

Enhancing scientific discoveries using Photonic Lanterns on the SCExAO system

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S. Sallum, P. Gatkine, T. Kotani, M. Tamura, E. Spalding, Y. Xin*



2023 Subaru Users Meeting



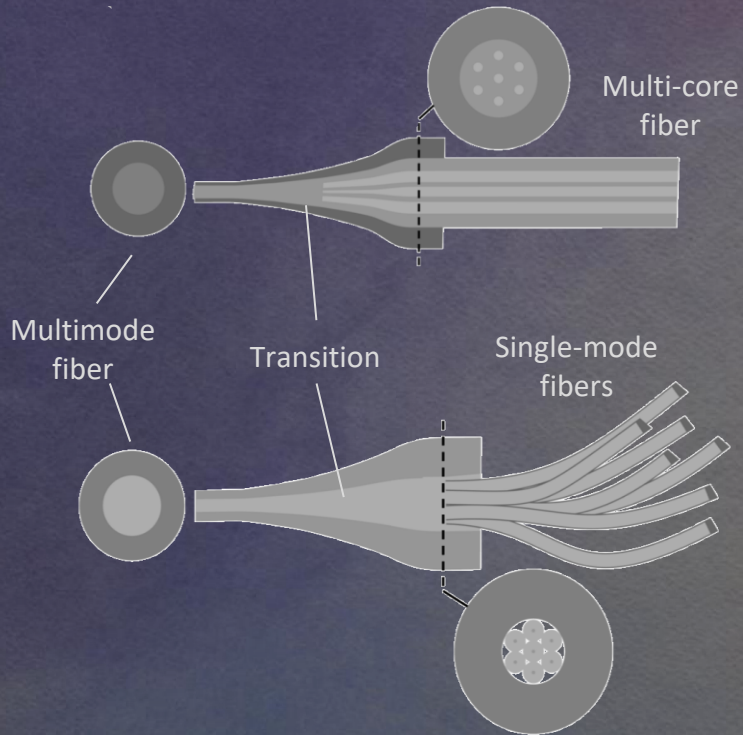
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Natural Sciences
**Astrobiology
Center**



What is a Photonic Lantern ?

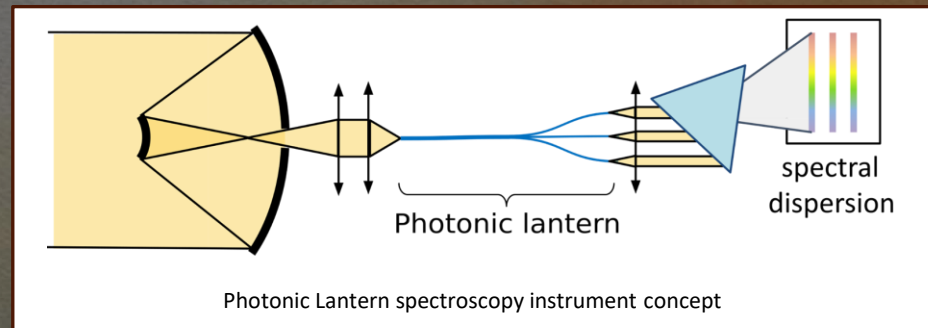


Schematics representation of two different approaches for the fabrication of Photonic Lanterns; Top: Multicore fibre approach; and Bottom: Standard single-mode fiber combiner/splitter technique.

From Leon-Saval et al. "Photonic Lanterns: a study of light propagation in multimode to single-mode converters" OSA 2010

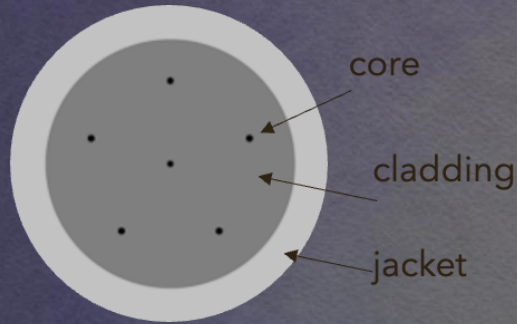
Photonic Lantern (PL) = Fibered device with Multi-Mode input and several Single-Mode outputs

- ❖ Adiabatic transition between MM to SM mode very efficient ($>90\%$, Birks et al. 2015)
- ❖ Allows for SMF-fed spectroscopy with high throughput



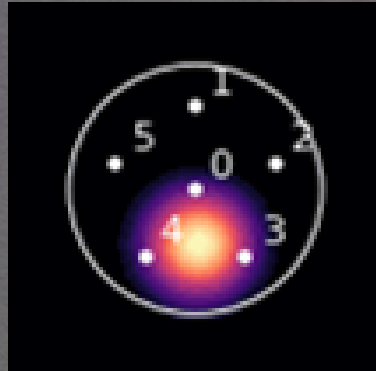
Properties of Photonic Lanterns

The intensity distribution at the SMF outputs depends on the **input scene**



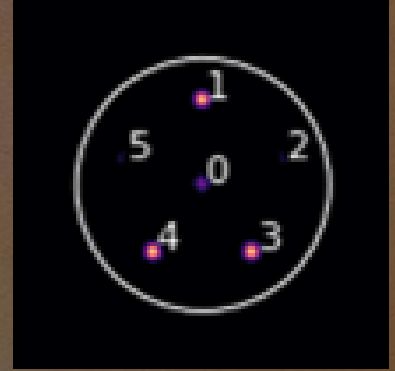
Photonic Lantern with
6 SMF outputs

PSF at the multi mode input



Photonic Lantern

Intensities at the single mode
outputs

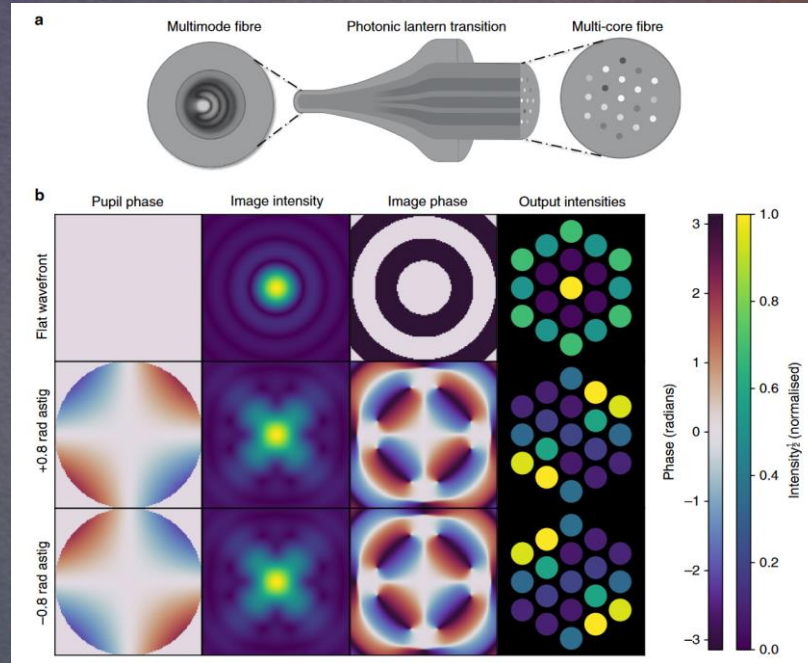


Simulation by Yoo-Jung Kim

→ PL outputs can be used for **constraining the input scene for science applications.**

Properties of Photonic Lanterns

The intensity distribution at the SMF outputs depends on the **input wavefront**



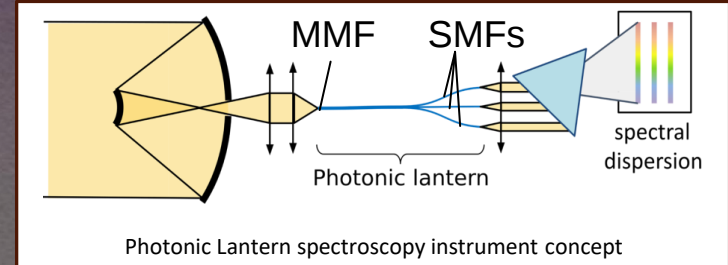
B. Norris et al, Nature, 2020

→ PL outputs can be used for **wavefront control**

Photonic Lanterns capabilities in a nutshell

High throughput

Enabling SMF-fed spectroscopy with high efficiency (especially in the Visible)



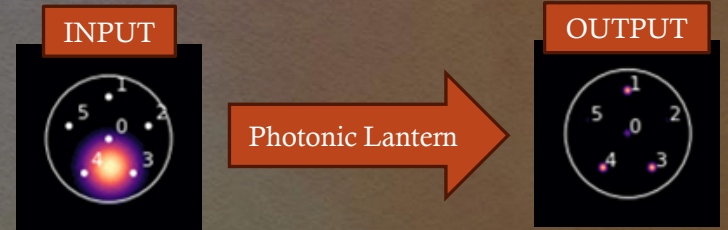
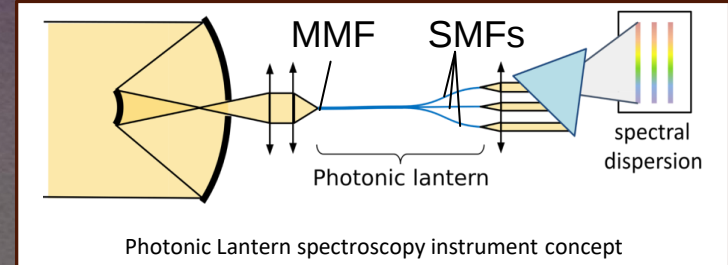
Photonic Lanterns capabilities in a nutshell

High throughput

Enabling SMF-fed spectroscopy with high efficiency (especially in the Visible)

Sensitive to input scene

Output flux distribution depends on the observed object

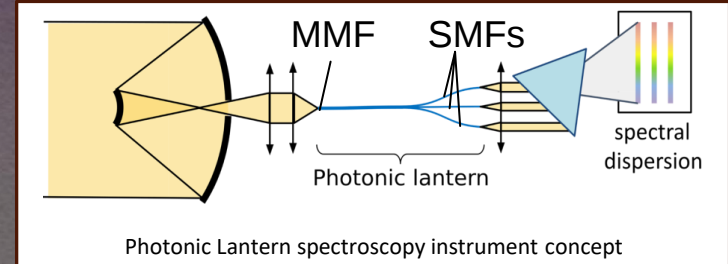


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Photonic Lanterns capabilities in a nutshell

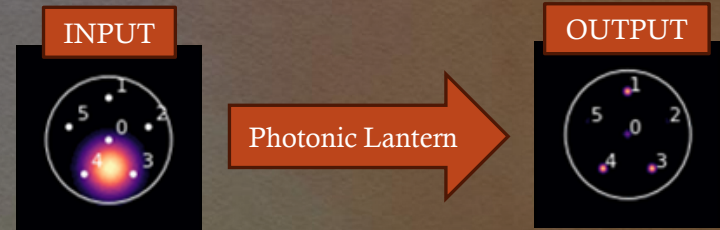
High throughput

Enabling SMF-fed spectroscopy with high efficiency (especially in the Visible)



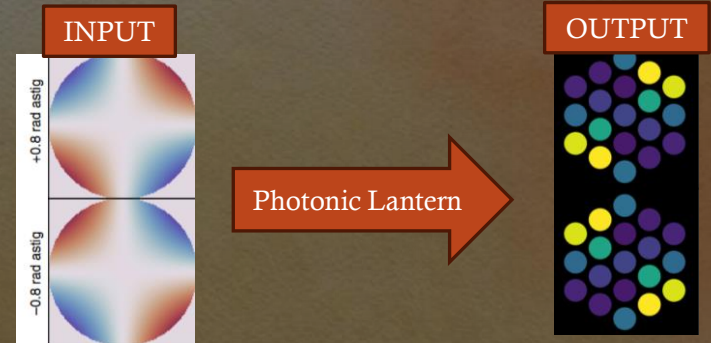
Sensitive to input scene

Output flux distribution depends on the observed object



Sensitive to input wavefront

Output flux distribution depends on the input wavefront



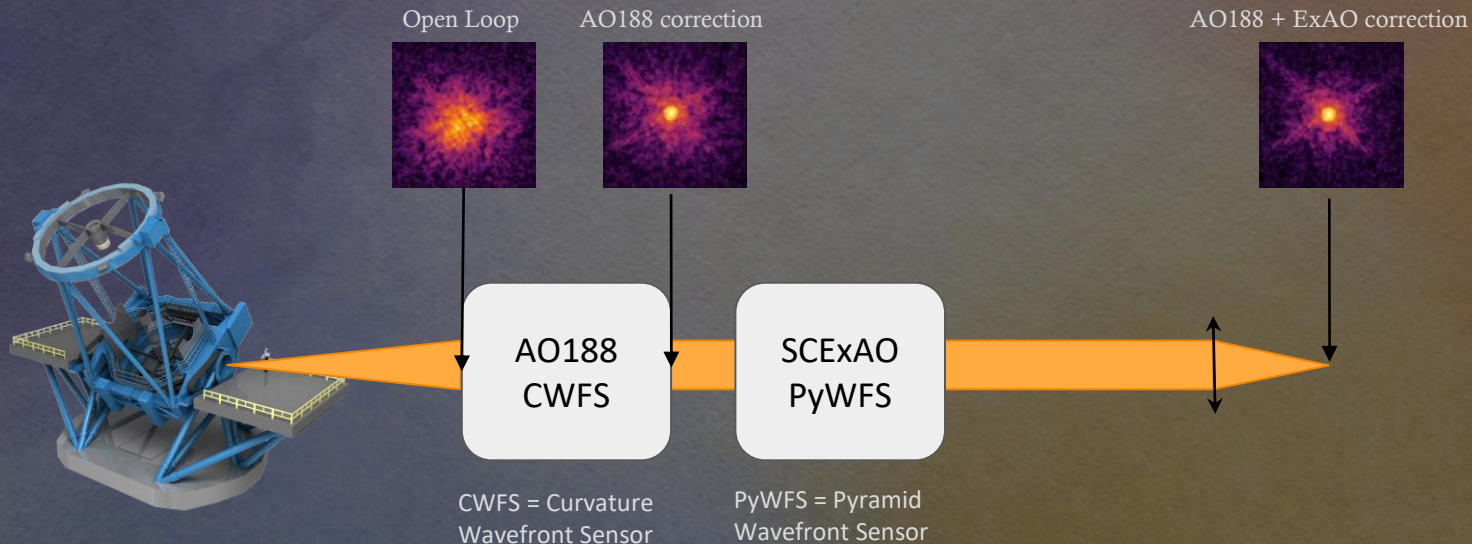
Science applications – SMF-fed spectroscopy

- **Spectrally dispersed wavefront sensing and close-loop control (Island Effect)**
J. Lin
- **Image reconstruction, spectro-imaging**
Y-J Kim, B. Norris, O. Guyon
- **Efficient injection into photonic devices (FIRST)**
E. Huby, S. Lacour + FIRST team, Y-J Kim
- **Emission line source detection (H-alpha accreting planets)**
M. Lallement, S. Vievard, O. Guyon
- **Emission line source characterization (H-alpha accreting planets - injection into RHEA R~60,000)**
S. Vievard, O. Guyon
- **Spectro-astrometry**
D. Levinstein
- **Extreme Precision Radial Velocity (EPRV)**
T. Kotani, M. Tamura, O. Guyon

Photonic Lanterns on SCExAO

SCExAO = Subaru Coronagraphic Extreme Adaptive Optics

→ High contrast imaging instrumentation

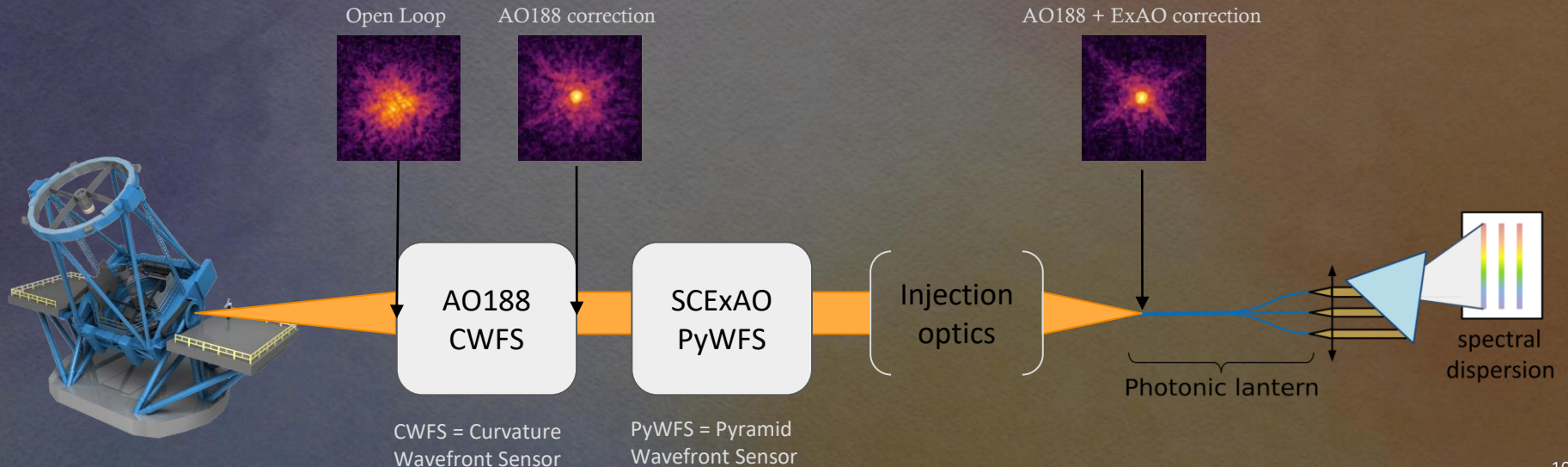


Photonic Lanterns on SCExAO

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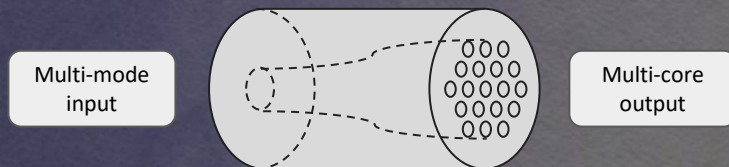
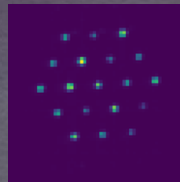
→ High Strehl PSF injected into the multi-mode photonic lantern input + single-mode fiber-fed spectroscopy



Photonic Lanterns on SCE_xAO

Near InfraRed

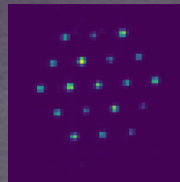
- 19-port photonic lantern (No dispersion)



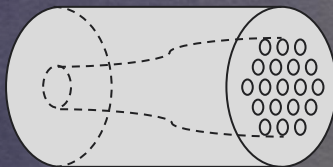
Photonic Lanterns on SCE_xAO

Near InfraRed

- 19-port photonic lantern (No dispersion)

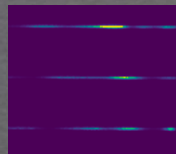


Multi-mode
input

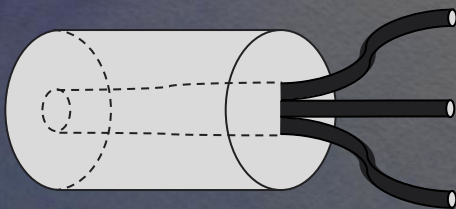


Multi-core
output

- 3-port photonic lantern ($R \sim 300$ spectro.)



Multi-mode
input

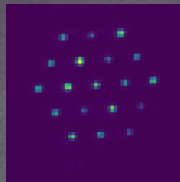


Multi-SMFs
output

Photonic Lanterns on SCE_xAO

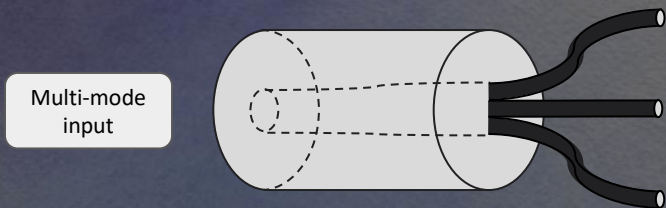
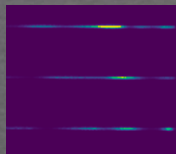
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Multi-core
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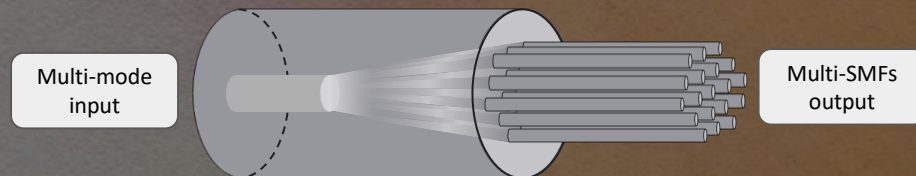
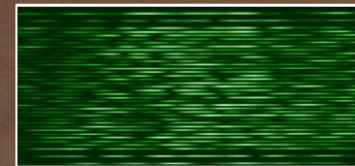
- 3-port photonic lantern ($R \sim 300$ spectro.)



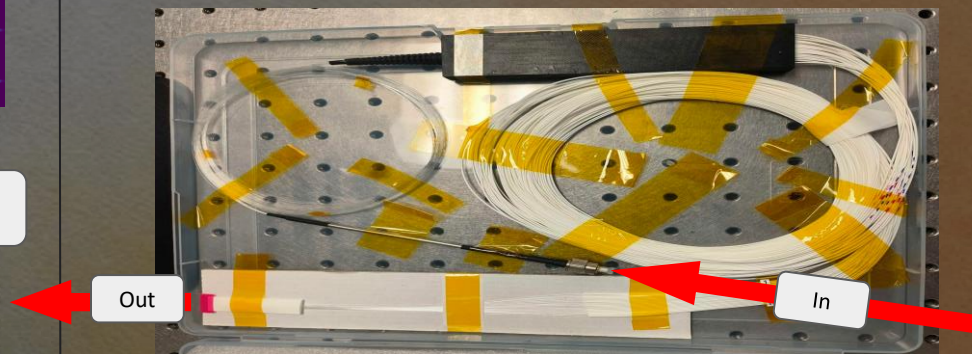
Multi-SMFs
output

Visible

- 19-port photonic lantern ($R \sim 4000$ + polarization splitting)



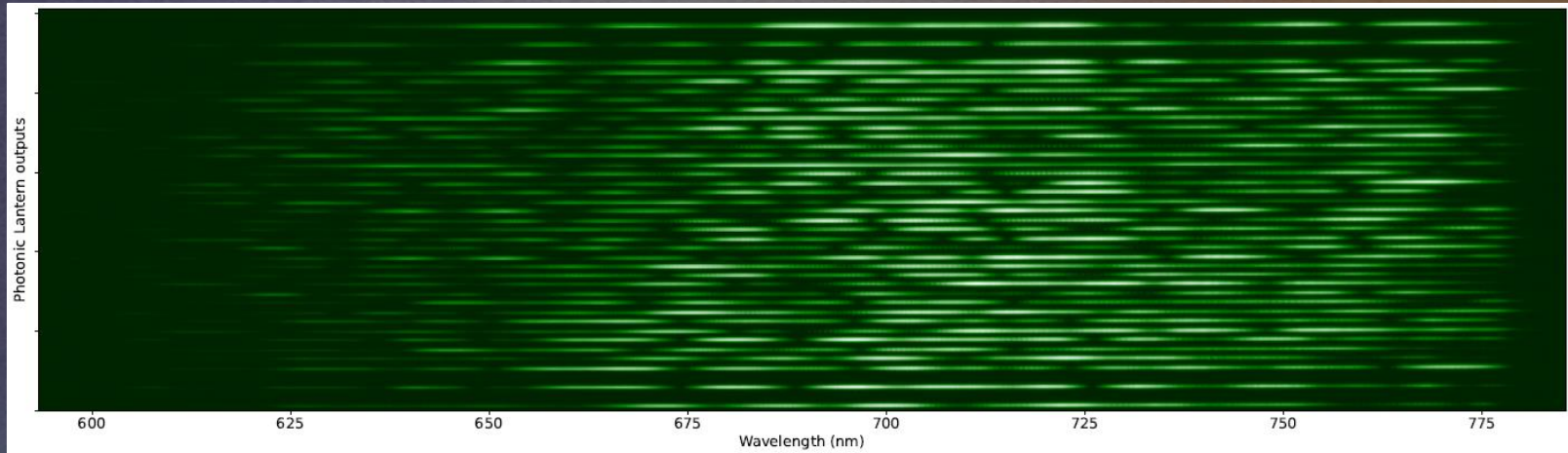
Multi-SMFs
output



Visible Photonic Lantern – Data acquisition

38 spectra on the detector (19 x 2 polarizations)

- **Flux repartition** between the PL outputs depends on **the source position/shape** and the **injected wavefront**
- Spectral structures appear depending on the injected PSF size ($\sim$ injected beam focal ratio or F#)



PL imaging the SCExAO internal super continuum source (F#4)

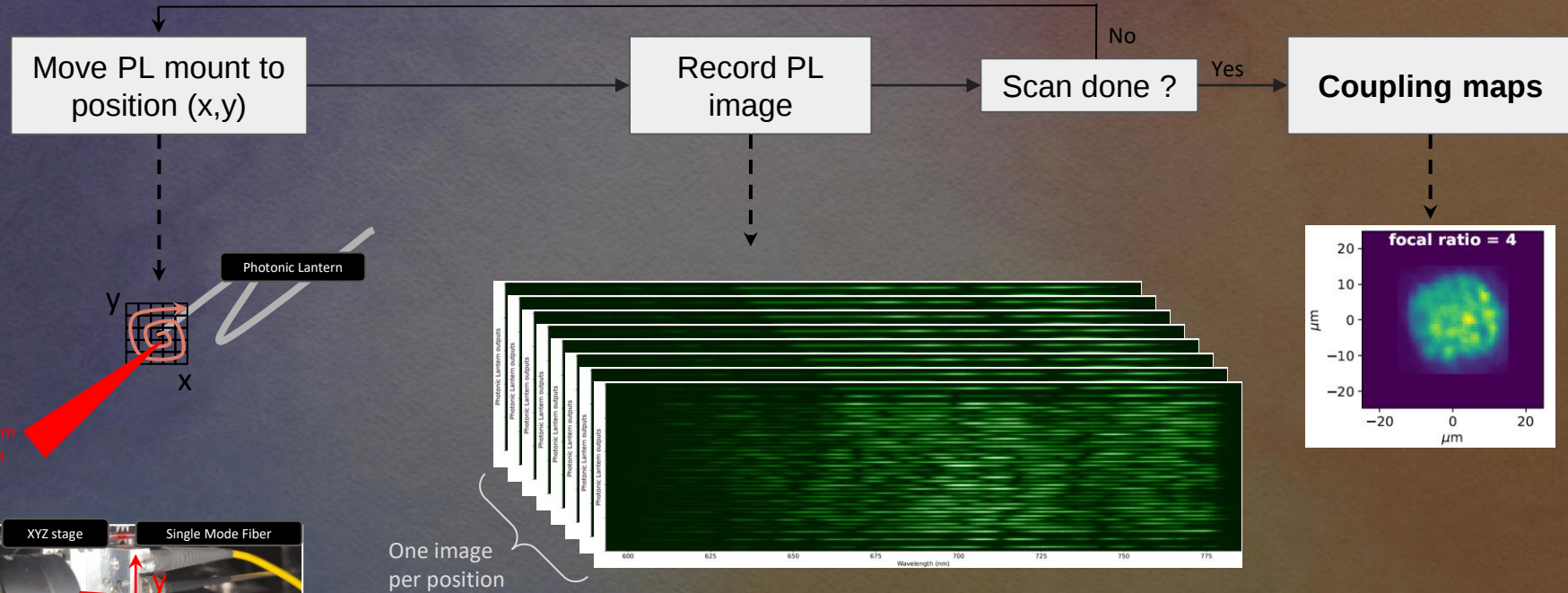
→ How to properly inject light into the Photonic Lantern ?

Visible Photonic Lantern – Coupling maps

To efficiently couple light into the PL, the PL core ($\sim 25\mu\text{m}$) must be centered on the injected PSF. This optimal position of the PL is found using **coupling maps**.

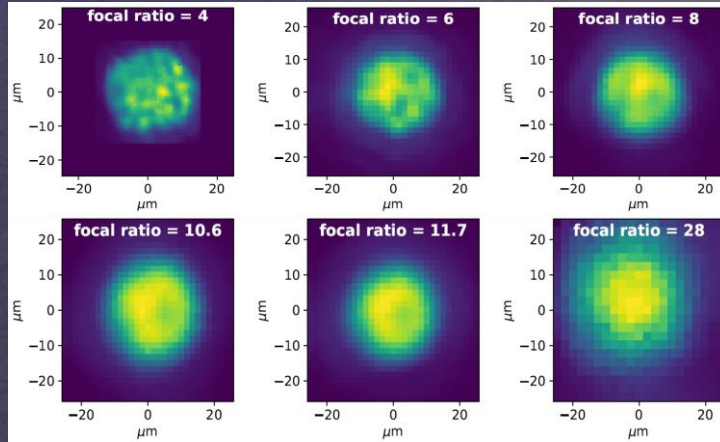
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Visible Photonic Lantern – Coupling maps

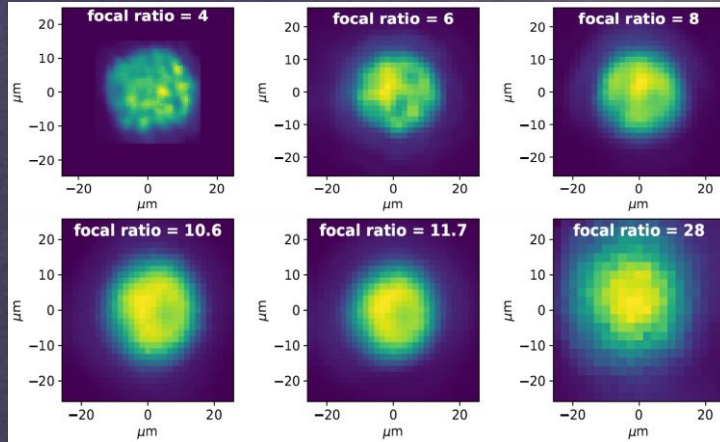
- ❖ Coupling maps for various focal ratios (or PSF size)



→ Less structures when focal ratio increases

Visible Photonic Lantern – Coupling maps

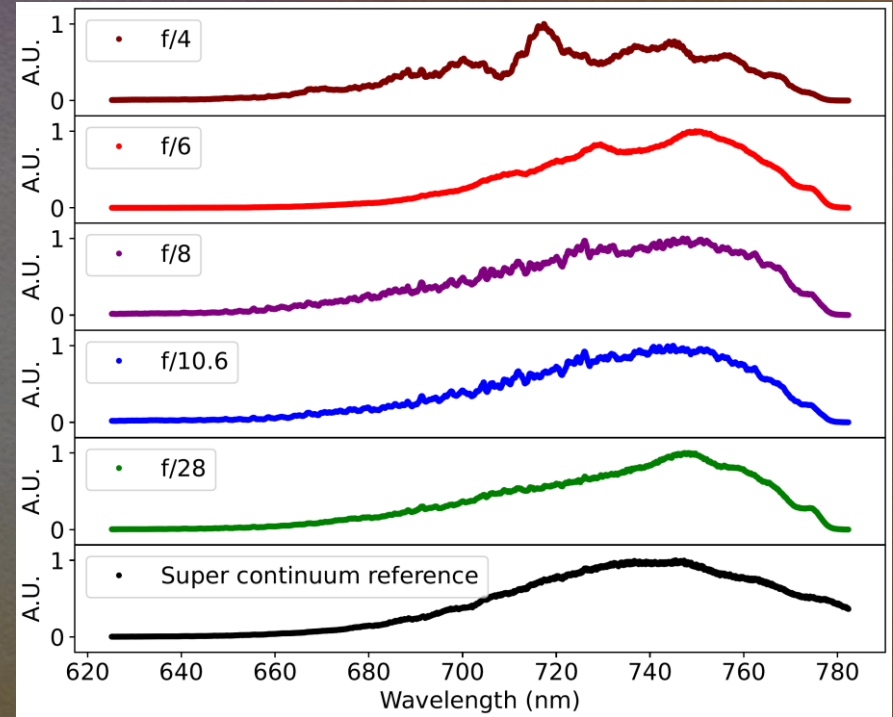
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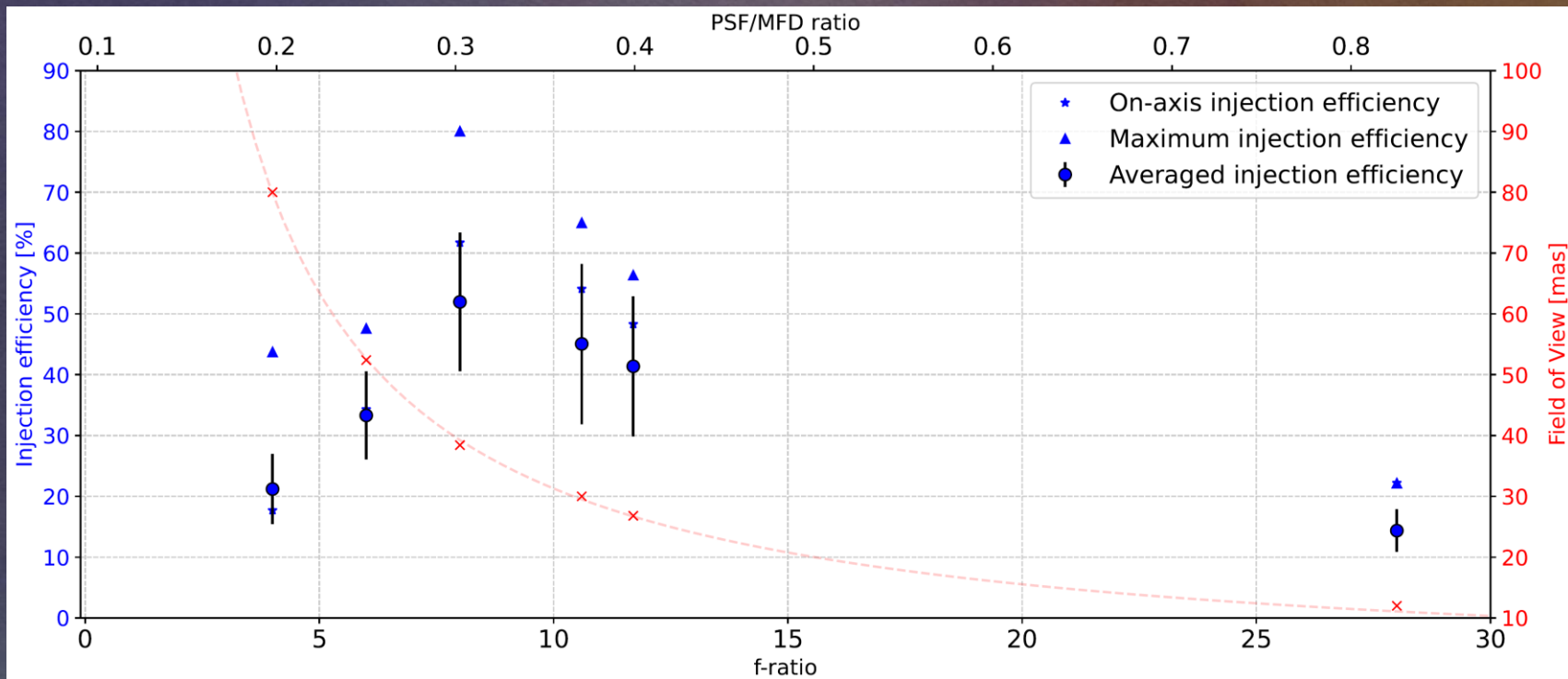
→ Less structures when focal ratio increases

❖ Spectrum reconstruction at various focal ratios

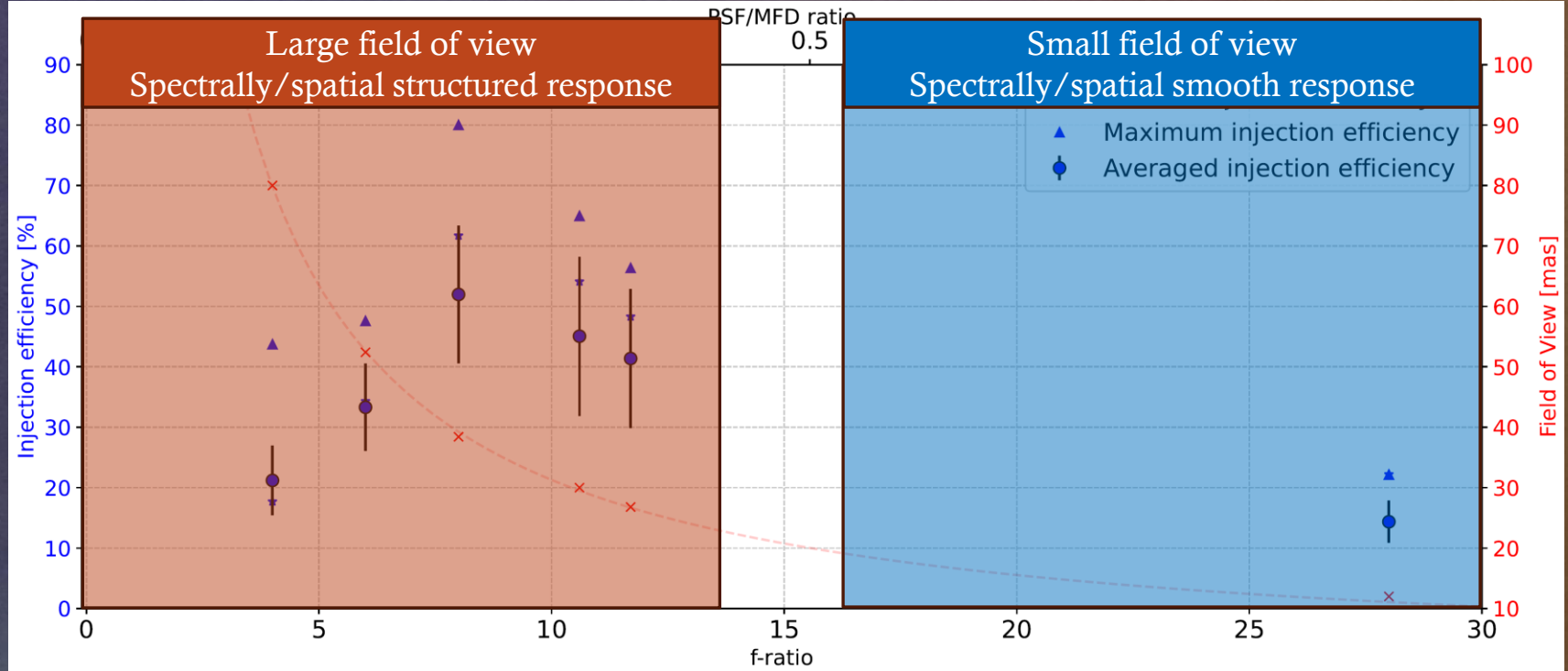
→ Spectral information loss for low focal ratios



Visible Photonic Lantern – Injection efficiency



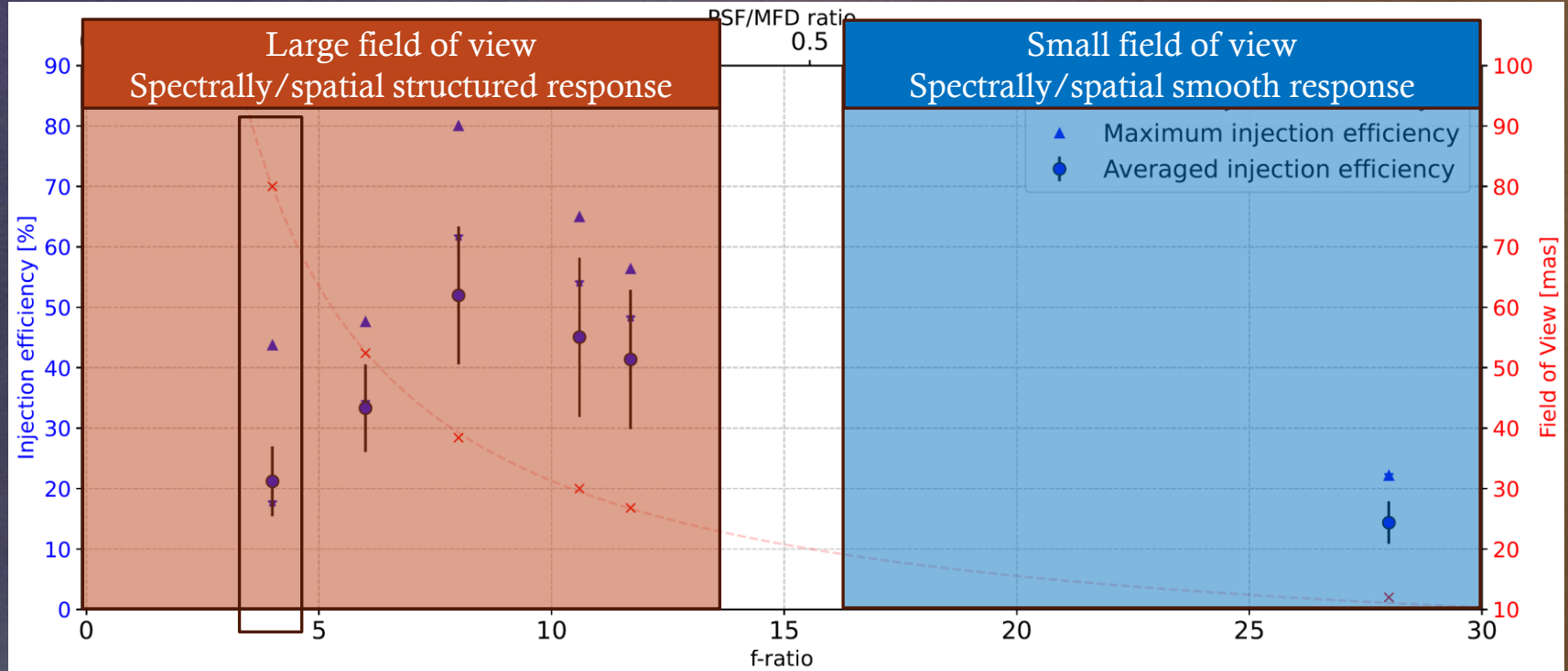
Visible Photonic Lantern – Injection efficiency



Modal overfilling

Geometric overfilling

Visible Photonic Lantern – Injection efficiency

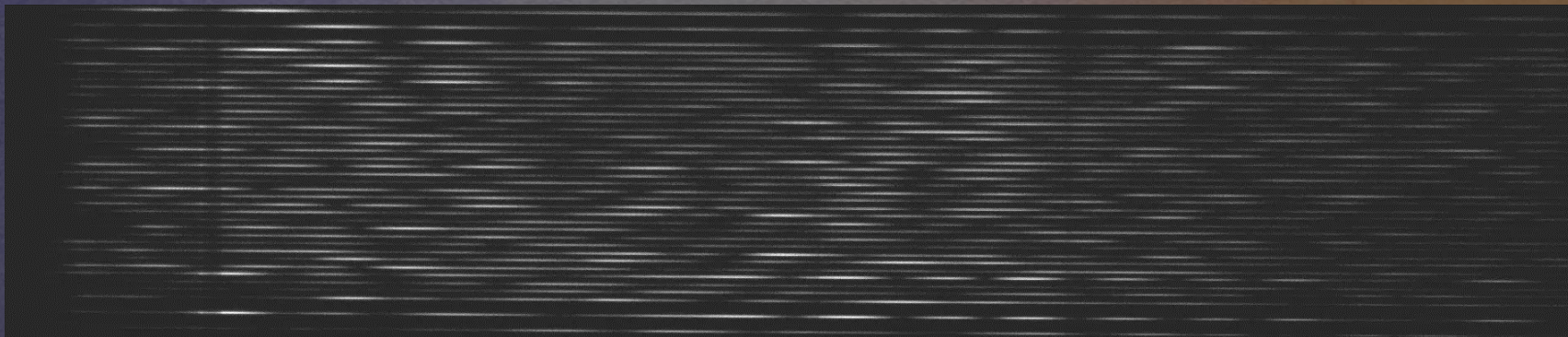


Modal overfilling

Geometric overfilling

Visible Photonic Lantern – On-sky testing on **Humu**

- Observation of Humu [*Altair*]
- Raw data – 5 ms exposure
- Medium seeing conditions (0.6 arcsec reported)

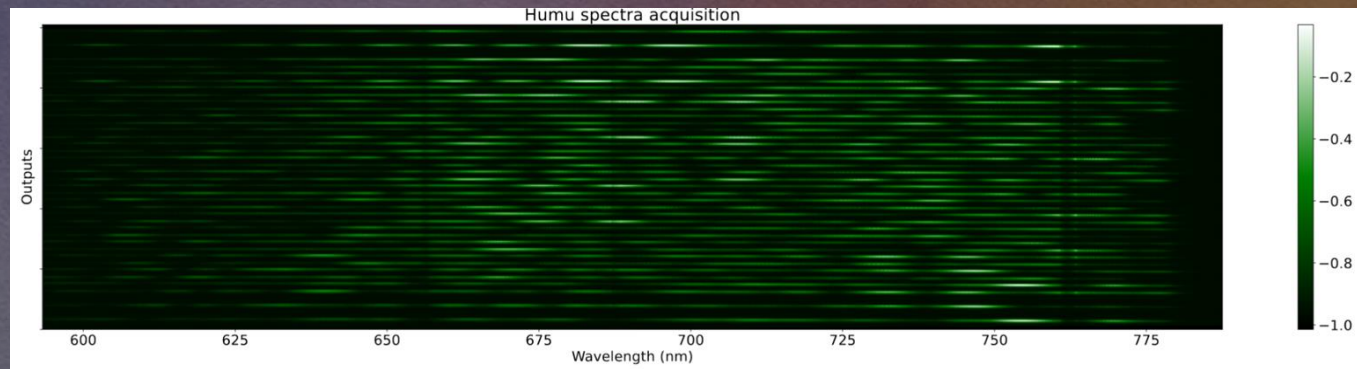


- Fast residual wavefront error ($\sim 100\text{Hz}$) redistributes light among the 38 spectra $\rightarrow$ Fast chromatic modulations

Visible Photonic Lantern – On-sky testing on Humu

Humu [*Altair*] test on-sky:

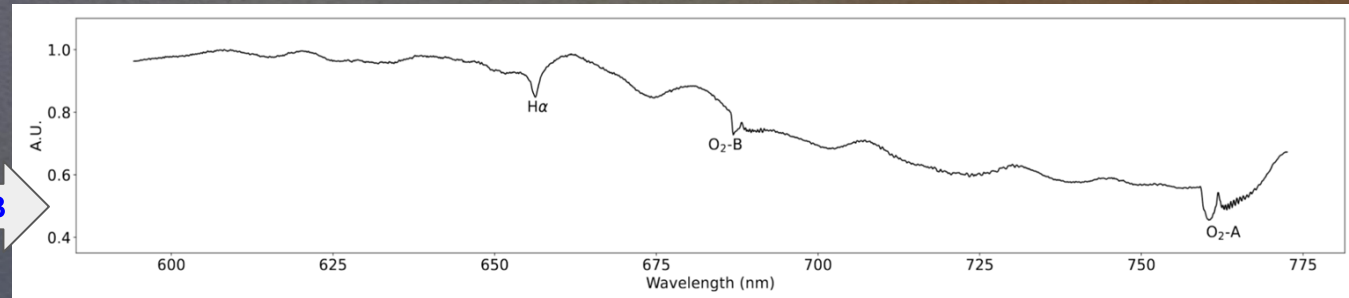
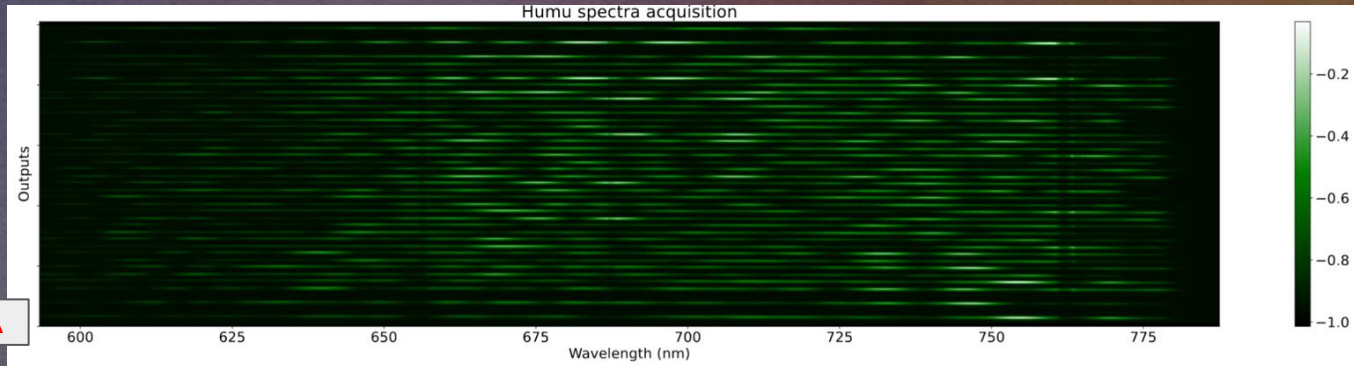
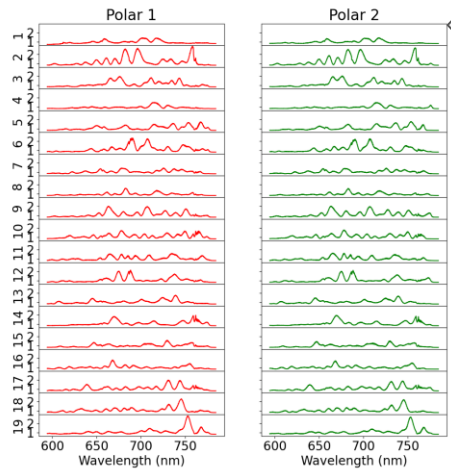
- Large FoV - f-ratio ~ 4
- Medium seeing conditions
- 5ms exposure
- 12000 frames averaged
- 1D spectrum extraction :
Extract each row (A) and co-add spectra (B)



Visible Photonic Lantern – On-sky testing on Humu

Humu [Altair] test on-sky:

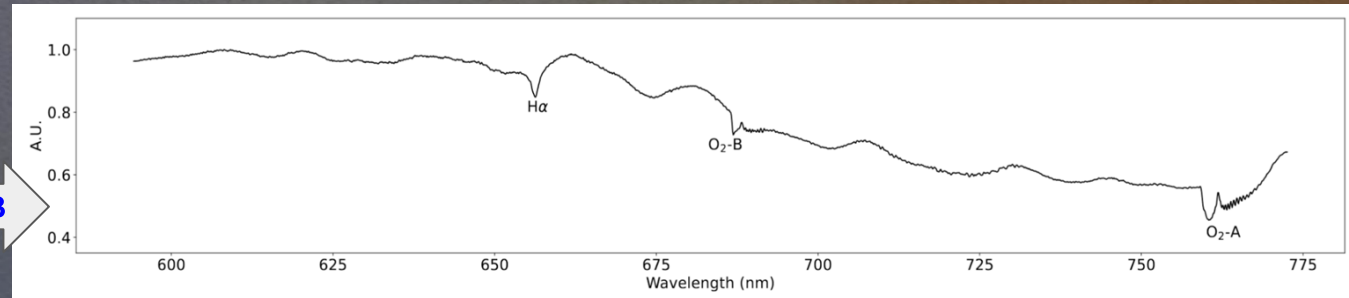
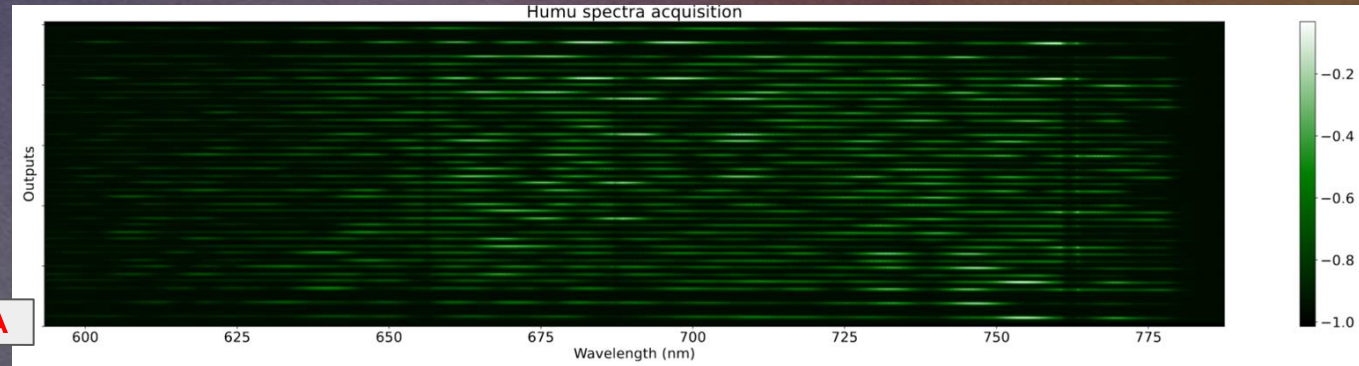
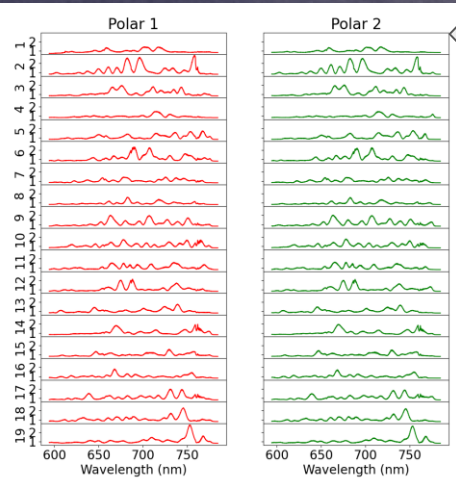
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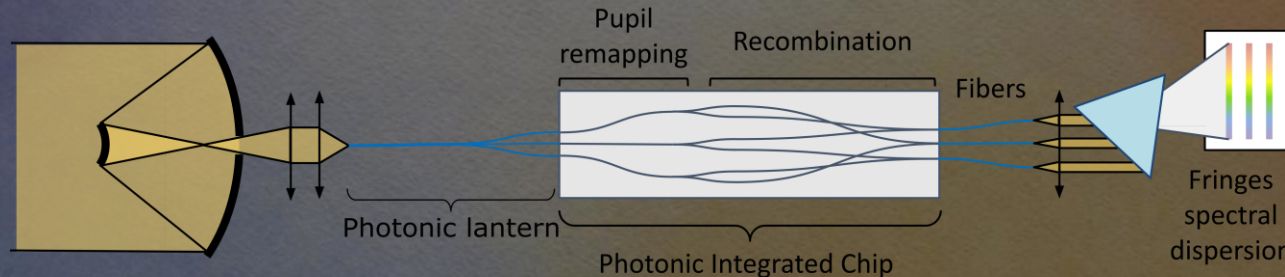
- Large FoV - f-ratio ~ 4
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Extract each row (A) and co-add spectra (B)



→ Coupling efficiency estimation : 6% average - $\sim 11\%$ at best (in lab, 18% obtained)

Conclusion & Perspectives

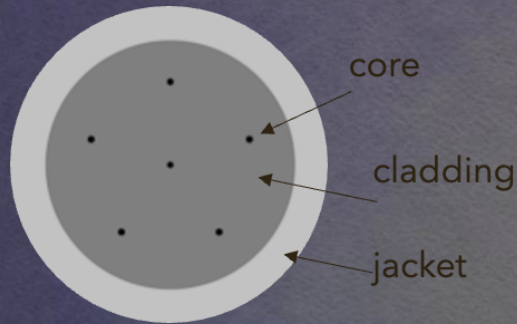
- ❖ Photonic Lanterns can enable self calibrated **high efficiency spectroscopy** with SMFs (deployable mini IFUs, radial velocity...) → various applications depending on the selected focal ratio
- ❖ More tests to come on-sky for instrument commissioning
- ❖ **Key perspectives for the future :**
 - ❖ Coupling with high resolution spectrograph (R-60,000) for H-alpha line characterization
 - ❖ Enabling Extreme Precision Radial Velocity measurements
 - ❖ Combination with photonic chips



- ❖ Waverfont sensing (petalling, cophasing, NCPAs, Self calibration)

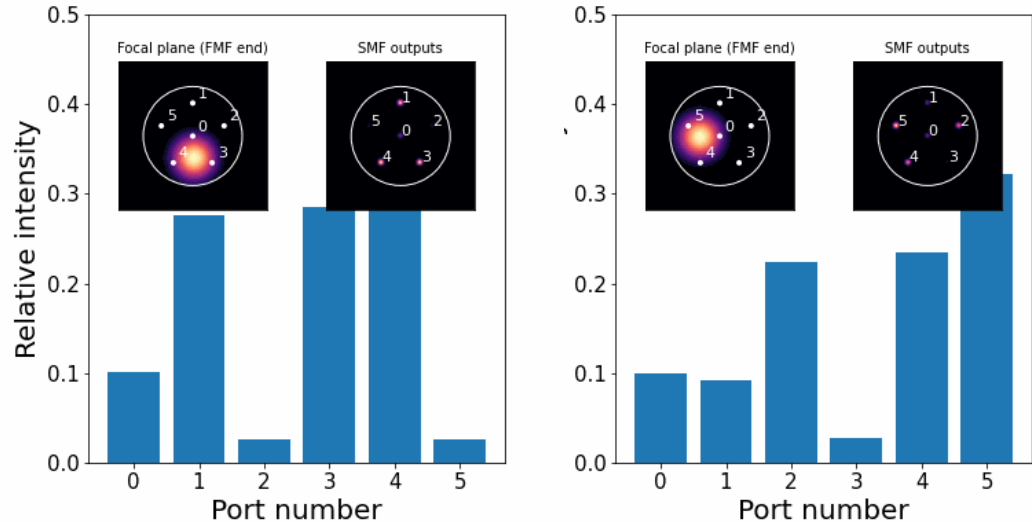
Properties of Photonic Lanterns

The intensity distribution at the SMF outputs depends on the **input scene**



Photonic Lantern with
6 SMF outputs

Numerically simulated intensity responses for a standard 6-port PL

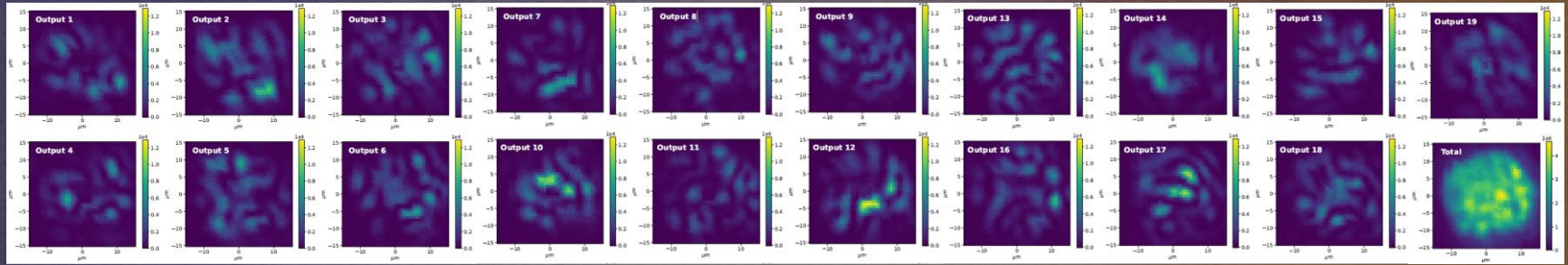


Simulations by Yoo-Jung Kim

→ PL outputs can be used for **constraining the input scene for science applications.**

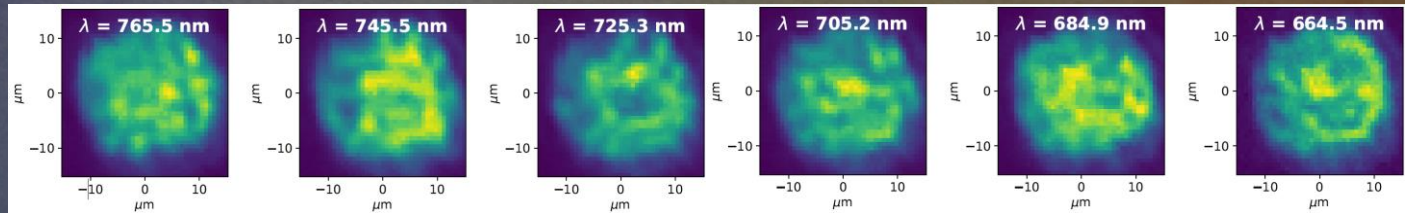
Visible Photonic Lantern – Coupling maps

❖ Coupling maps for 19 outputs + Total @ 765nm



→ **High order structures** on each output maps - smoother on the total (ratio darkest/brightest ~ 0.4)

❖ Coupling maps at various wavelengths

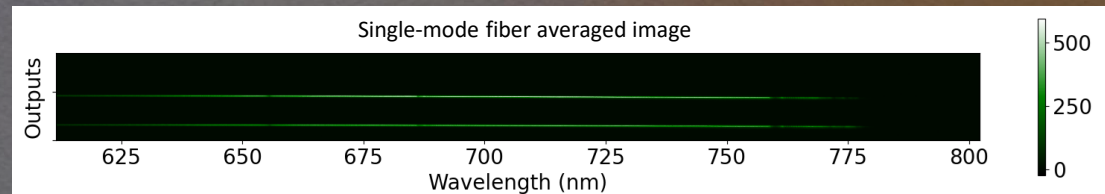
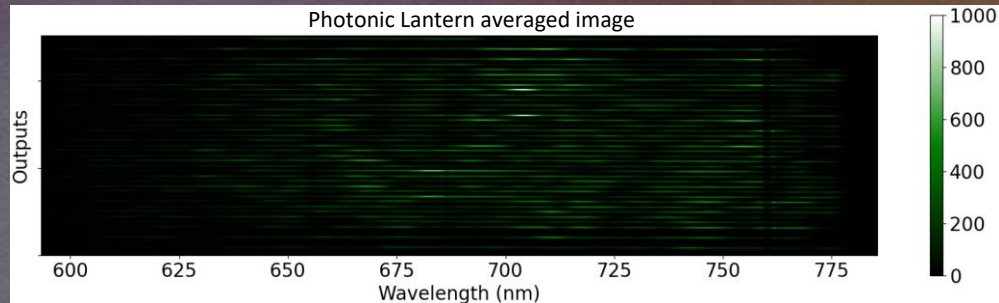


→ **Pattern changes with the wavelength**

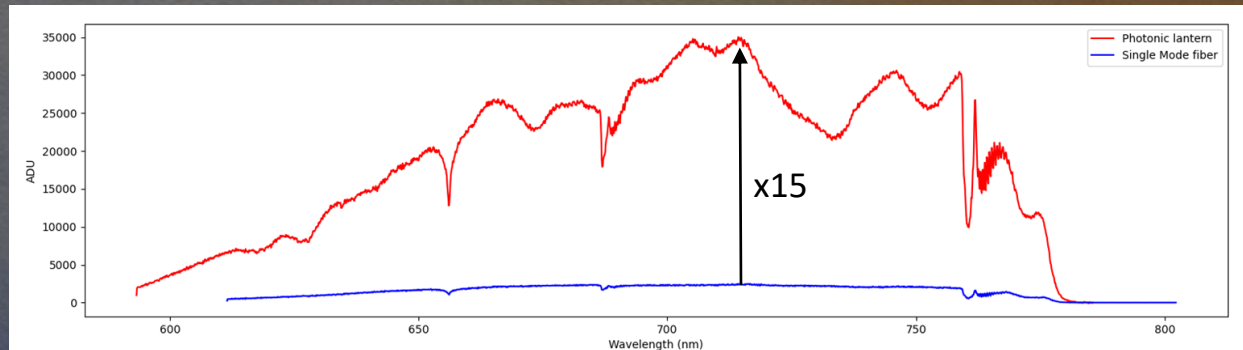
Visible Photonic Lantern – On-sky testing on Pō‘ā

Pō‘ā [Algol] :

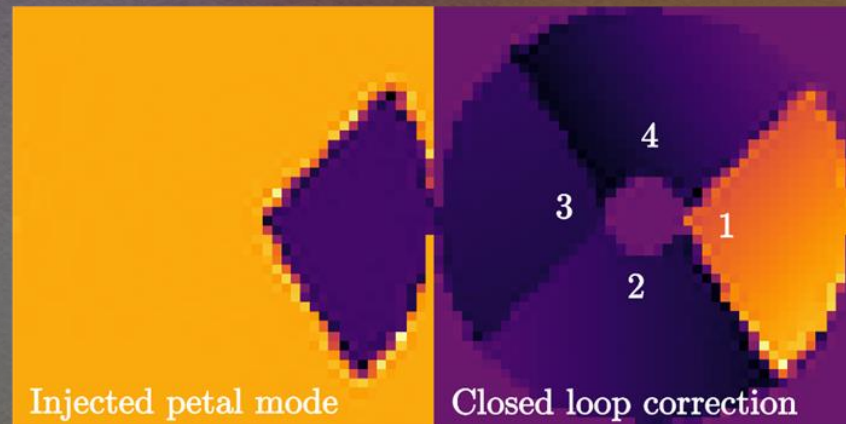
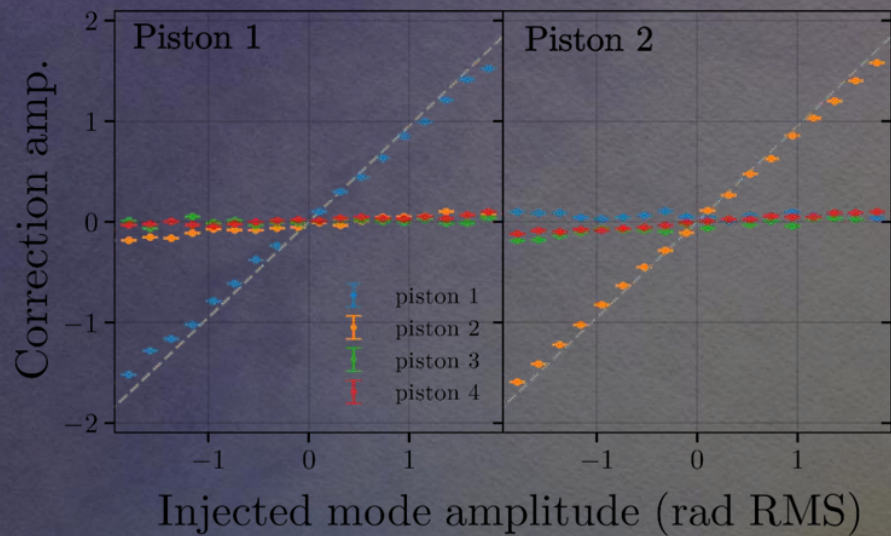
- 20ms exposure
- 3000 frames averaged
- 1D spectrum extraction :
Extract each row and co-add spectra



- Comparison with SMF spectroscopy (same exposure time, spectrum extracted from 3000 averaged frames)
- At best, gain x15
- Overall, x12.5 more flux

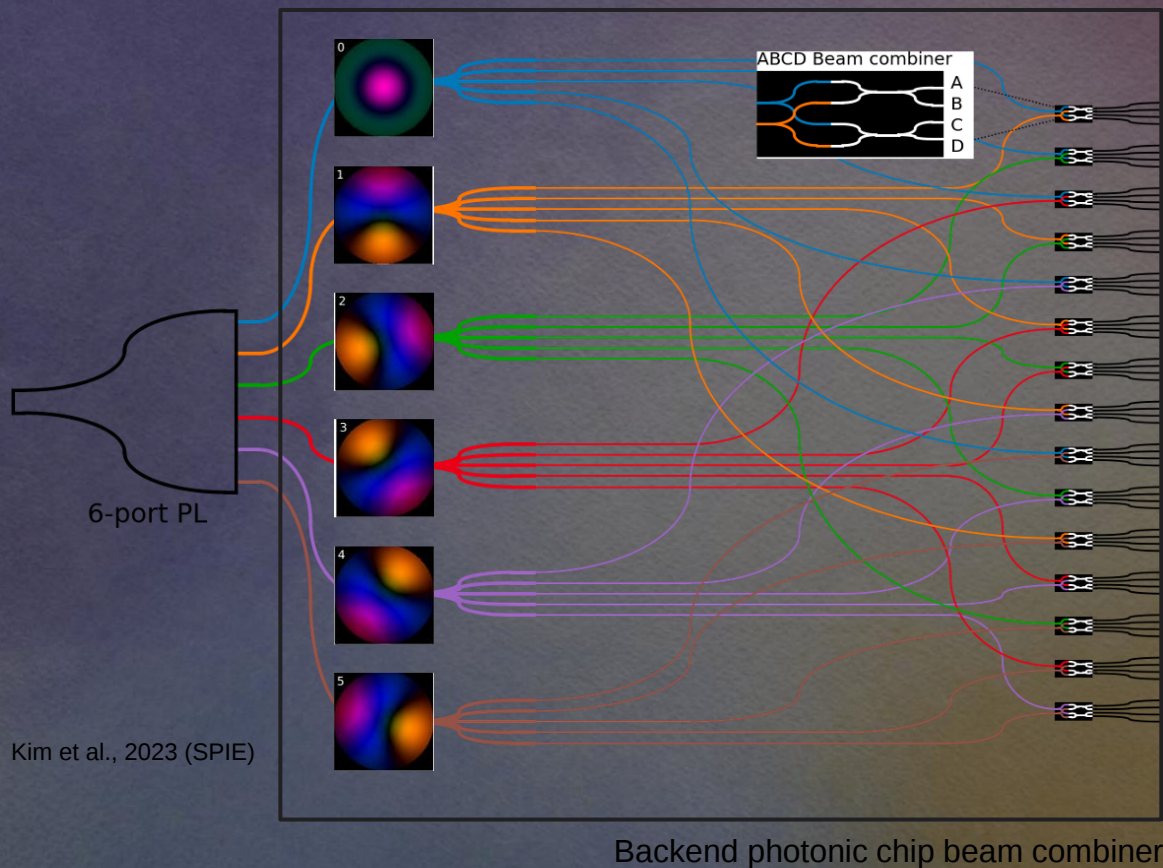


Correction of static aberrations: petalling / low wind effect – Work by Jon Lin



Correction of the **four petal-piston modes** out to ~ 2 radians RMS

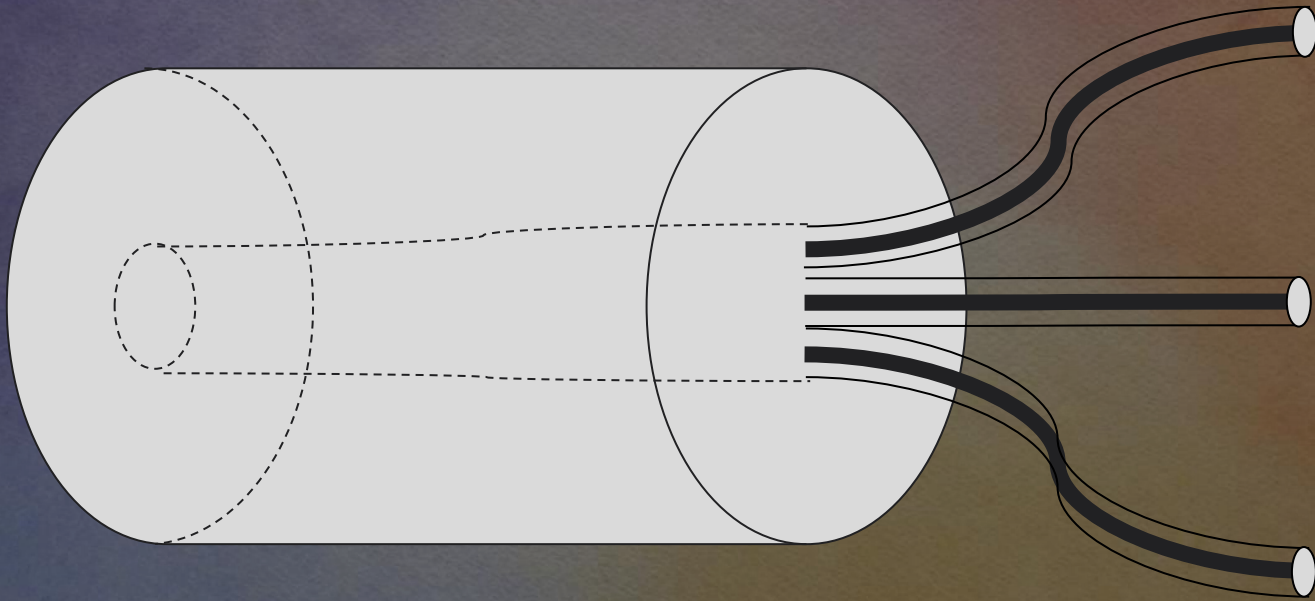
Interferometric imaging with Photonic Lanterns

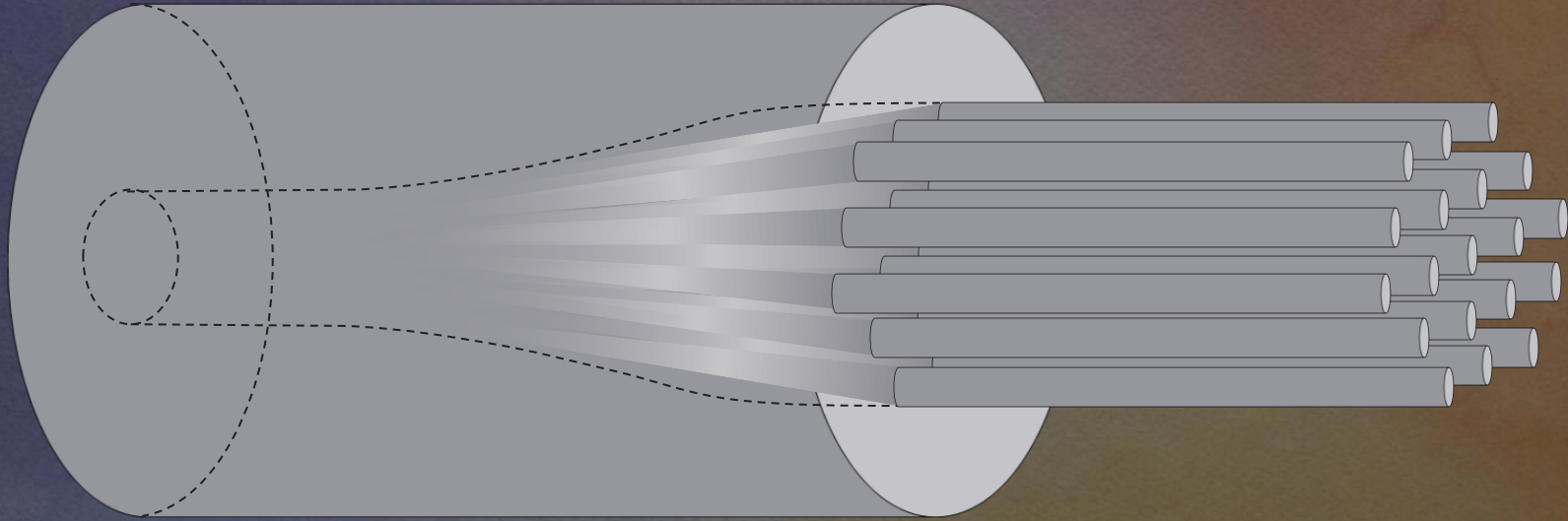


Kim et al., 2023 (SPIE)

A **pairwise beam combination** of PL outputs (e.g., ABCD beam combiner) to measure coherence properties in the incoming wavefront filtered by different apertures (PLPMs). (Single telescope interferometry)

$$\mathcal{V}_{ij} = \frac{(I_{ij,A} - I_{ij,B}) + (I_{ij,C} - I_{ij,D})i}{I_{ij,A} + I_{ij,B} + I_{ij,C} + I_{ij,D}}$$





Multi-mode
input

Adiabatic taper

Multi-SMFs
output

