


Current Status of Our AO-Assisted Optical IFU Study on Galaxy-BH Coevolution

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Does the quasar-mode feedback really exist?
Is it powerful enough to quench star formation?

* Preliminary analysis for our **Kyoto 3D II + AO188 / Subaru** data

3.1'' ~ 2.2kpc


< An example >

* Structure in the **velocity**
map of narrow line region