

Precise Radial Velocity Survey with IRD: Searching for Earth-like planets orbiting M dwarfs

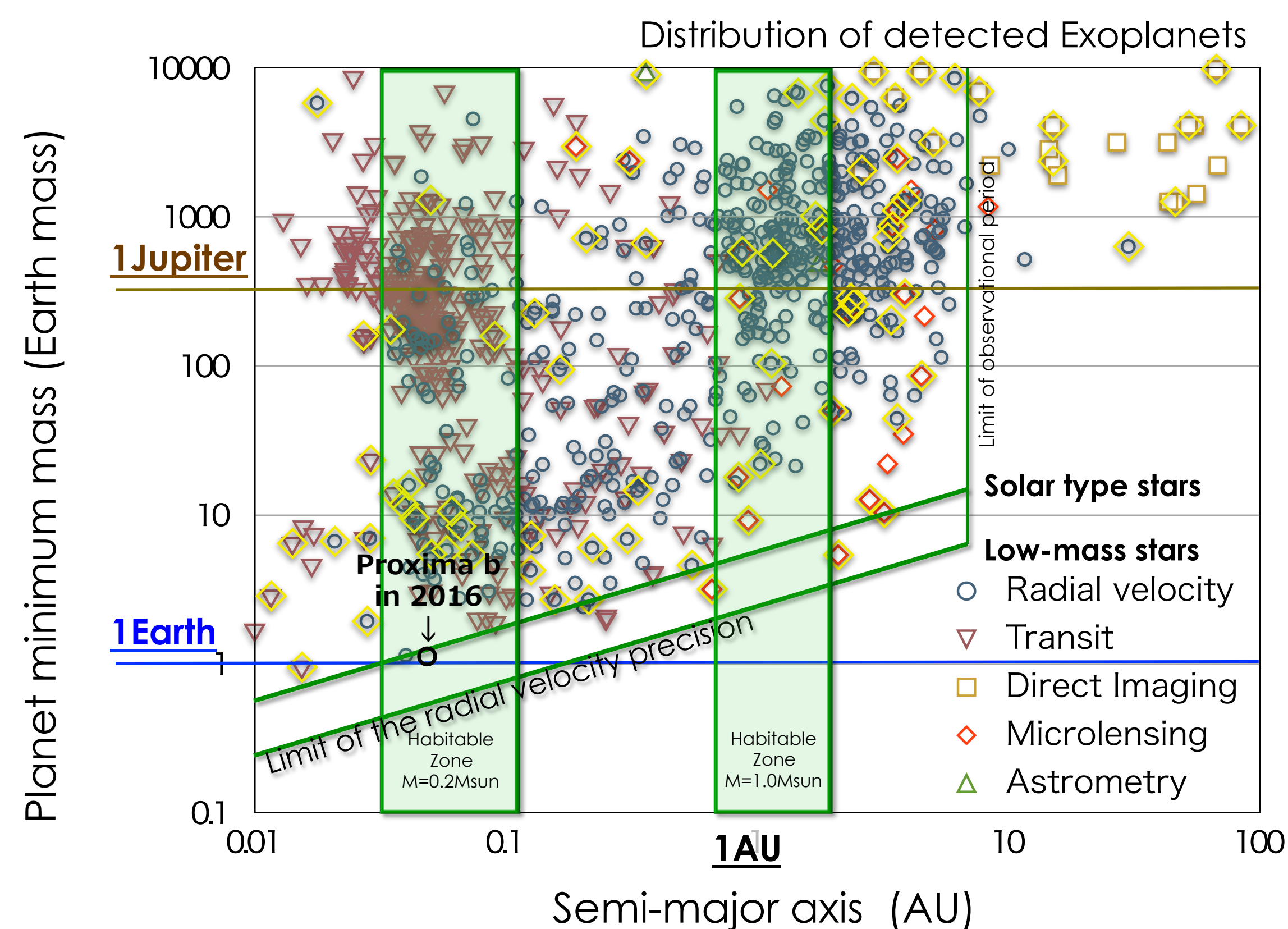
P20

Abstract

We are planning to perform a new planet search for Earth-like planets around late-M dwarf stars using **the InfraRed Doppler instrument for the Subaru telescope (IRD)**. The planet search project aims detecting Earth-mass planets in the habitable zone (HZ) and understanding configurations of planetary systems including Earth-mass planets around low-mass stars by the infrared Doppler method. Although a planetary companion with Earth-mass around Proxima Centauri in the HZ has been reported by using a most advanced optical spectrograph, HARPS, **a very strategic survey with IRD** is necessary for detecting significant number of Earth-like planets and achieving our scientific goals. In this poster, we describe a scientific background, a strategy, and simulation results of the Subaru/IRD planet search.

Introduction

Toward Earth-mass planets in HZ



- ☆ Optical planet searches around solar type stars

There are a few reports of the earth-radius planets by Kepler

However, it is difficult to detect earth-mass planets around nearby stars

- ☆ Optical planet searches around low-mass stars

There are some candidates of the planets.

However, it is difficult to search for Earth-mass planets around many stars because of small stellar flux in optical.

- ☆ Infrared planet search around lower-mass stars (Mass < 0.3 M_{Sun}):

Larger signal → Detectable for 1 M_{Earth} planets

→ Can Detect planetary systems including Earth-, Neptune-, Jupiter- mass planets

Stellar flux peak is in near infrared region.

Closer Habitable Zone (HZ) (<0.3AU).

A Strategy to detect Earth-mass planets

Difficult to detect Earth-mass planets around solar-type stars

■ Search for exoplanets using IRD/Subaru

- IRD + Subaru → Can target on faint low-mass stars
- High precision → Easier to detect Earth-mass planets
- Target stars → Can select more than 100 stars

Our targets: Late-M type dwarfs (M4V-M9V, 0.1-0.3 M_{Sun})

Large signal ($K = 0.5\sim 2$ m/s) → Easier to detect Earth-like planets

Short orbital period ($P < 40$ d) → Easier to detect Earth-like planets

Flux peak is in infrared → Effective in infrared observation

Many stars in solar neighborhood → Many bright sample

InfraRed Doppler (IRD) instrument

Telescope: **Subaru telescope** (8.2m)
Wavelength : 0.97-1.75 μm (Y, J, H -band)
Resolution : 70,000 (3pixels-sampling)
Calibrator : **laser-frequency comb**

Goal: 1 m/s radial velocity precision
for non-active low mass stars

Start obs. with IRD in 2017 spring
IRD/Subaru planet search in 2018



Kotani+14

Main Science Goals of Subaru/IRD planet search

1. Discovery of Many Earth-mass planets (in the habitable zone)
2. Statistical understanding planetary systems around low-mass stars

Subaru/IRD planet search for Earth-mass planets around low-mass stars

1. Screening observations of low-mass stars using small telescopes

To select slowly rotating stars for samples of the planet search

Small rotational velocity: <4 km/s (RV error ~ 1 m/s)

Small stellar jitter: ~1 m/s

Optical medium resolution observations

R~1500, 5500~8000Å (including H α)

At five observatories and telescopes:

Okayama(1.88m), NHAO,(2.0m)

KPNO(3.5m), Calar Alto(2.2m),

APO(3.5m)

2. Target : ~100 late-M dwarf stars

Inactive M4V-M9V type stars

with measurements of parallax

without X/UV emission

without H α emission

with masses of 0.1-0.3 M_{Sun}

They are IRD-unique targets!

Too faint for other Doppler projects using 3m-class telescopes

Tess would detect many planets around late-M stars with J>12 (too faint)

3. Strategic observation using the Subaru 8.2m telescope and IRD

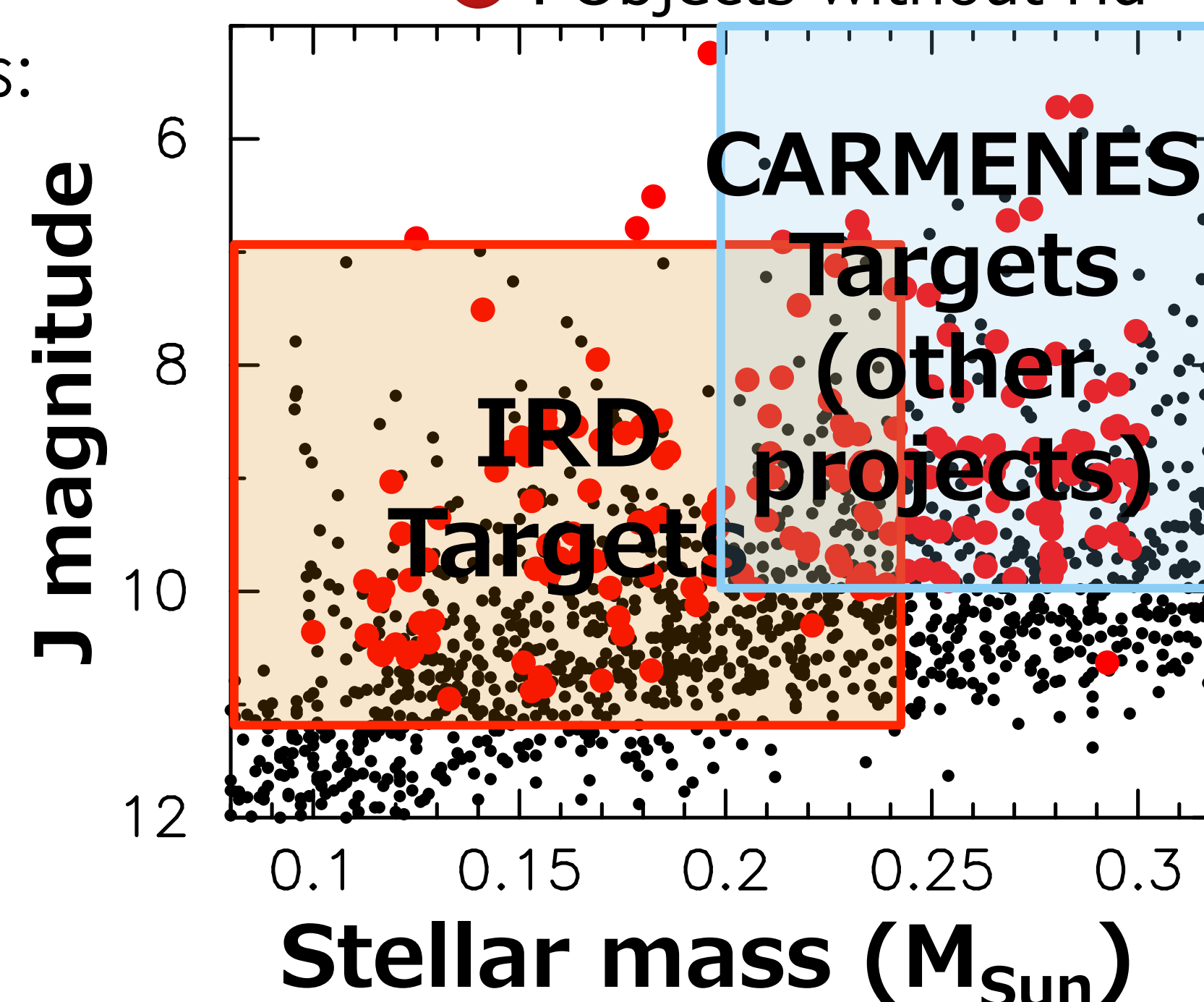
half night-allocated mode: ex. ~70 half nights/yr 8(-6h/1night)

Period of the survey : 5 yr (Success rates of 70%)

Number of observations : Nobs~ 50 or 80 / star

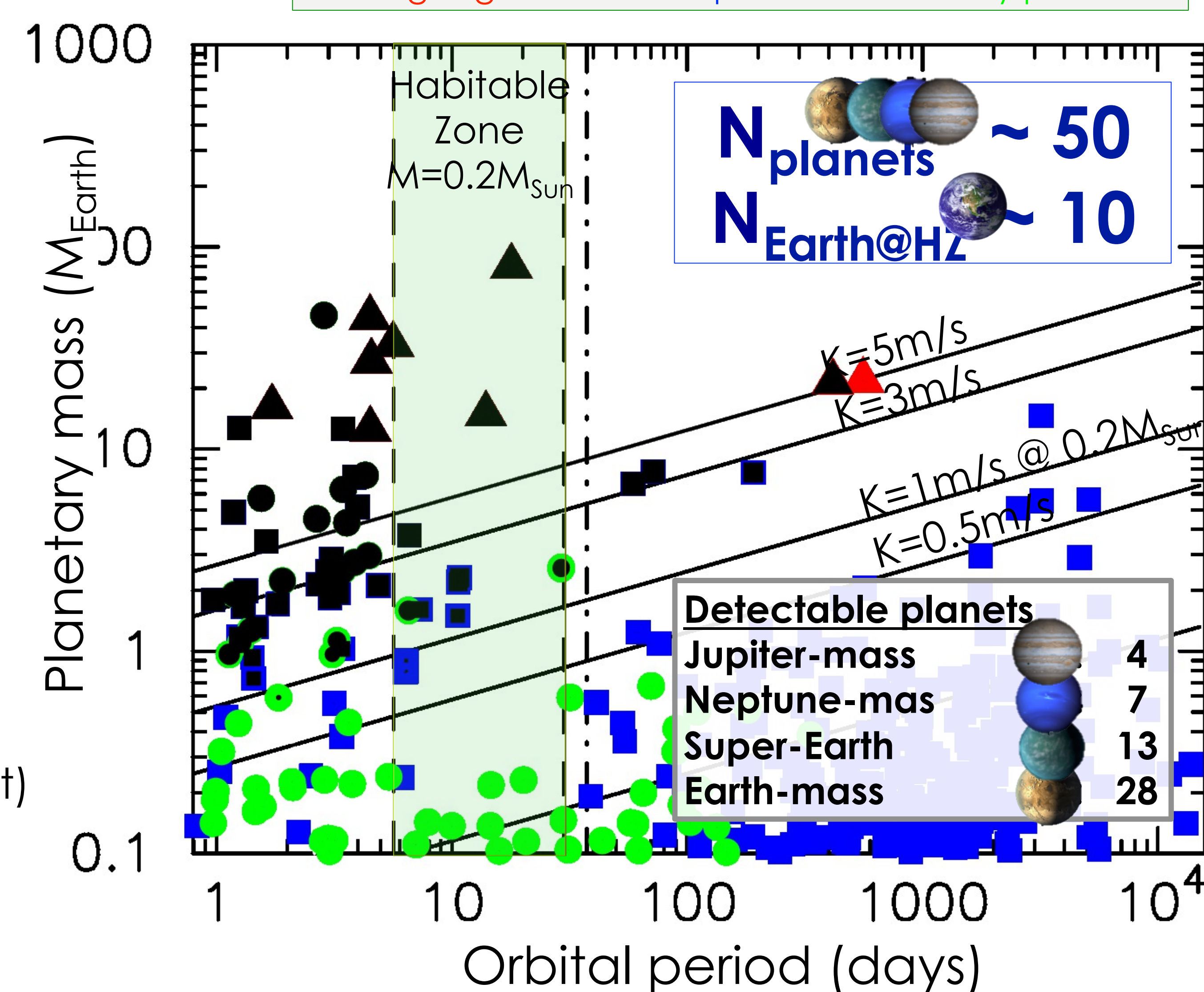
Stellar mass and brightness

- : Objects with parallax
- : Objects without H α



Observational simulation of detectable planets (black, Size indicates detection probability.)

Theoretical Simulation (Pop. synthesis by Hori, Y et al.)
▲ : gas giants ■ : ice planets ● : Rocky planets



Summary

In Subaru/IRD planet search with the specific strategies for 5 yrs, we aim to detect Earth-mass planets orbiting low-mass late-M dwarfs in their habitable zone. We would like to start a unique planet search in 2018.

Subaru/IRD detects many Earth-mass exoplanets in the habitable zone!