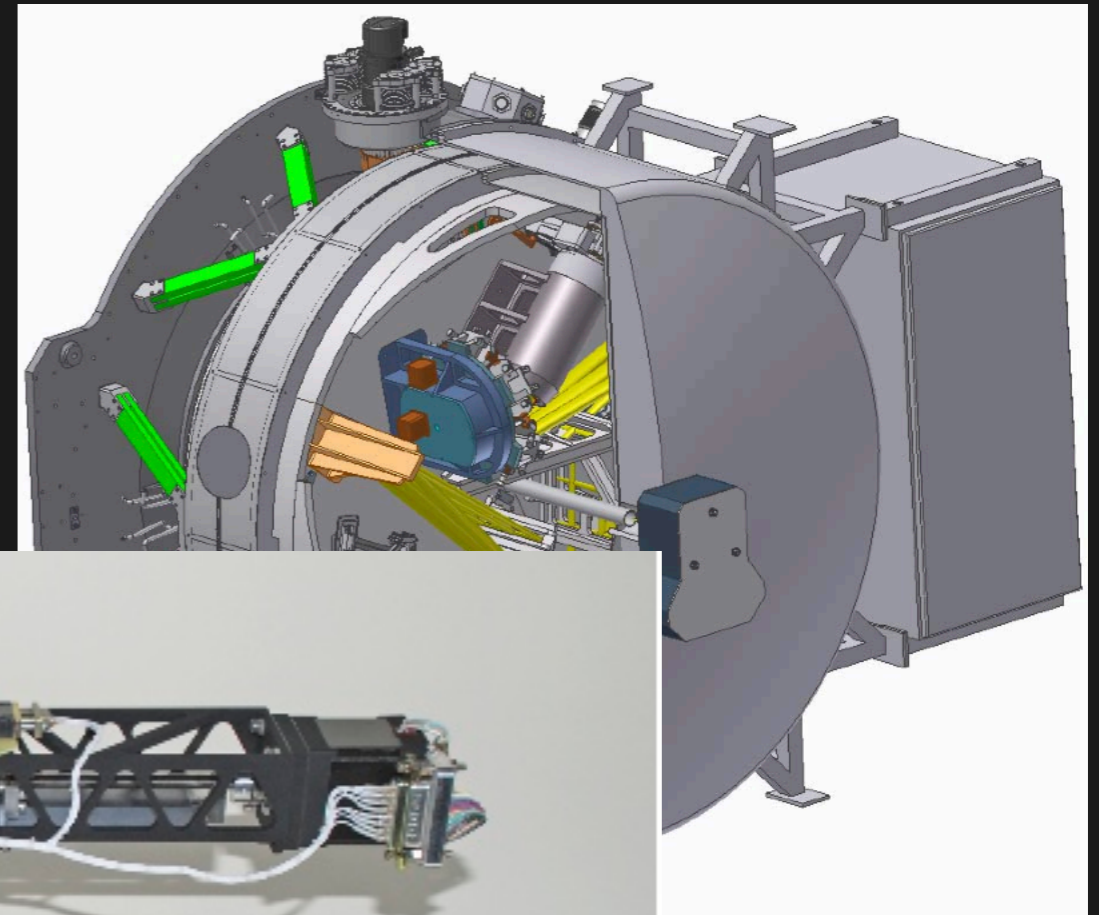
- FLAMINGOS-2:

- FOV 6.1'
- $R=1,200-3,000$
- with MCAO: $0.09''/\text{pix}$ for $2'\Phi$ FOV



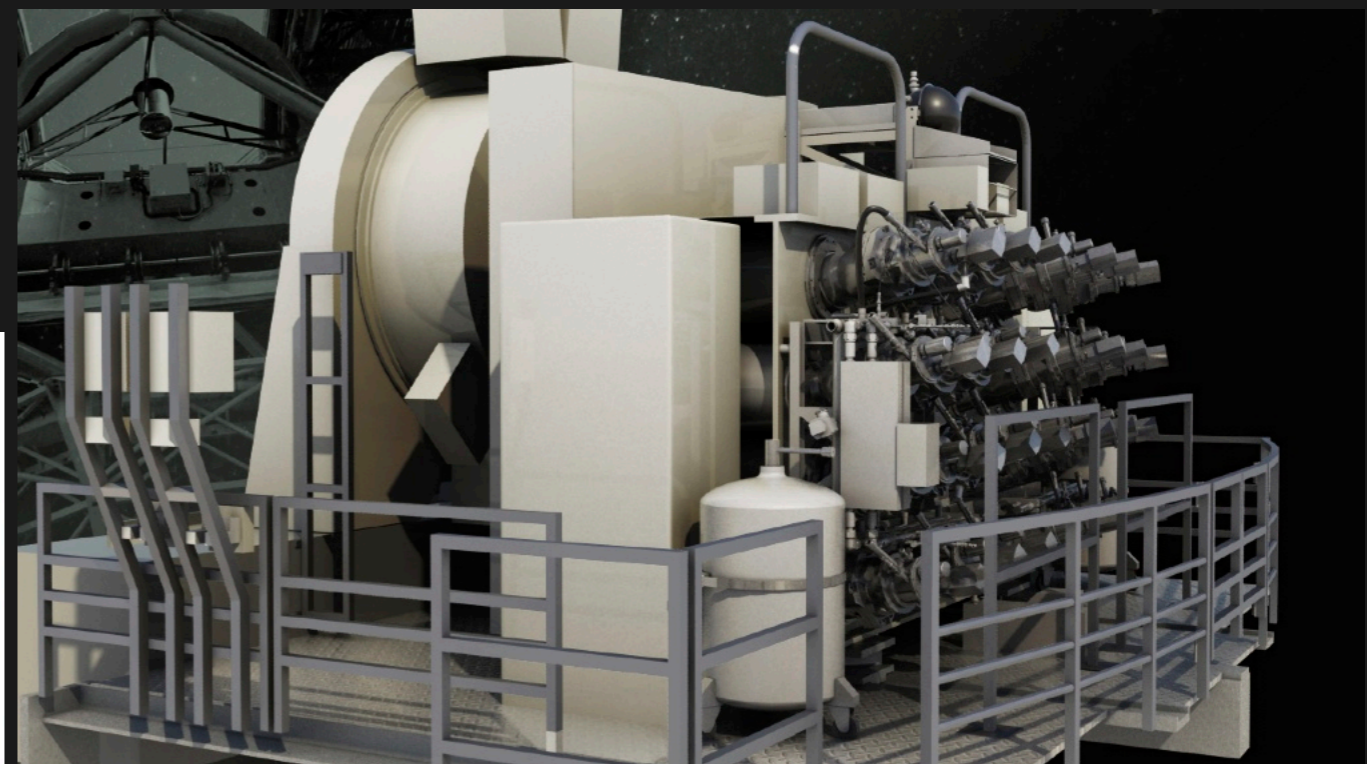
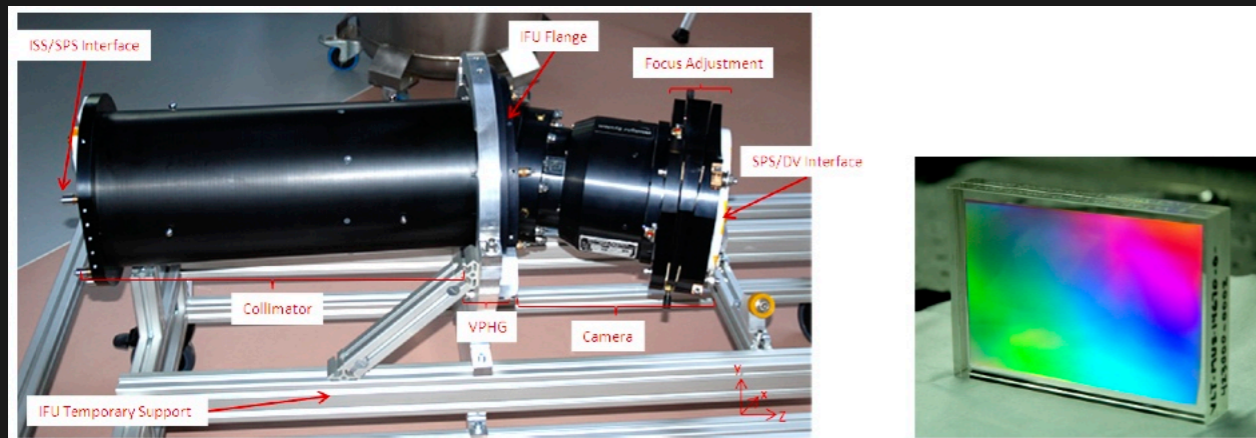
KMOS for VLT

- Near-Infrared Multi-Object IFS (Integral Field Spectrograph)
- 24 IFUs (2.8" x 2.8" each), 0.2"/pix
- Patrol Field: 7.2' Φ
- 1.0 to 2.5 μm
- $R \sim 2,000$ to 4,000
- Currently in Assembly Phase, First Light in 2011



MUSE for VLT

- Optical Multi-Object IFS (Integral Field Spectrograph)
- 24 IFUs + 24 Spectrographs
- Wide-Field Mode: $0.2''/\text{pix}$, FOV $1' \times 1'$
- Narrow-Field Mode: $0.025''/\text{pix}$, FOV $7.5'' \times 7.5''$
- 0.465 to $0.93 \mu\text{m}$
- $R \sim 2,000$ to $4,000$
- Ground-Layer AO (GALACSI), a part of AO Facility



Strategy for Subaru Next-Gen Instruments

- 将来装置計画ワークショップ Future Instrumentation Workshop
 - #1 2010/9/9-10 @ IPMU
 - PFS and Next-Generation AO
 - #2 2011/1/18 @ Mitaka
 - PI-type / Carry-in Instruments
 - Instrument Development in Universities
 - #3 2011/6/2-3 @ Mitaka

本ワークショップの目的

- すばるが活躍を続けるためには、不断の装置の更新、アップグレードが必要
 - 継続的な装置開発が必須
1. すばるへの取り付け、観測を希望する複数の装置計画のコミュニティへの紹介
 2. 大学での装置計画の在り方
 - 国立天文台(ハワイ観測所、ATC)と大学の協力、連携

Advantages of Carry-in Instruments

- 比較的少ないコストとマンパワー、短い期間で開発し、観測に結びつけることが可能
 - 装置開発の継続的な活動性の維持
- 汎用的ではないが、特定の性能に特化した仕様
- 汎用的共同利用装置が持たない機能を実現
- 次世代の大型汎用装置へのテストベッド、技術実証

Procedures for Carry-in Instruments

1. Contact Support Astronomer or Any Collaborators in Subaru
 2. Submit A Proposal to Subaru Director
 3. Subaru Internal Review Committee Makes A Report
 4. Director Reports to SAC and User Community
 5. Approval / Disapproval
- Reviews / Tests
 - Readiness Review, Engineering Acceptance Test, Final Acceptance Review
 - Carry-in Approval Does Not Guarantee Science Observing Time
 - See Details at:
<http://www.subarutelescope.org/Observing/Instruments/ApprovalProcessPInstrument.pdf>

議論すべき事項

- 大学における装置開発と国立天文台の関係
 - 国立天文台によるサポートはどうあるべきか
 - 大学間の連携
- すばる望遠鏡用装置開発の在り方
 - TMTにむけた装置開発との関係
- 装置が大型化し、開発が長期化する中で人材を育てていくためには

Strategy for Subaru Next-Gen Instruments

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#1 2010/9/9-10 @ IPMU

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「すばる次世代AOワークショップ」

次世代赤外線装置 Next-Gen IR Instrument

- 2010年代 - 2020年代にすばるに必要な赤外線装置は?
 - 多天体分光能力の強化
 - 面分光機能
- 次世代の補償光学は?
 - Ground-Layer AO による広視野での解像度の向上
 - 多天体AO (MOAO)
- スペース、TMTがあっても価値のある装置とは?

次世代赤外線装置 Next-Gen IR Instrument

- Working Group Just Founded
 - GLAO, MOAOの基礎的study
 - 次世代AOの仕様案策定 = 望遠鏡機能のUpgrade案
- 次世代AOを活かす装置案の設定
- サイエンス面からの要請
 - 6/2-3 「次世代AOワークショップ」にて