

Dynamical Properties of Massive Quiescent Galaxies at $z \sim 4$

Masayuki Tanaka (NAOJ)

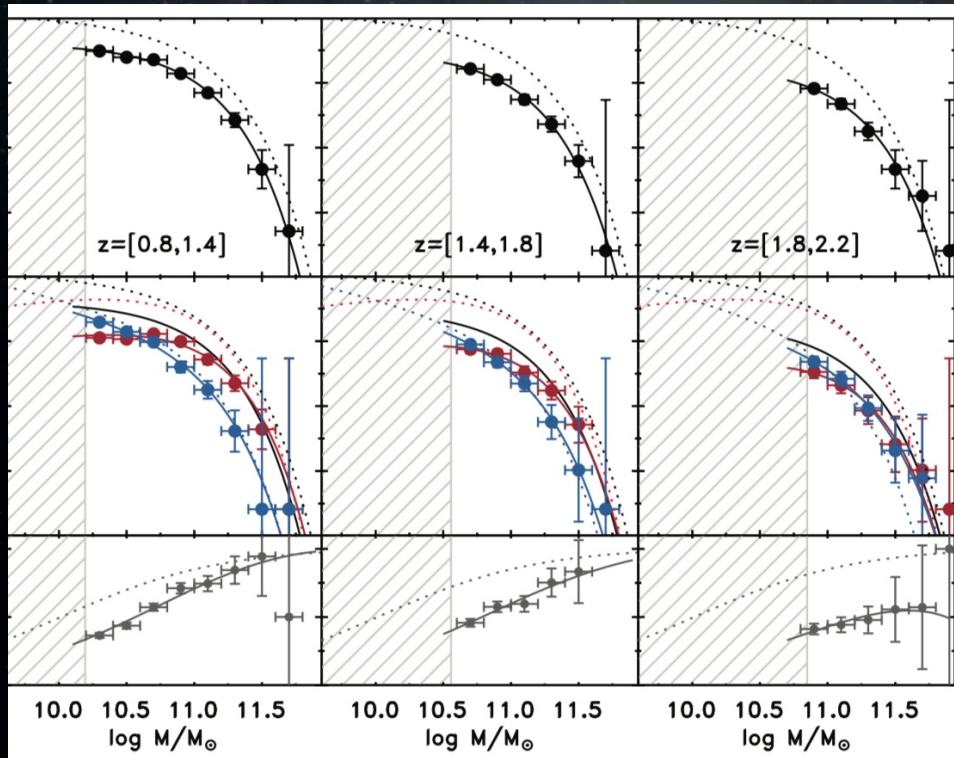
Masato Onodera, Rhythm Shimakawa, Po-Feng Wu, et al.

A deep-field astronomical image showing a vast field of galaxies and stars against a black background. The image is filled with numerous galaxies of various shapes and sizes, including spiral, elliptical, and irregular forms. Some galaxies are bright and prominent, while others are faint and distant. The stars appear as small, bright points of light, some with distinct colors like blue or red. The overall scene depicts the large-scale structure of the universe.

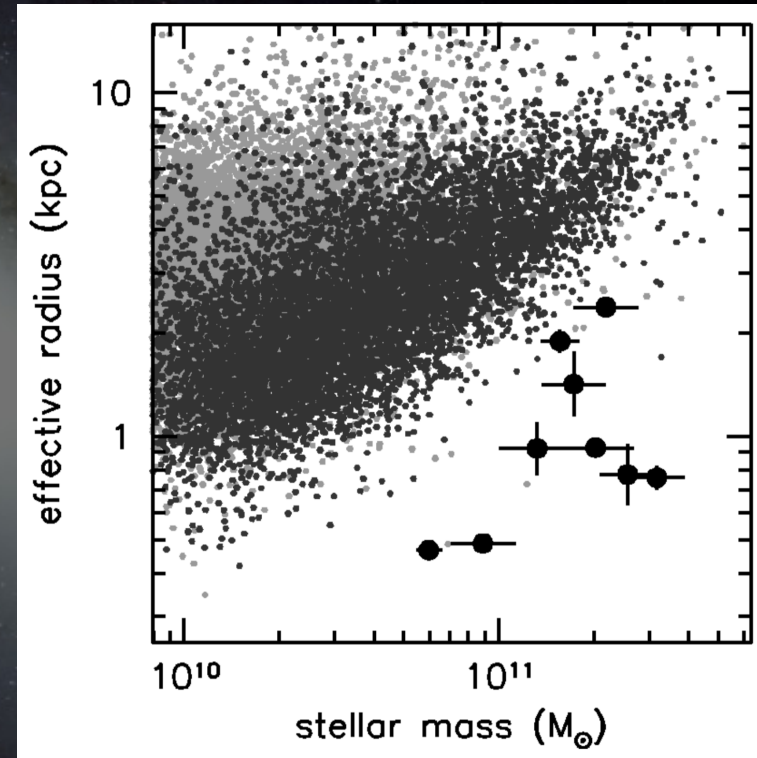
Massive early-type galaxies are dominated by old stellar populations and have not formed stars since an early epoch. What is the physical mechanism to form a massive galaxy in the early Universe and shut off their star formation?

Massive quiescent galaxies at high redshifts

Brammer+11

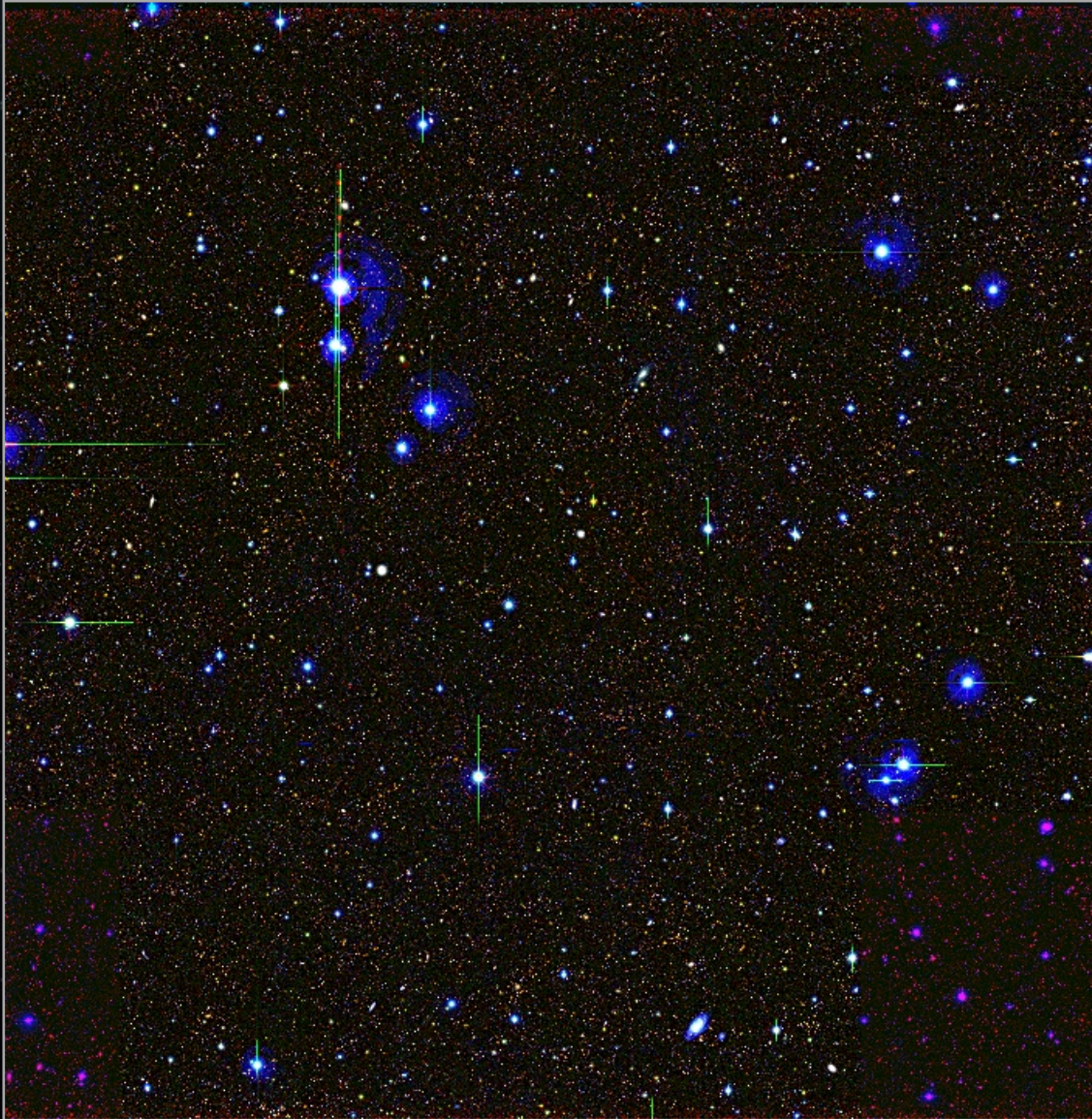


van Dokkum+09



~50% of the massive galaxies at $z \sim 2$ are compact and quiescent.

Deep Multi-Wavelength Data in UDS

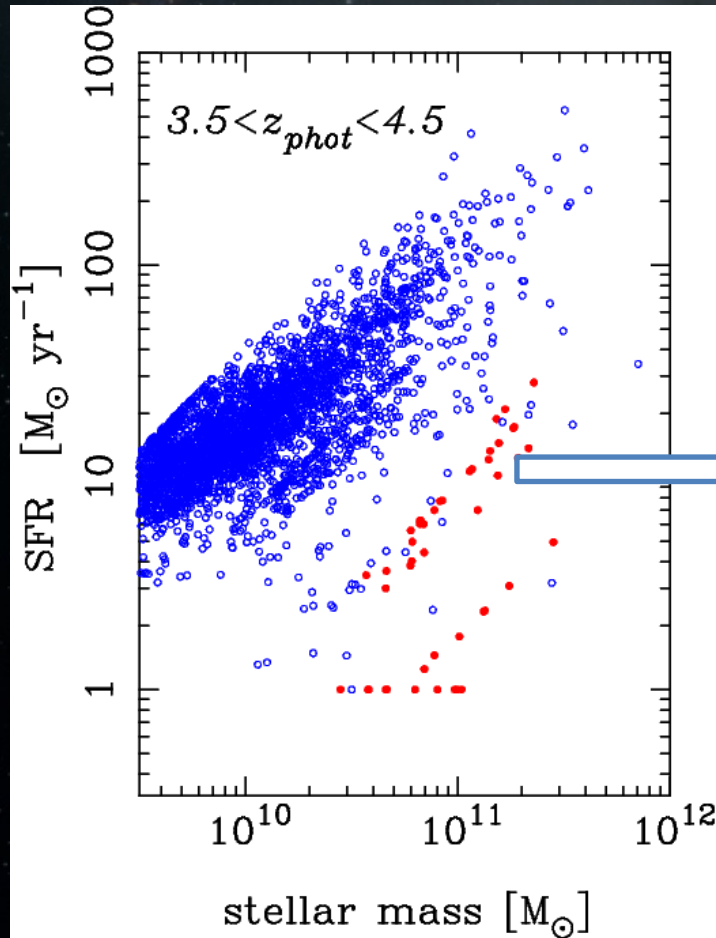


BzK composite

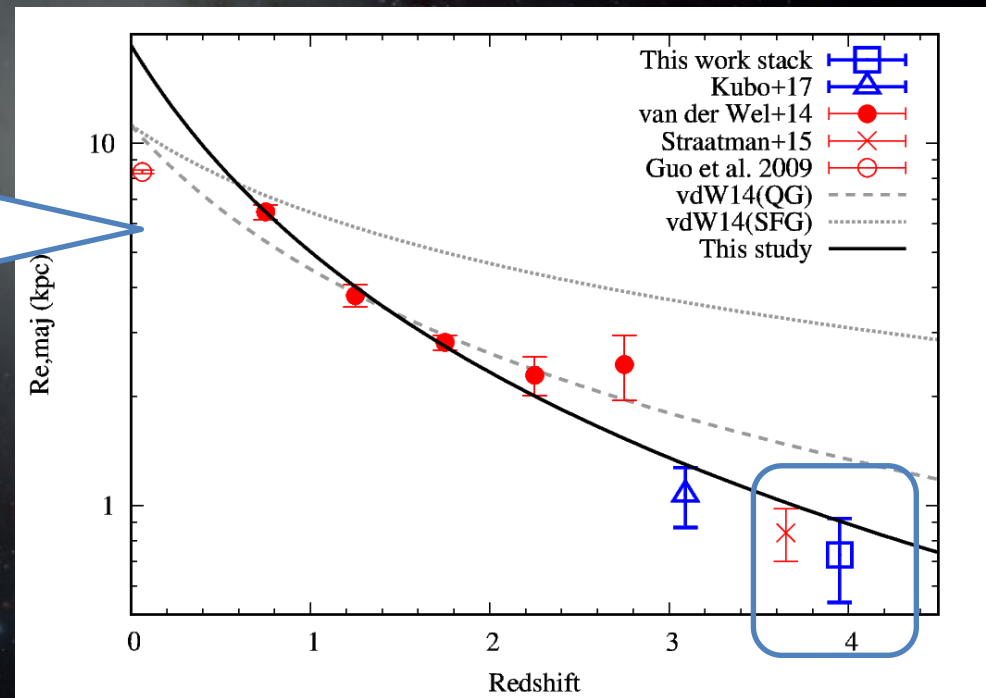
u:	26.8
B:	27.6
V:	27.3
R:	27.1
i:	27.0
z:	26.0
J:	25.2
H:	24.6
K:	25.0
[3.6]:	24.8
[4.5]:	24.3
[6.5]:	22.6
[8.0]:	22.5

5sigma in 2" aperture

Massive quiescent galaxies at $z \sim 4$



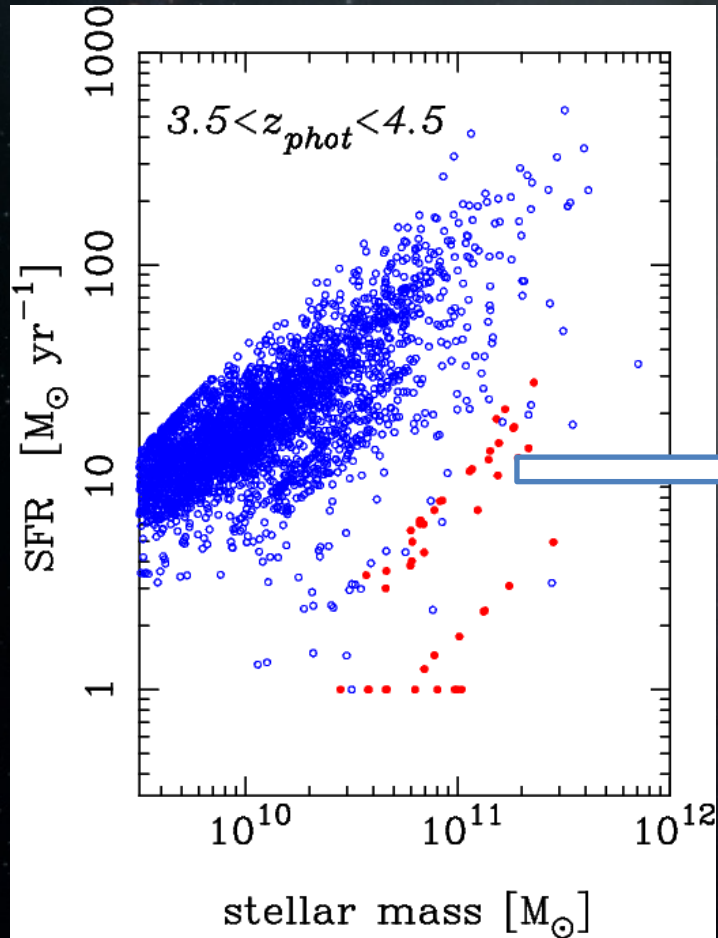
Extremely compact sizes from the AO-assisted K-band imaging!



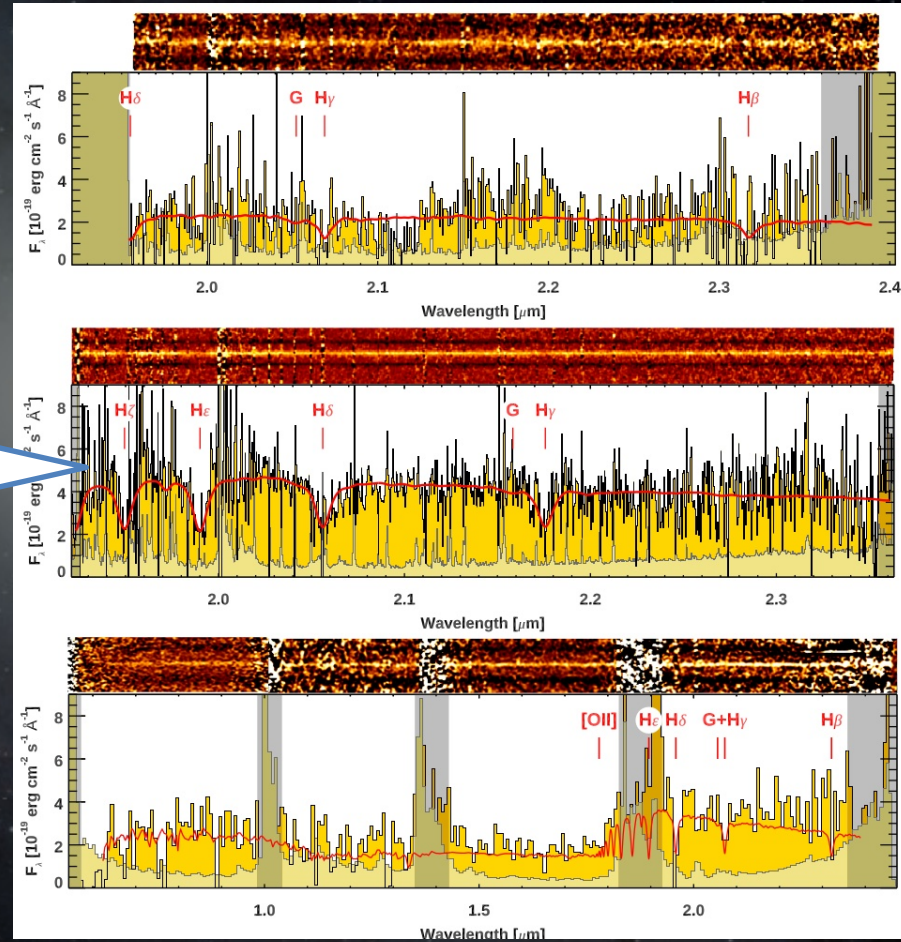
Quiescent: $s\text{SFR} < 10^{-9.5} \text{ yr}^{-1}$

Kubo, MT, et al. 2018

Massive quiescent galaxies at $z \sim 4$



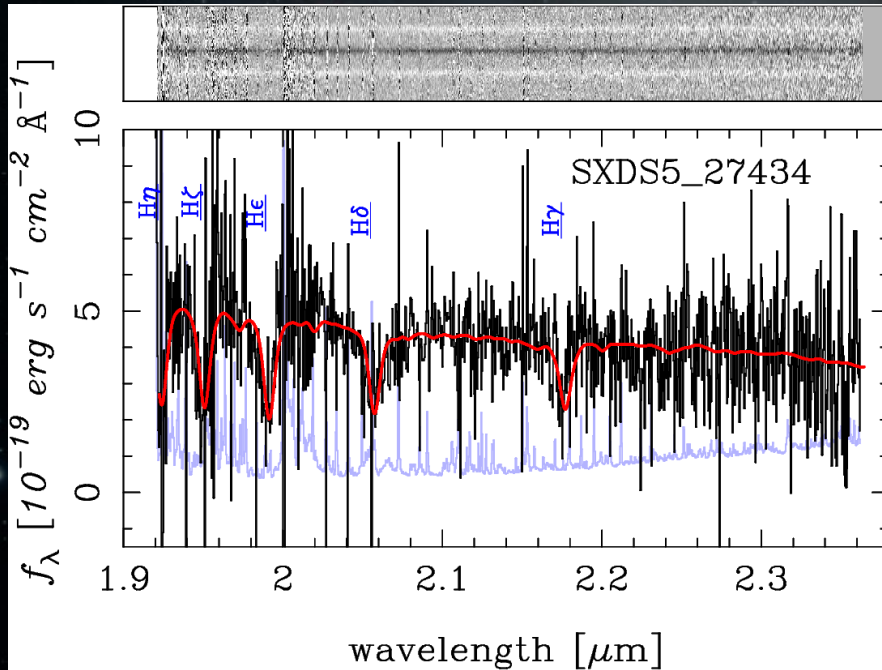
Quiescent: $\text{sSFR} < 10^{-9.5} \text{ yr}^{-1}$



Valentino, MT, et al. 2020 ApJ

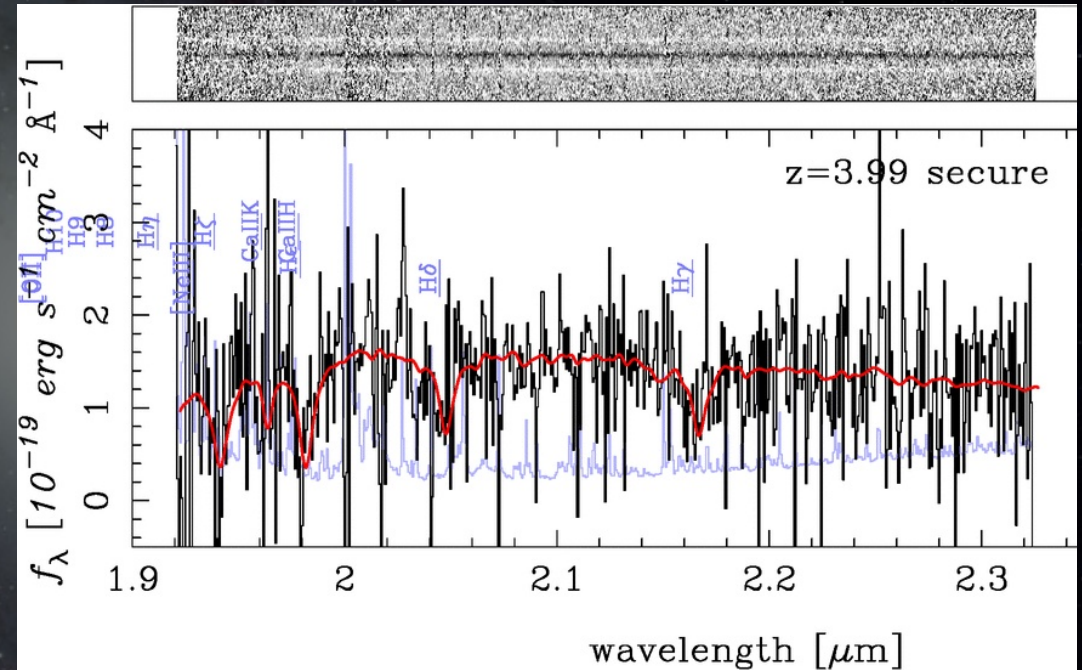
MOSFIRE spectra

$z=4.01$



~7 hours integration with MOSFIRE
(Tanaka+ 2019, ApJ, 885, L34)

$z=3.99$

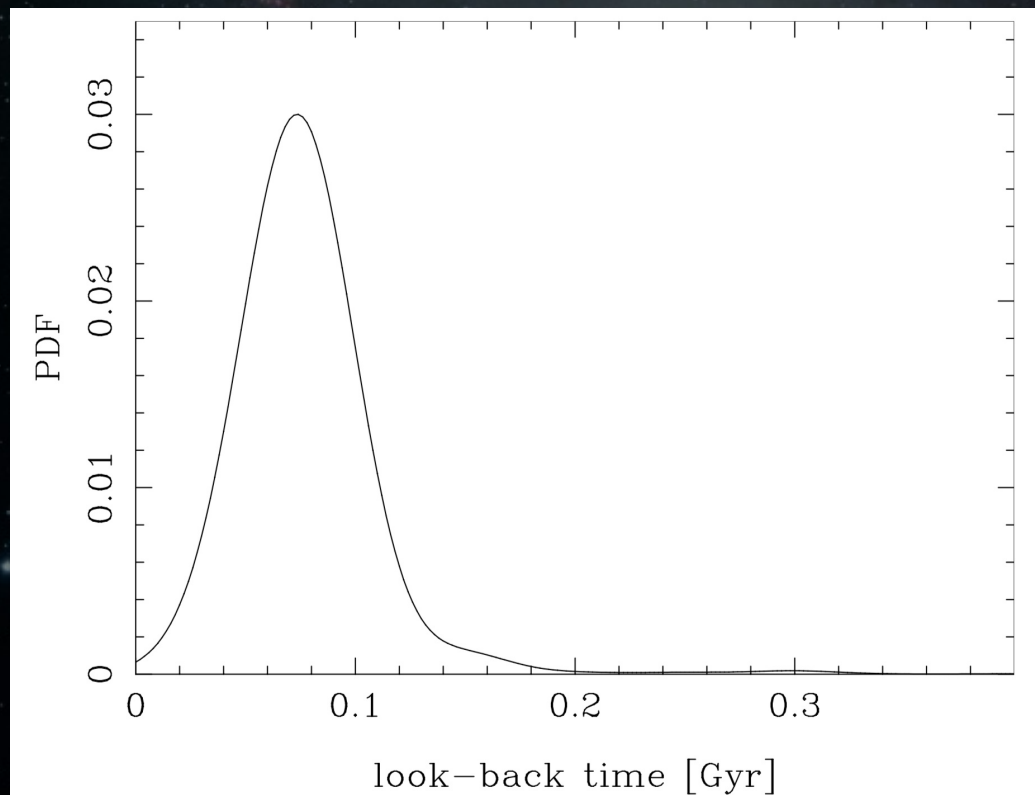


~15 hours integration with MOSFIRE
(Tanaka+ in prep)

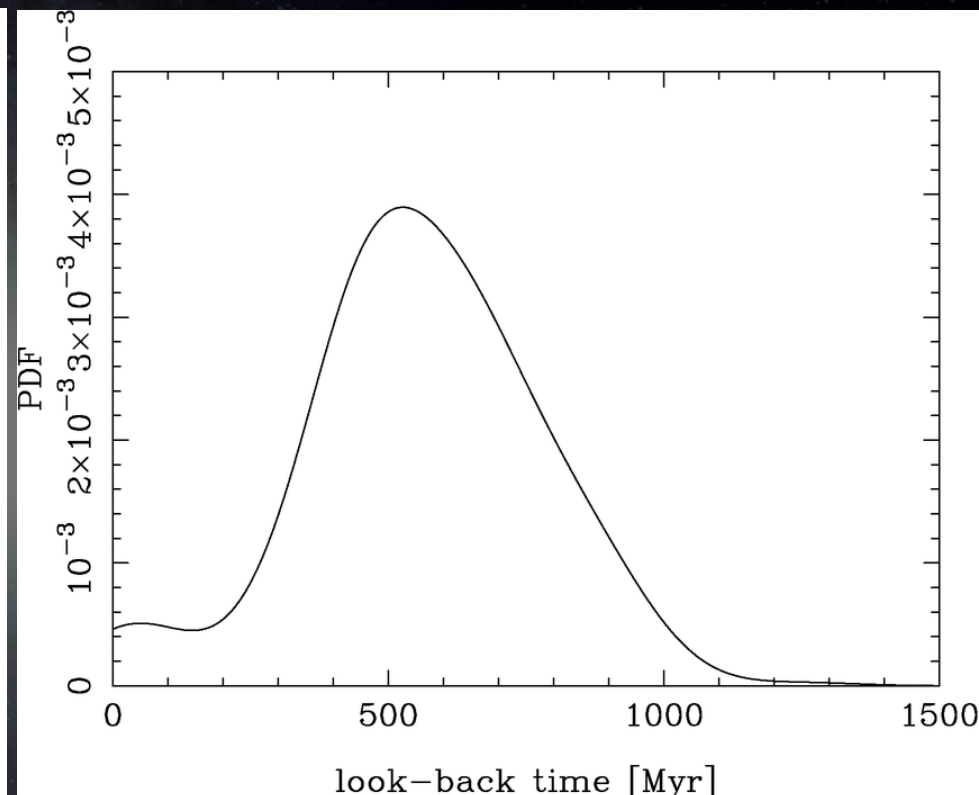
Highest redshift quiescent galaxies known to date!

Star formation histories

$z=4.01$

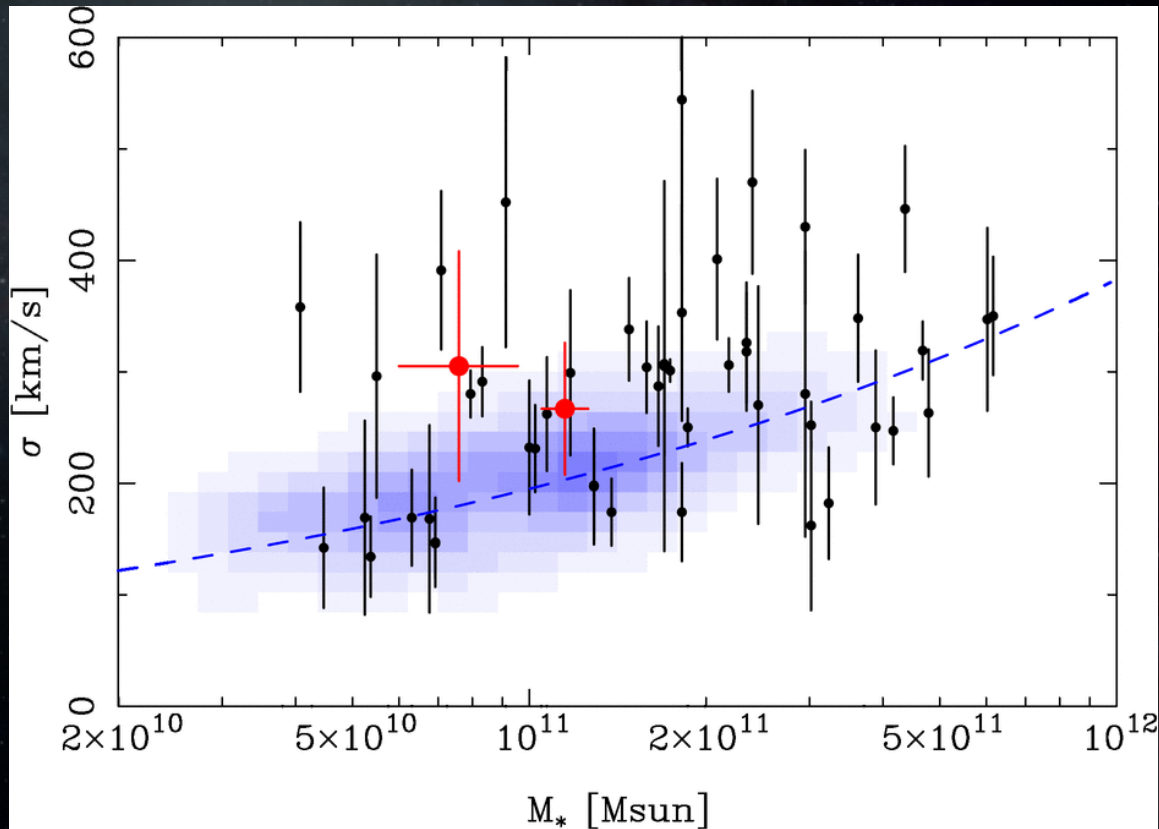


$z=3.99$



SFH from pPXF (age distribution of the SSP templates contributed to the best-fit spectrum). The $z=4.01$ is a young poststarburst, while the $z=3.99$ is a more typical poststarburst. The latter object has $z_{\text{form}} \sim 6$.

Massive quiescent galaxies at $z=4$



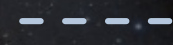
$z=4$



$z \sim 2$



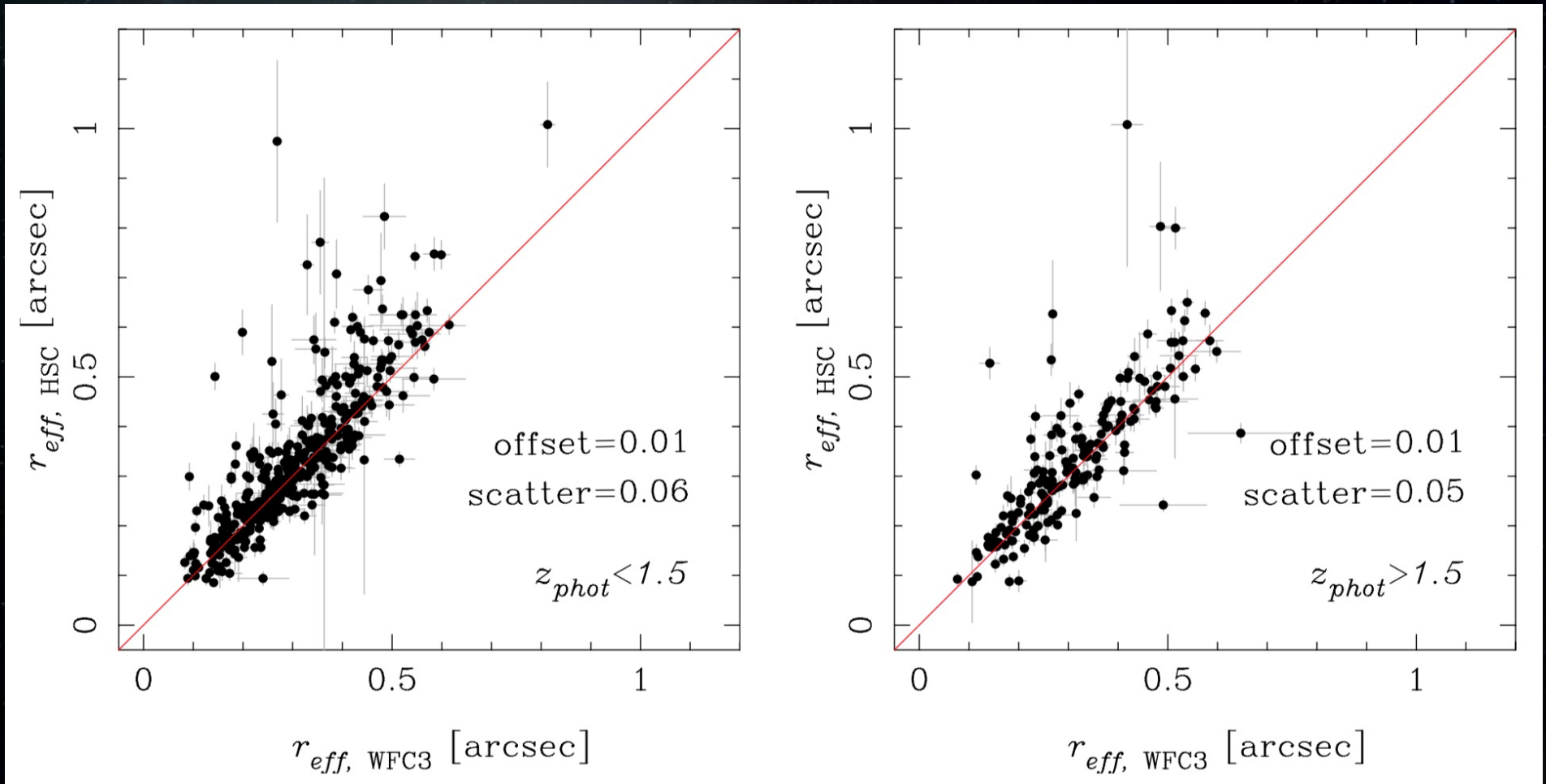
$z \sim 0.7$ (LEGA-C)



$z=0$

Stellar velocity dispersion does not significantly evolve over the last 12 Gyr

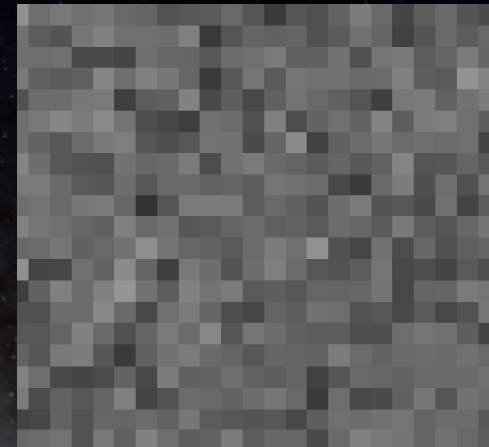
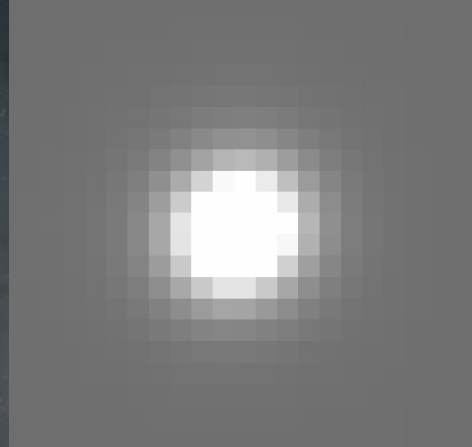
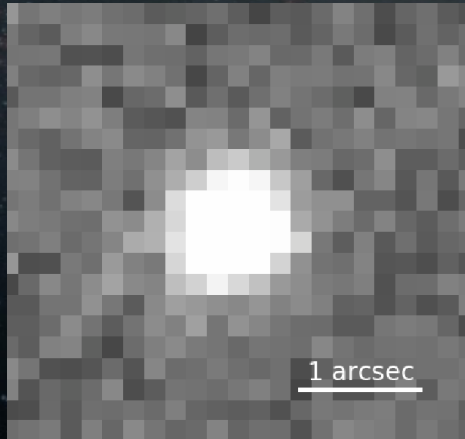
Rest-frame UV sizes from HSC-SSP



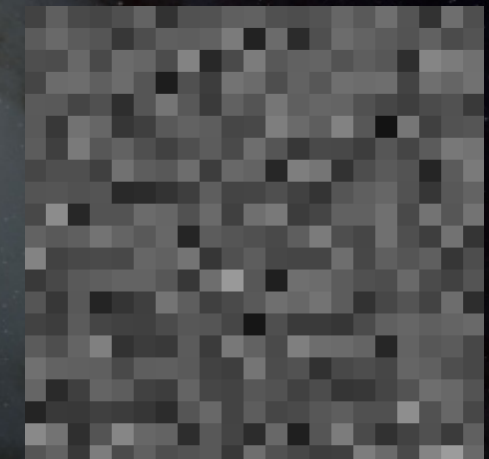
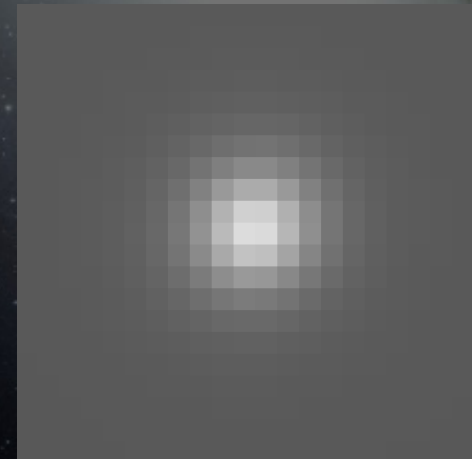
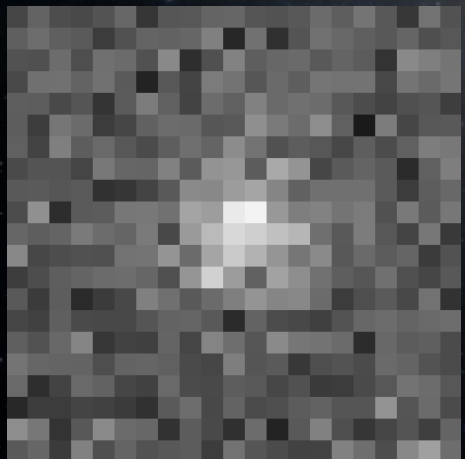
HSC-SSP is pretty good for constraining sizes.

Rest-frame UV sizes from HSC-SSP

$z=4.01$



$z=3.99$



Original image

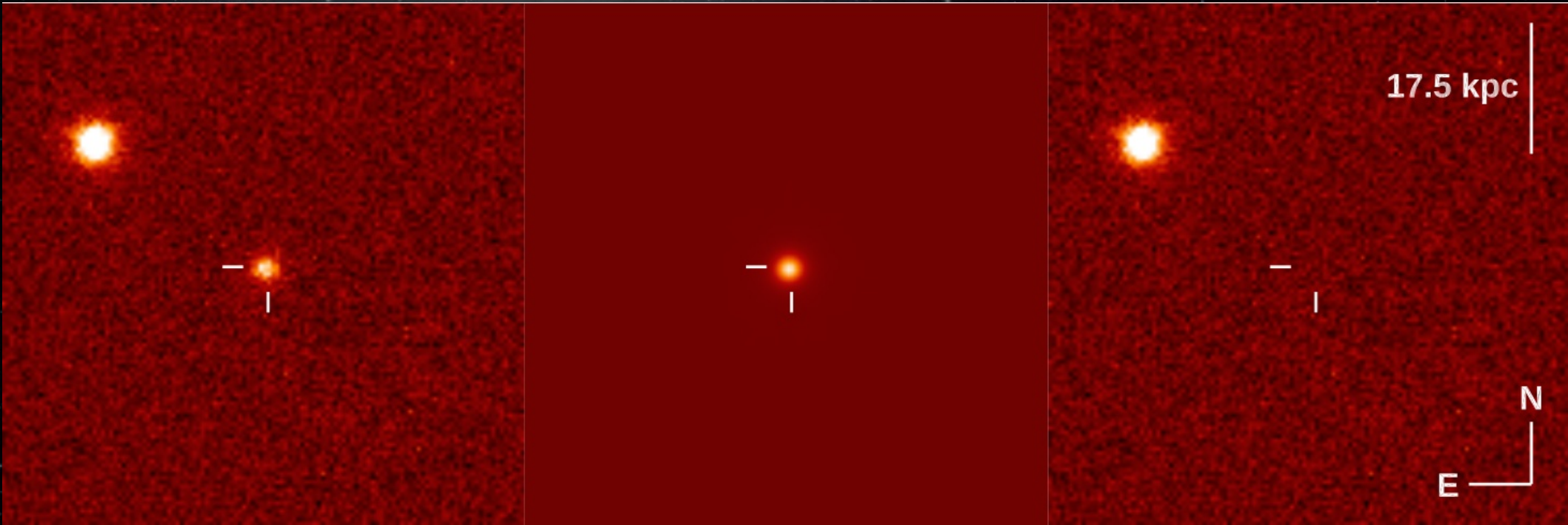
Best-fit model

Residual

$$z=4.01: r_{\text{eff}} = 0.76 \pm 0.20 \text{ kpc}$$

$$z=3.99: r_{\text{eff}} = 1.0 \pm 0.50 \text{ kpc}$$

Rest-frame optical sizes from HAWK-I + GLAAL



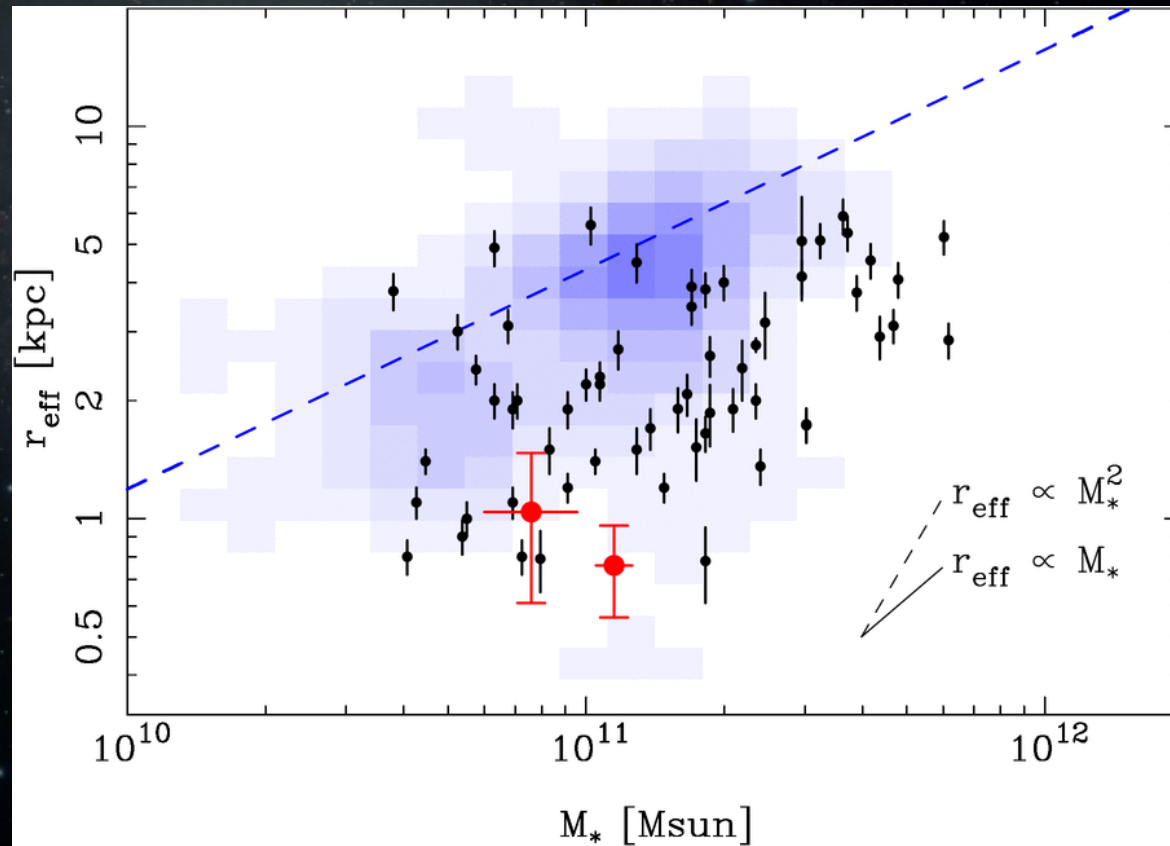
Original image

Best-fit model

Residual

The $z=4.01$ galaxy has $r_{\text{eff}} = 0.74 + 0.17 / -0.08$ kpc

Size-Stellar Mass relation



● $z=4$

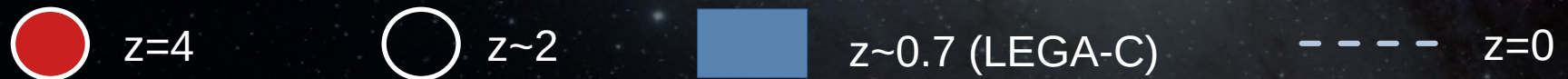
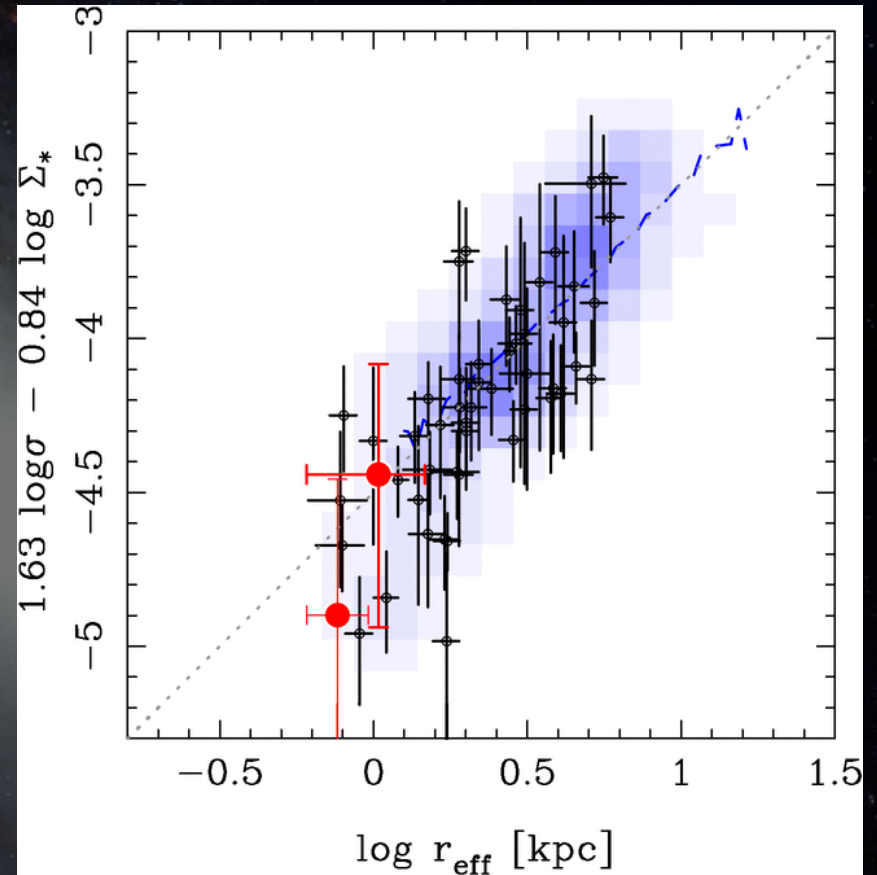
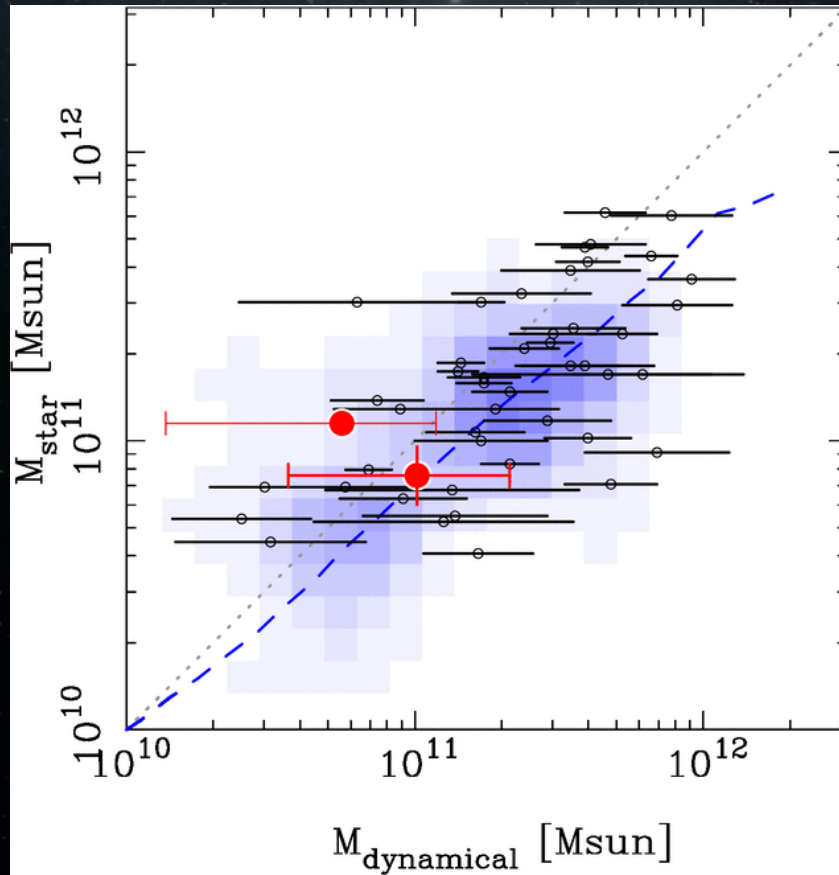
○ $z \sim 2$

■ $z \sim 0.7$ (LEGA-C)

--- $z=0$

The $z=4$ objects are very compact for their mass. Their sizes are Consistent with the typical size of $z \sim 4$ galaxies from Kubo+ 2018.

First Dynamical Analysis at $z=4$



- The dynamical mass is consistent with the stellar mass from photometry.
- The $z=4$ object is on the mass fundamental plane.

Discussion and Summary

- We have confirmed the highest redshift quiescent galaxies at $z=4$.
- We made the first measurement of stellar velocity dispersion at such a high redshift.
- Stellar velocity dispersion does not significantly evolve over the last 12 Gyr.
- But, we know that sizes and masses increase with time.
- This suggests that the mass growth does not occur equally at all radii; mass growth mostly occurs in the outer parts. That also drives the size evolution.
- This is consistent with the 2-phase formation scenario.
- But, this galaxy may be rotating!
- Step forward to a larger sample and also to JWST.