

# SWIIMS

## A New NIR MOS Spectrograph for TAO 6.5m Telescope

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# TAO

## The University of Tokyo Atacama Observatory Project

- \* Construct an infrared-optimized 6.5m telescope at the world's highest site of 5650m(18000ft) altitude, Co. Chajnantor at Atacama desert, Chile
- \* Current status:
  - \* Installed a 1m telescope (called miniTAO), and started observations in the near- to mid infrared



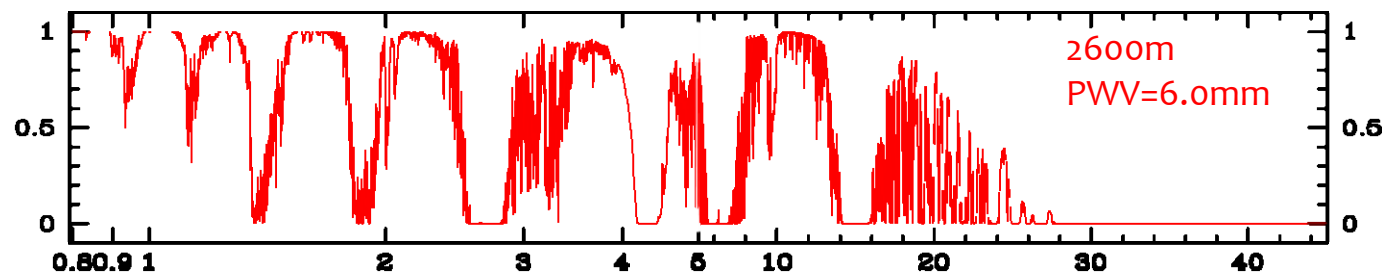
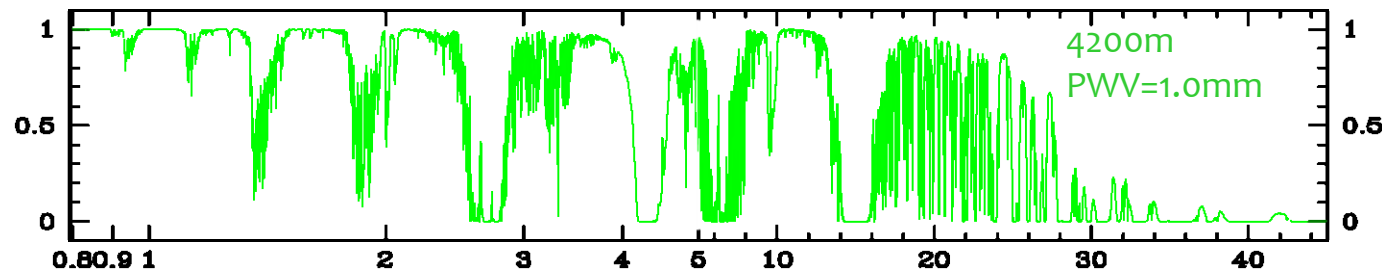
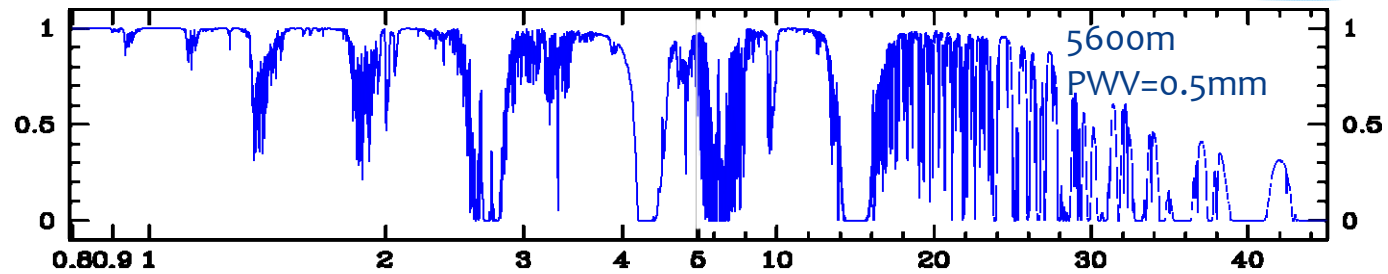


# Advantage of Co. Chajnantor

- \* Nice climate, with photometric nights  $> 60\%$ , and usable  $>80\%$  (Miyata+08)
- \* Seeing is as good as Mauna Kea; DIMM measurement shows median seeing of  $0.7''$  (KM+08)
- \* Thanks to the low pressure ( $<0.5\text{atm}$ ) and dry climate, extremely low perceptible water vapor (PWV) is achieved :  $0.5\text{mm}$  (25%tile) (MK has  $1\text{mm}$ )
- \* This results in excellent observational condition in both NIR( $1-2.5\mu\text{m}$ ) and MIR ( $5-40\mu\text{m}$ )



# Atmospheric Transmittance

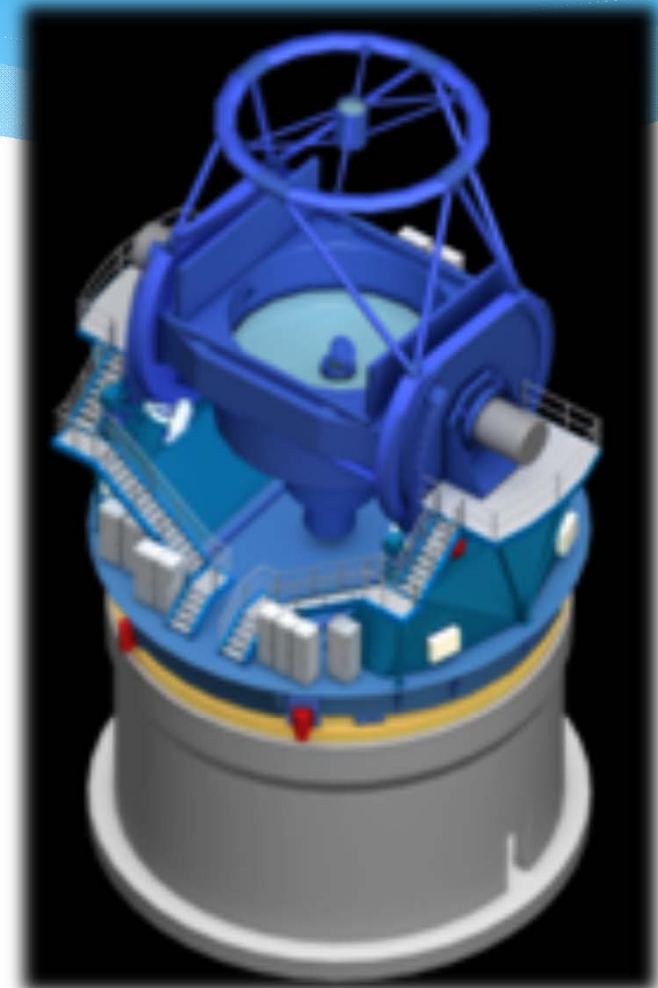




# The 6.5m Telescope

Optics	Ritchey-Chretien
Primary Mirror	$\Phi$ 6,500 mm (effective $\Phi$ 6,154 mm)
2ndry Mirror	$\Phi$ 897 mm
Foci	Cassegrain, Nasmyth x 2
FoV	$\Phi$ 25'
PM F-ratio	1.25
Final F-ratio	12.2

**Synergy with  
the Subaru Telescope**

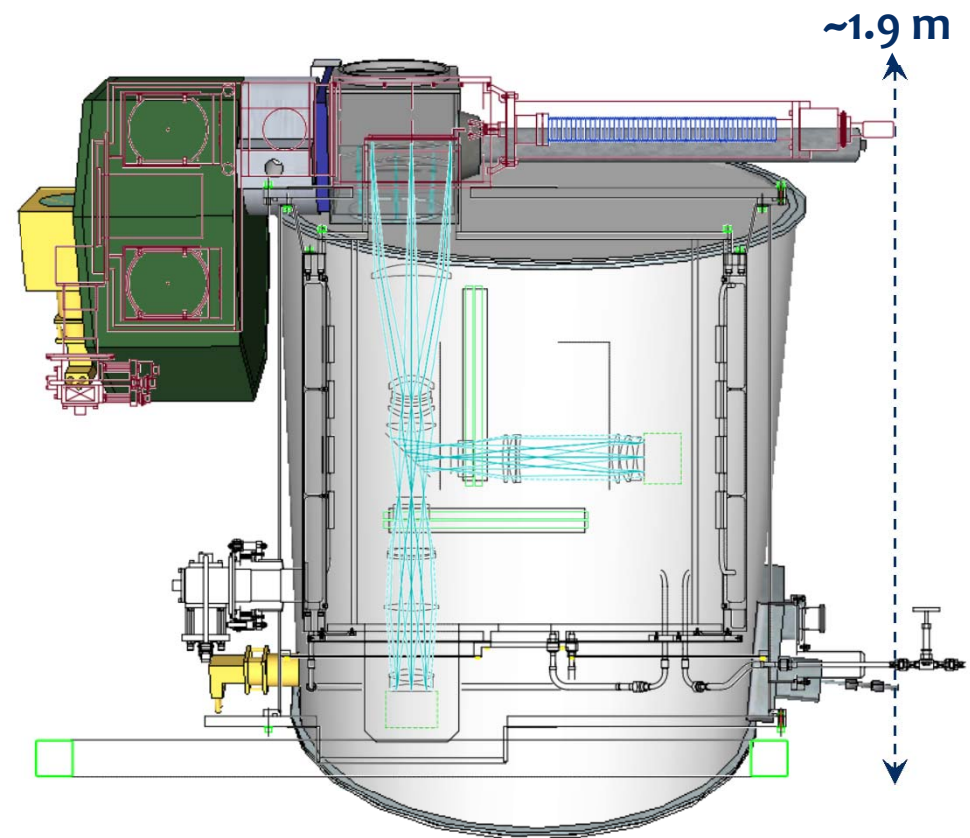




# SWIIMS

## Simultaneous-color Wide-field Infrared Imager and Multi-object Spectrograph

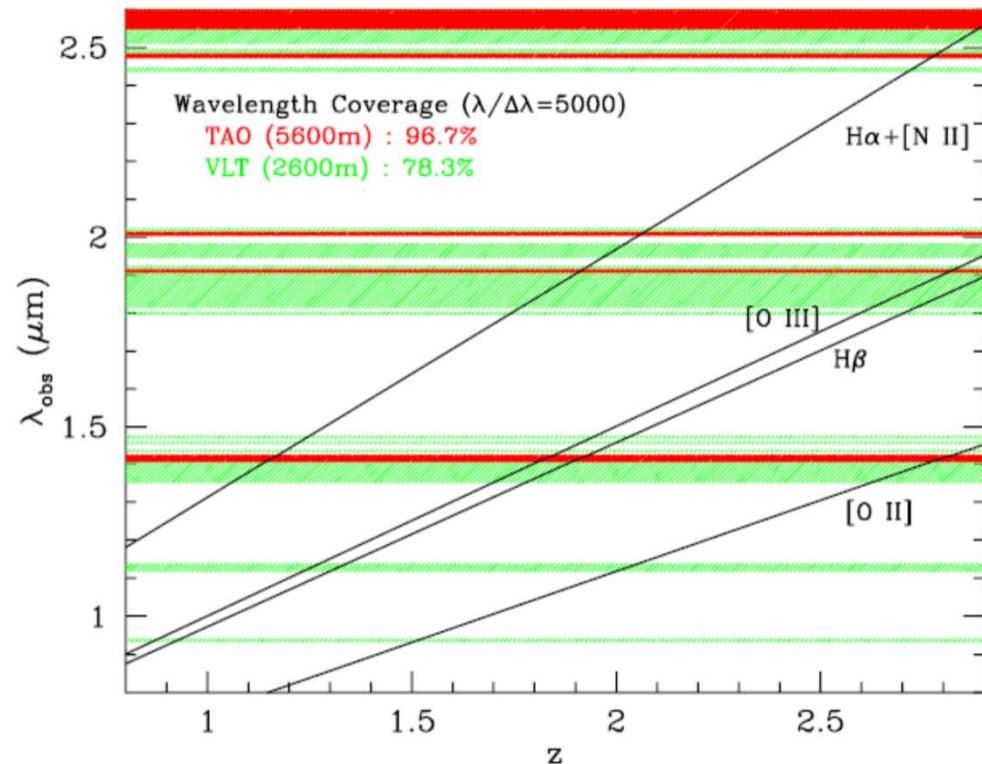
- \* NIR Spectrograph for TAO  
6.5m
- \* **Wide-Field** (9.6'  $\Phi$  FoV on TAO)
- \* **MOS** Spectrograph
- \* **2-band simultaneous** Imaging/Spectroscopy  
: 0.9-1.4 and 1.4-2.5 $\mu$ m





# Science with SWIIMS

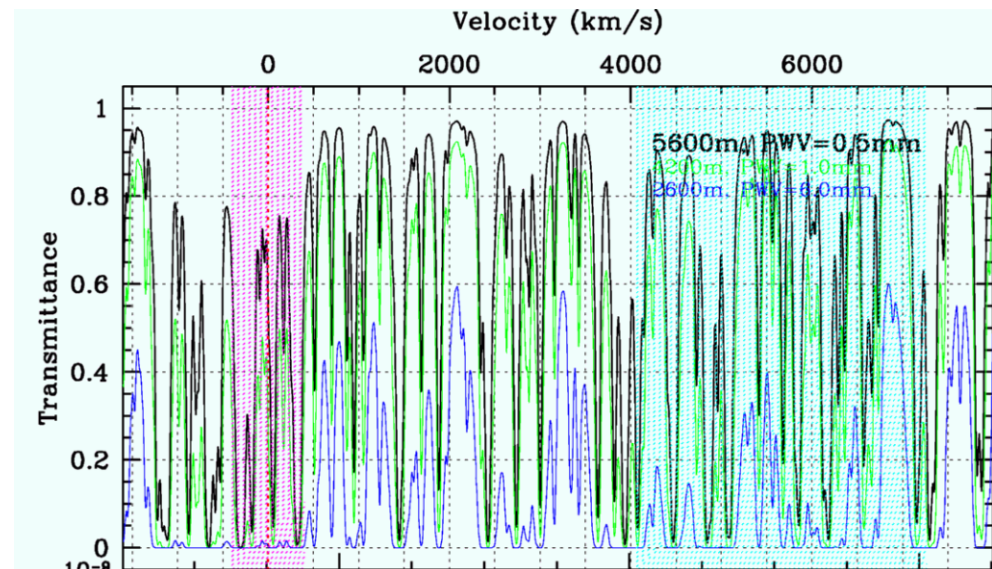
- \* High transmittance at TAO site  
: Ideal for the **redshift survey** in the infrared
- \* Full 0.9-2.5mm spectra can be obtained simultaneously  
: **precise continuum / line ratio measurement** possible
- \* Simultaneous Hi-z Narrow-band Imaging (e.g. H $\alpha$ /OIII, H $\beta$ /OII)





# Paschen $\alpha$ at Co. Chajnantor

- \* Hydrogen Pa $\alpha$  line (1.8751 $\mu$ m) available
  - \* Strongest Hydrogen Line in NIR
  - \* Good probe in dusty environment





# Pa $\alpha$ Imaging with ANIR at Co. Chajnantor



Galactic Center (Sgr A\*)

Starburst Galaxies IC4686/4687



J, H, Ks



J, H, Ks+Paschen- $\alpha$



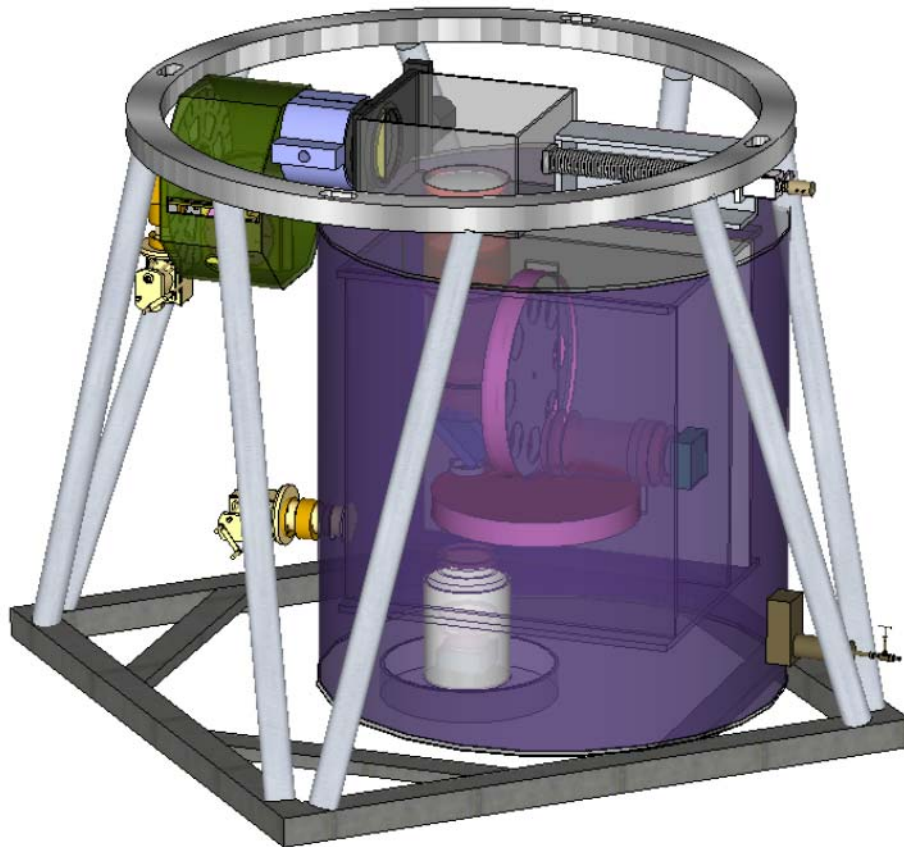
# 1<sup>st</sup> SWIIMS Science Workshop

(Sep. 11, 2009)

- \* Hi-z Cluster Survey by NBFs
- \* YSO eclipsing binary survey
- \* Elliptical Galaxy distance measurement by Mira variables
- \* Chemical properties of Hi-z QSOs
- \* SN follow-up
- \* AGN monitoring
- \* Synergy with ALMA
- \* Synergy with HSC
- \* GRB follow-up
- \* ...



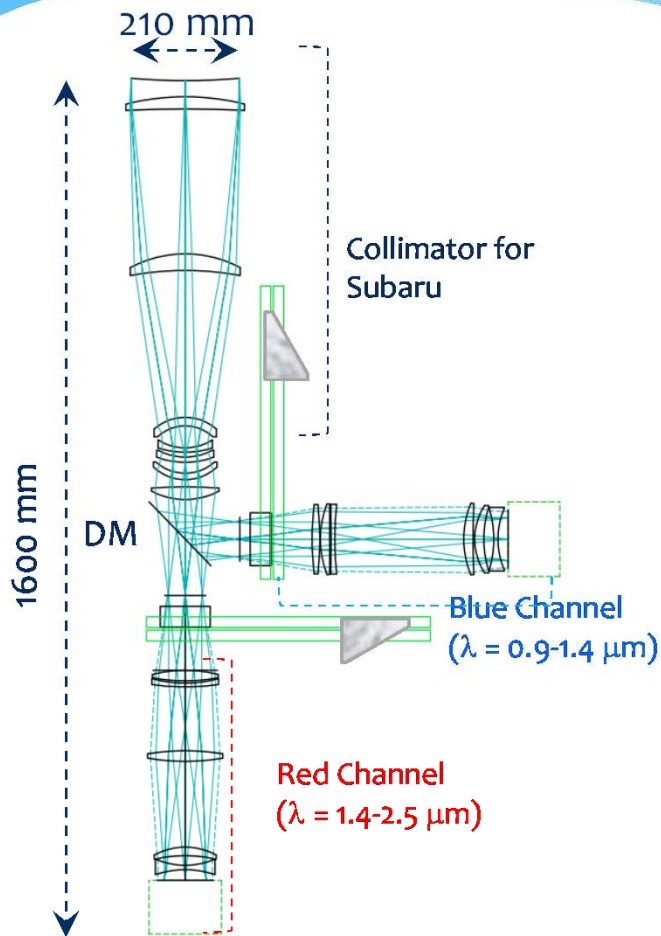
# SWIIMS on Subaru/Cassegrain



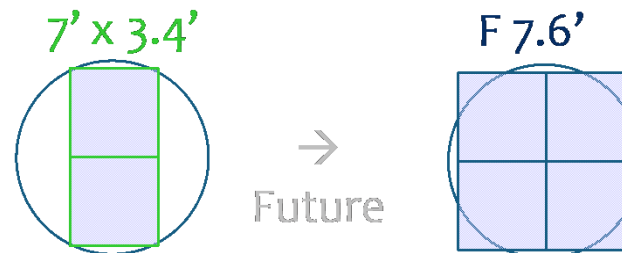
- \* TAO 6.5m not available before 2015
- \* First science with Subaru
- \* Collimator design optimized for Subaru
- \* FoV of 7' x 3.4' (MOIRCS : 7'x4')



# SWIIMS Optics for Subaru

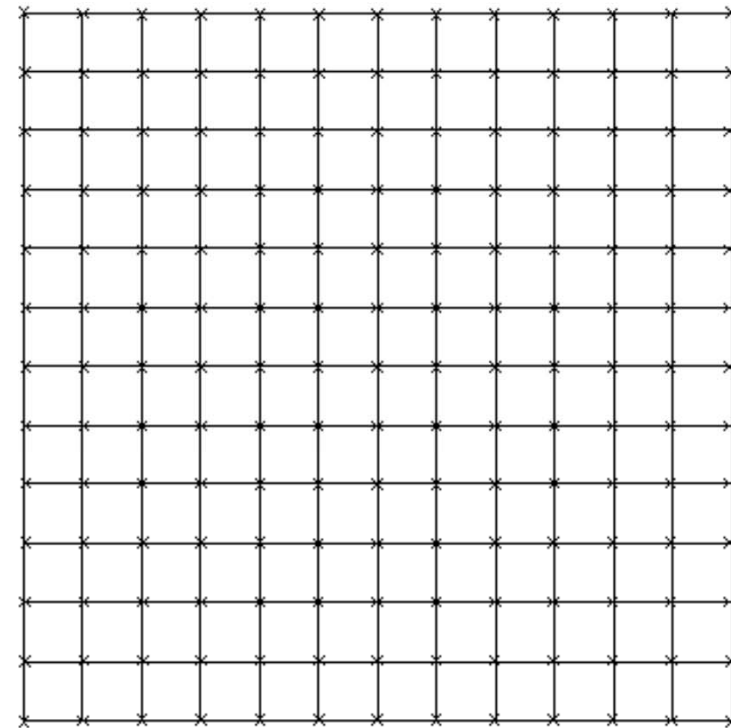
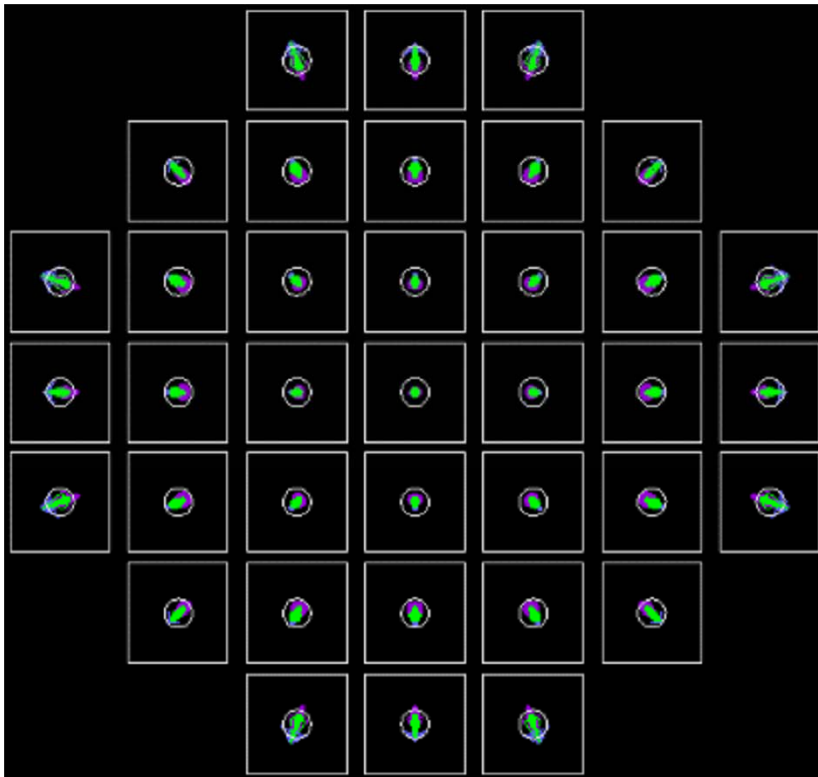


- \* Split collimated beam using a Dichroic Mirror(DM) at 1.4mm
- \* Final F-ratio=4.8
- \* As only 4 arrays were procured, each focal plane is covered by 2 HAWAII2-RGs with FoV of 7'x3.5'





# Performance of Optics (Red Channel)

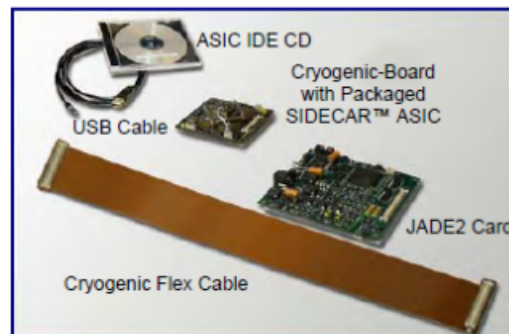




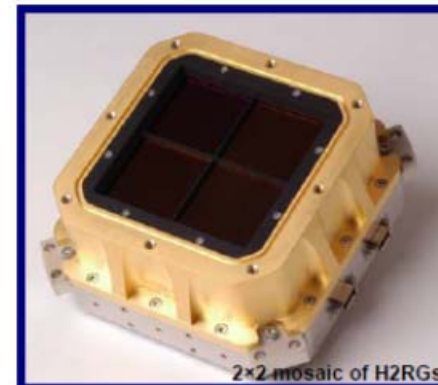
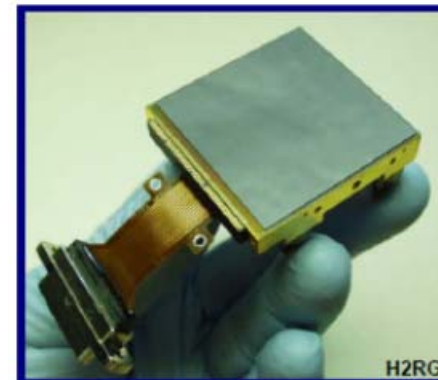


# Focal Plane Arrays

- \* 4 HAWAII2 RGs /  $2.5\mu\text{m}$  cutoff
- \* SIDECAR ASICs for readout
- \* Covers central  $7' \times 3.5'$  FoV
- \*  $0.1''/\text{pix}$  sampling



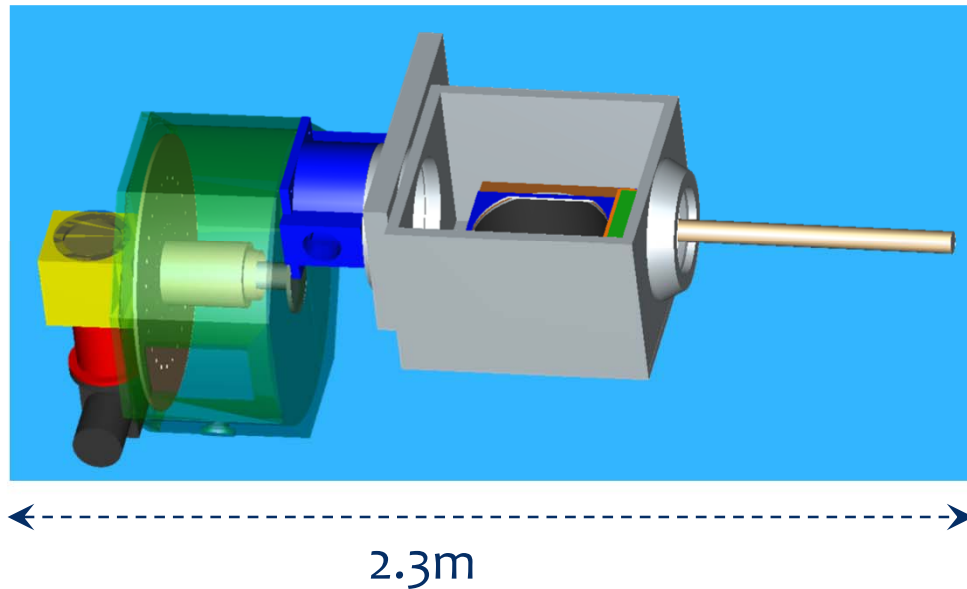
Contents of the ASIC Cryogenic FPE Kit (Temp: 32 to 300 K)





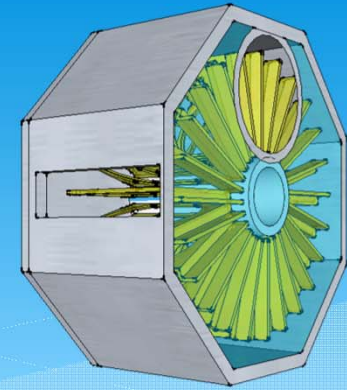
# MOS Slit Plate Exchanger

- \* Based on the design of MOIRCS
  - \* Turret-type mask stocker
  - \* Robotic mask exchanger

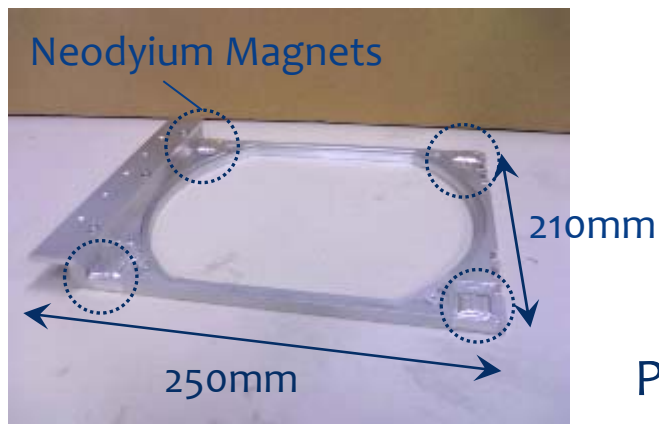




# MOS Unit



- \* Maximum # of mask slit : 20 or more
- \* New design with high reliability
  - \* Magnet – latch (Both at the focus and the stock)
  - \* Operative even at the TAO Nasmyth focus
  - \* Curved mask possible



Prototype Mask Holder





# Schedule

- \* 2010/Q3 : Fix the design of dewar, MOS, optics
- \* 2010/Q4 : Array delivery
- \* 2011/Q1 : Delivery of the dewar, MOS, optics componetns
- \* 2011/Q2-2012/Q2 : Assembly, adjustment, etc...
- \* 2012/Q3 or later : Transport to Subaru, setup
- \* 2013/Q1 : First Light at Subaru ?