omography Adaptive optics Research experimenT **: Subaru** 1

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ABSTRACT

We present the current status of the ULTIMATE-START (Subaru Tomography Adaptive optics Research experimenT) project, an on-going laser tomography adaptive optics (LTAO) project on the Subaru telescope. The main purpose of this project is to realize high Strehl Ratio AO correction not only in NIR bands, but also in visible bands above 600nm. The LTAO system will be operated by four 32x32 Shack-Hartmann wavefront sensors (SH-WFSs) and four LGSs. The LTAO WFSs will be installed behind AO188, which is the current facility AO system on the Nasmyth platform of the Subaru telescope. We will use the low-order WFS and DM of AO188 for Tip-Tilt measurements with a NGS and wavefrot correction, respectively. The DM of AO188 will be replaced with a 3228 element DM. Assembling of the LTAO WFS system has completed in 2022. Currently WFS data acquisition and tomographic wavefront estimatation are under testing. We also performed on-sky engineering observations of one prototype SH-WFS with a NGS and LGS with the Subaru telescope. A new laser lauching system has been installed. A single LGS is under on-sky performance verification for the open-use observations, and for LGS sysmtem, which can make an asterism with 10-40arcsec diameter, will be installed in mid-2024. The first light of the entire LTAO system is planned in late 2024. \bigcirc

OVERVIEW



ALPAO 64 x 64 DM (3228 usable actuators)

✓ SHARPEST: Turbulence profiling with Shack-Hartmann WFSs (check Ogane's talk !) · Initial results are published in Ogane et al. 2024.

Moffat FWHM vs Wave arcsec] GLAO WHM4 (0.2" @ K) Noffat LTAO (0.05" @ all λ) 1 1.5 Wavelength [um]

• In optical wavelength, AO performance by LTAO is significantly improved

SCHEDULE

SH WFS prototype test at Subaru: 2022 4 LGS launching system: 2023 AO188 DM upgrade: 2024

2024

2024 202X

LTAO WFS installation on Subaru:

NIR imaging with LTAO (IRCS): Optical IFU with LTAO (3DII):

compared to the AO188 SCAO model. LTAO will achieve the close to

diffraction limit of the Subaru telescope in the optical and NIR.

Expected FWHM versus wavelength of current AO system (SCAO) and the LTAO

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