Toward WFOS IFU from FOCAS IFU

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Our project overview

FOCAS@Subaru





2025/01/30

WFOS@TMT



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FOCAS IFU

- FOCAS is an optical imaging spectrograph on the Subaru telescope.
- Development of FOCAS IFU was started in 2010 and completed in 2018.
- FOCAS IFU has been open for common use since 2019.
- This development was supported by
 - TMT strategic R&D fund (TMT戦略基礎開発経費) (FY2012, FY2013, FY2014, FY2016, FY2017) and,
 - KAKENHI Grant-in-Aid for Scientific Research (C) (科研費基盤C) (FY2014-FY2016).



FOCAS IFU parameters



Slice number	23
Slice widht	0.435"
Slice length	13.5"
Field of View	13.5" x 10.0"



Main optical components in FOCAS IFU



Image slicer



Pupil mirror array



These opt-mechanics were fabricated in NAOJ ATC, and these components were assembled in NAOJ ATC.



Slit mirror array

Vignetting of FOCAS IFU

- The vignetting was mainly caused by the assembling error of the slicer due to tight tolerance.
- We started R&Ds toward WFOS IFU right after the completion of FOCAS IFU.



Wide Field Optical Spectrograph (WFOS)

- The development is led by Caltech.
- Observing modes:
 - Imaging
 - Multi-slit spectroscopy
 - (Integral filed spectroscopy)
- FoV = 8.3 arcmin \times 3 arcmin
- R~1500, 3500, 5000@0.75" slit
 - R~4500, 10500, 15000@0.25" slit



WFOS IFU parameters

Considering a slit loss, IFU has great advantage against the slit spectroscopy.



Number of slices

Slice length (arcsec)

Slice width (arcsec)

R=1500 grating

R=3500 grating

FoV (arcsec x arcsec)

R=5000 grating (goal)

Optical IFUs on other ELTs

GMT



GMACS



MANIFEST



MANIFEST is a fiber feeding module and provides an IFU capability for GMACS.

E-ELT



HARMON



HRMONI for ELT is mainly for near IR, but its sensitivity extends to optical range.

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Comparison with other IFUs on ELTs





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Science cases

- Cosmic web
- Circum-galactic medium (CGM)



Tumlinson et al. (2017, ARA&A, 55, 389)



Expected Ly α surface brightness (Yajima et al. 2020) Lowest level: 1 σ detection limit for 10-hour exposure ~3x10⁻²⁰ erg/s/cm²/arcsec²

WFOS IFU status





Rotating structure in WFOS

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R&D efforts toward WFOS IFU

- TMT strategic R&D fund (TMT戦略基礎開発経費) (FY2019, FY2022, FY2023)
- NAOJ Leadership fund(国立天文台リーダーシップ経費)(FY2020, FY2022)
- KAKENHI Grant-in-Aid for Scientific Research (B) (科研費基盤B) (FY2024-FY2027). to



Optical layout with loosetolerance



Improvement of the slicer assembling accuracy.



Technology verification IFU (Tech. IFU) (See P20)



Reflectivity measurement system



Flat-fielding error correction

Summary

- We developed FOACS IFU.
- Based on the experiences, we are conducting some R&D efforts toward WFOS IFU.
- Our activity has been supported by **TMT strategic R&D fund**, KAKENHI and NAOJ Leadership fund.
- I am very happy if TMT strategic R&D fund supports us continuously!

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