

# Subaru-TESS Synergy



Norio Narita (U. Tokyo/ABC/NAOJ)  
on behalf of Japanese TESS consortium



# Outline

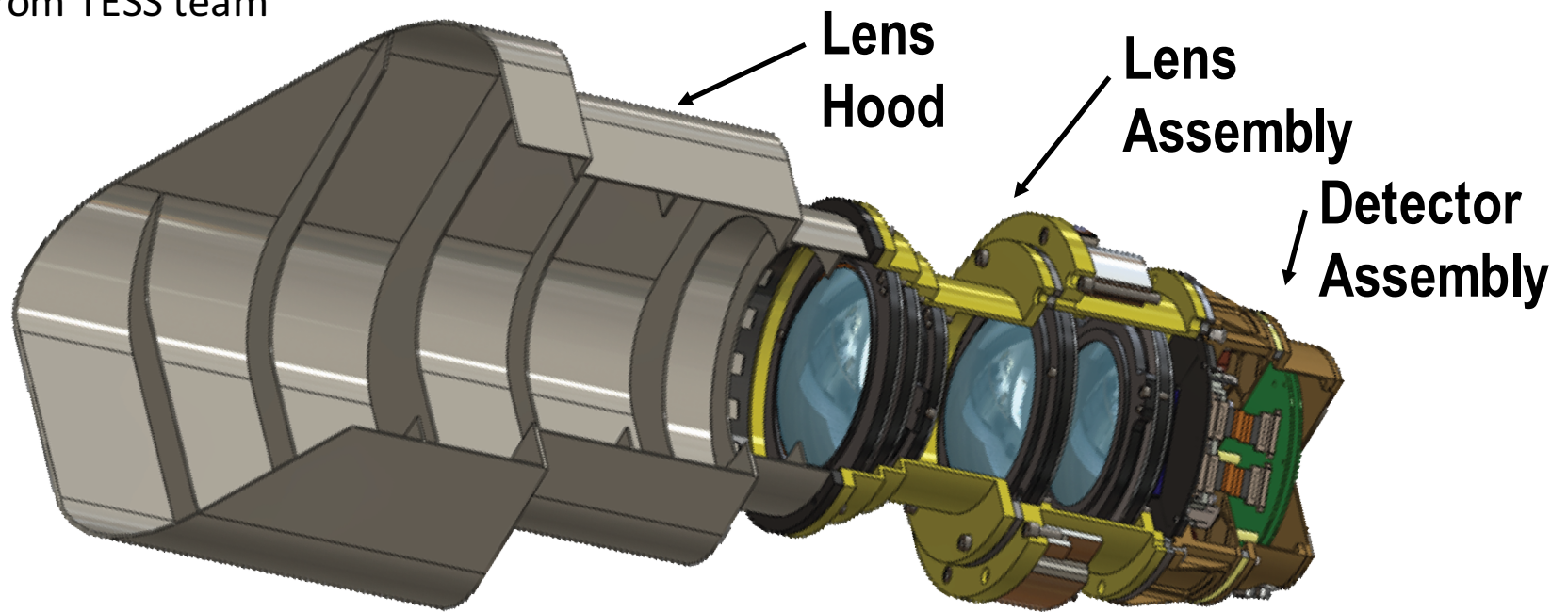
- Introduction of TESS
- Introduction of Japanese TESS Consortium
- Subaru-TESS Synergetic Science Cases
  - Exoplanets
  - Galactic Archaeology
  - Superflare stars
- Desired Instruments and Nights with Subaru
- Summary

# TESS Overview

- TESS is the next flagship space mission for exoplanet studies after Kepler mission, currently planned to launch in 2017 Dec
- Almost-all-sky survey of “transiting” planets around nearby 200,000 stars
- 2 year observation with 4 cameras (10cm diameter) each having  $24^\circ \times 24^\circ$  FoV
- Providing ~20 Earth-like transiting planets around HZ for further follow-up observations with ELTs and JWST



slide from TESS team

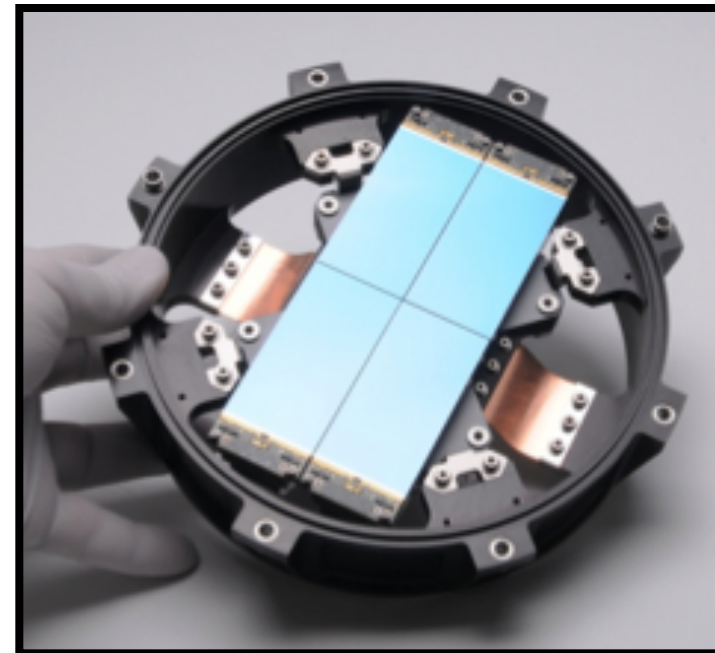


Entrance pupil diameter	10.5 cm
Bandpass	600-1000 nm
Field of view	24° x 24°
Cadence for target stars	2 min
Cadence for full frame images	30 min

four 2048 x 4096 CCDs

each CCD  
= 2048 x 2048 imaging  
+ 2048 x 2048 storing

in total  
4096 x 4096 imaging

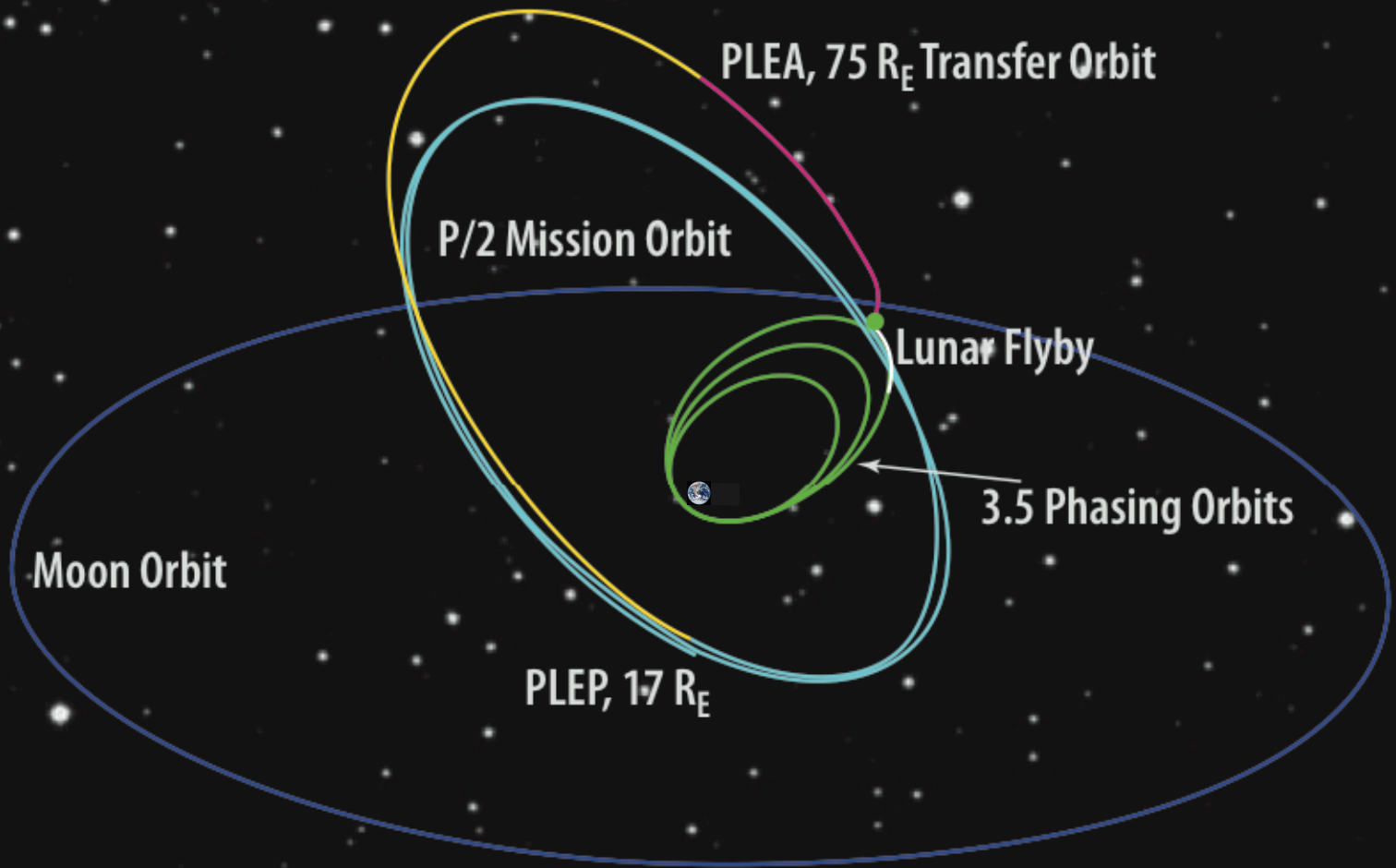




# Upcoming TESS Timeline

- Launch-9 month : Call for Guest Investigator Program (GIP)
- Launch-6 month : Proposal due for GIP
- Launch-1 month : Announcement of GIP results
- 2017 Dec 20: Launch date (currently scheduled, but may be postponed for a few months, no later than 2018 June)
  - Launch+2 month: Reach High Earth Orbit and engineering
  - Launch+3 month: Start observations from Southern hemisphere
- Start obs+6 month: Start data release (every 4 months)
  - Follow-up observations are possible from 19A semester

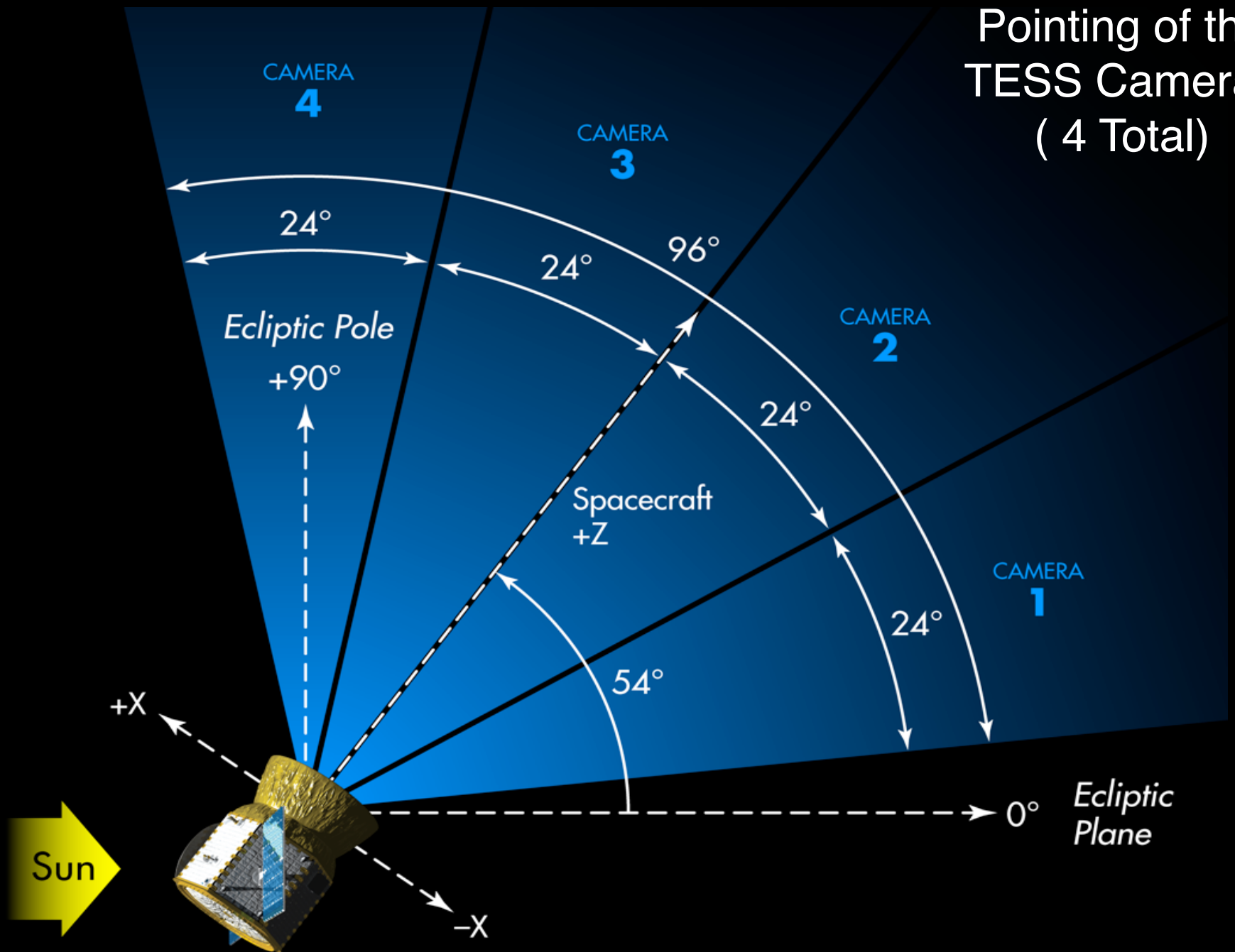
# TESS High Earth Orbit and Downlink



The orbit is stable over decades.

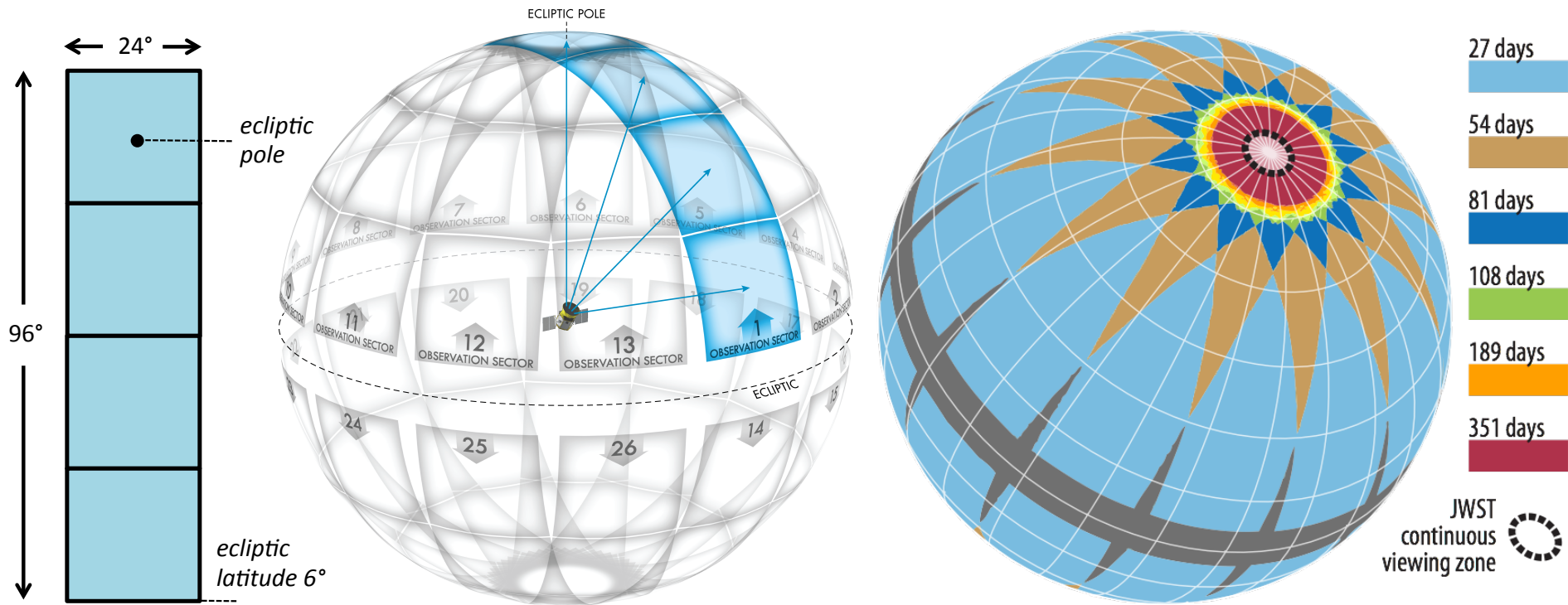


# Pointing of the TESS Cameras ( 4 Total)



slide from TESS team

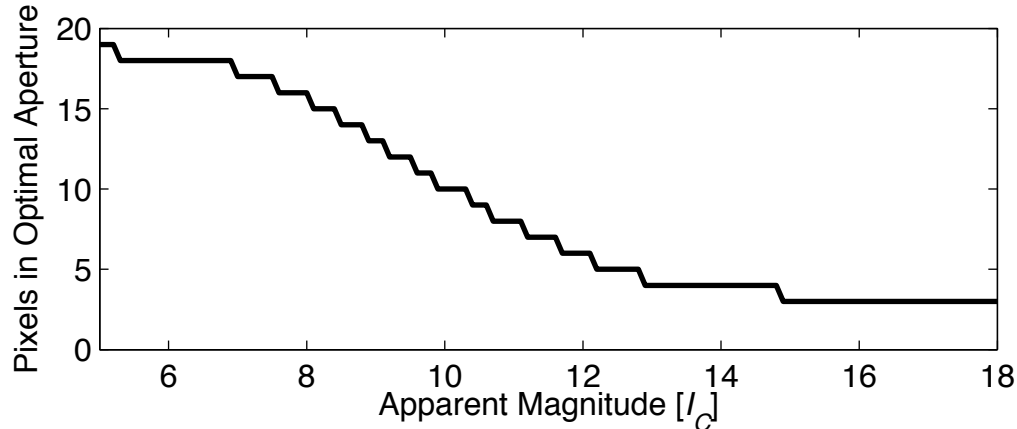
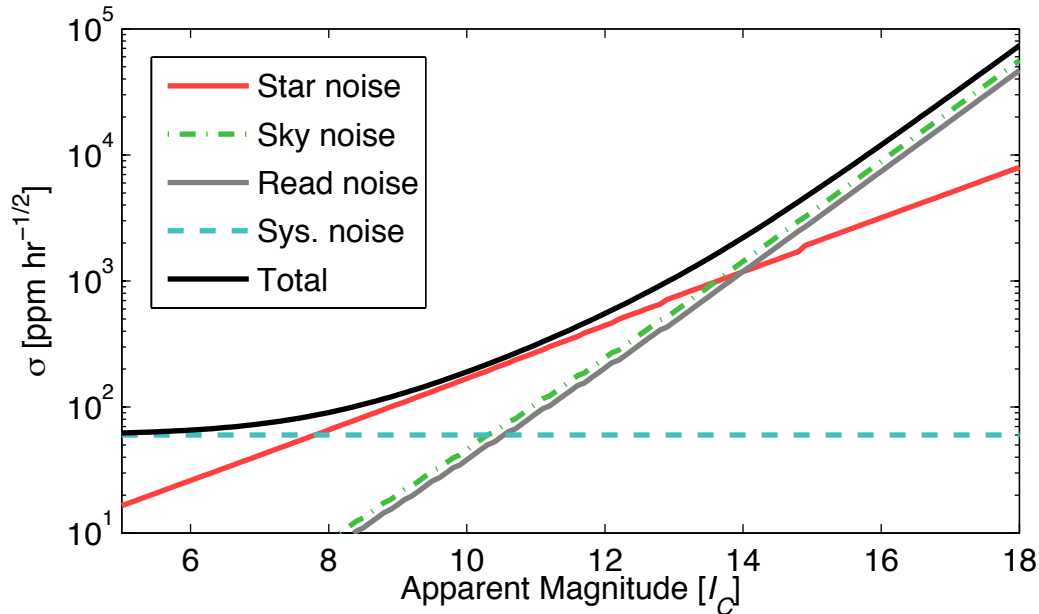
# TESS Field of View and Strategy



- TESS observes **27 days** for each rectangle sector (24° x 96°)
- Start with southern hemisphere (1 year) and then monitor northern hemisphere (1 year), and possible extension.



# TESS Expected Sensitivity



- Better than 0.1% (per hr) photometric precision for  $I$  mag < 12
- Sufficient to detect (sub)Earth-size planets around M dwarfs
- 1% precision for  $I \sim 16$
- It may be also useful for other science cases

# Guest Investigator Program

## Guest Investigator Proposals:

**\$2.5M funding per year**

### 2-minute cadence

~10,000 targets/year

“Survey” proposals encouraged

Data processed with core data

Targets and funding are decoupled

### Working with FFIs

Science that doesn't need 2-min cadence

Development of new software tools (high bar!)

Funding only

### Targets of Opportunity

Known science, unknown targets (e.g. supernovae)

ToO Science Leads identify targets for inclusion

Targets and funding are decoupled

### Discretionary Targets

Unknown science

Responding to new discoveries

Targets only

Ad hoc peer review

**TESS is also considering to make GCN alerts for transient events.**

slide from TESS team



# Introduction of Japanese TESS Consortium

## <Aims>

### 1. Detailed considerations of Subaru-TESS synergetic campaign

- What kinds of science?
- How many nights are desired?
- Which instruments are desired?

### 2. Providing latest information of TESS to Japanese community

- preparing Guest Investigator Program proposals
- 4 Japanese researchers are joining the TESS team as collaborators and can access to latest information

# Activities of Japanese TESS Consortium

## <Previous activities>

- 2015 Nov. 26-27: Japanese TESS Science Workshop
  - about 60 attendees
  - 19 talks for not only exoplanets, but also asteroseismology, galactic archeology, super-flare stars, supernovae, transient, variable stars, Wolf-Rayet stars
- 2016 Jan: call for consortium members
- 2016 August: start monthly regular meeting
  - Discussions of various science cases
  - Discussions of various available telescopes/instruments



# Current Consortium Members: 49

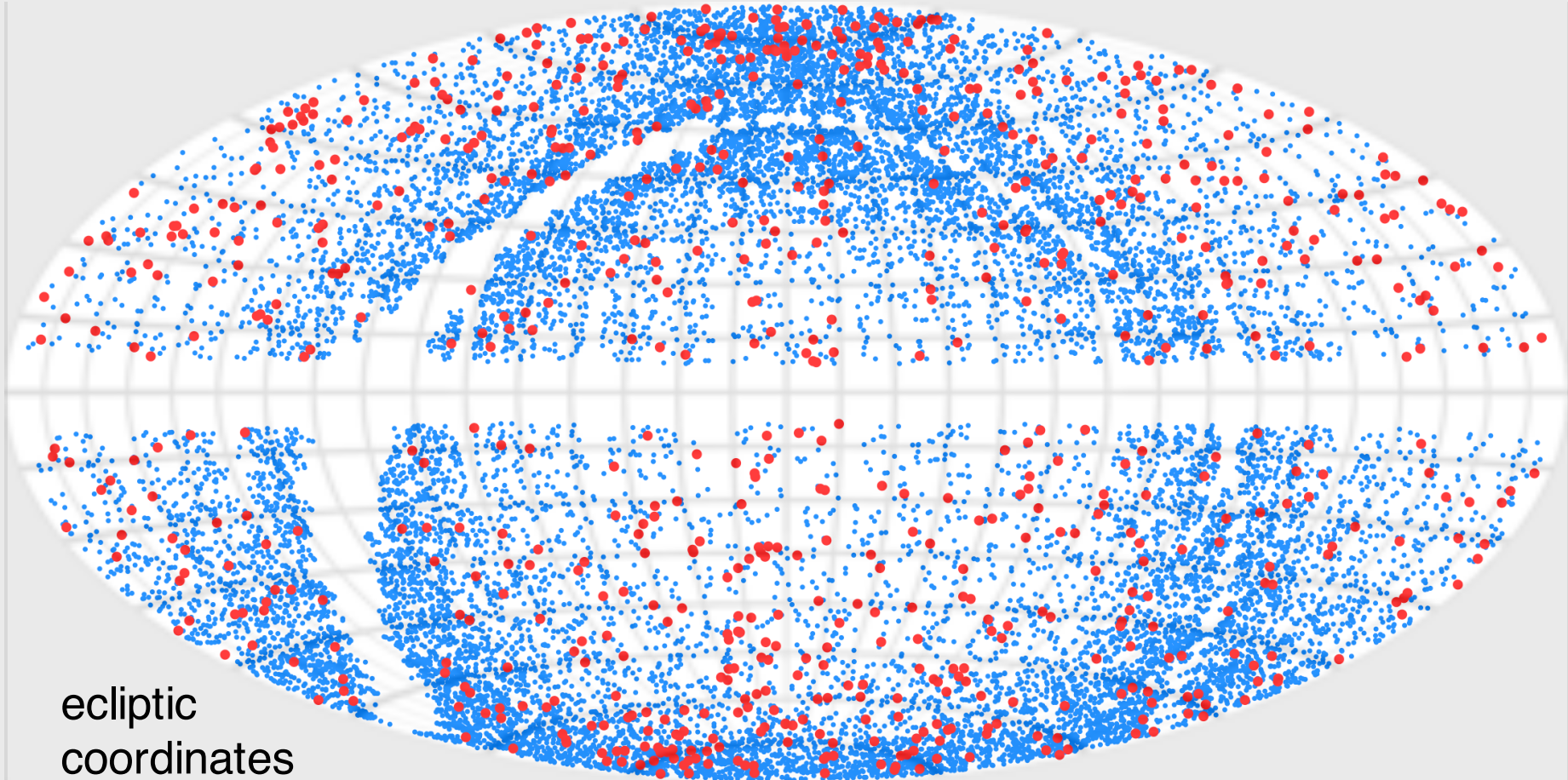
## Members' Affiliations

- University of Tokyo
  - Tokyo Institute of Technology
  - Astrobiology Center
  - NAOJ
  - Kyoto Univ.
  - Nagoya Univ.
  - Osaka Univ.
  - Japan Spaceguard Association
  - Tokyo University of Science
  - Univ. of Aizu
  - Saitama Univ.
  - Tokai Univ.
  - Chiba Institute of Technology
- Professors : 13
  - Assoc. Professors : 10
  - Assist. Professors : 10
  - Postdoc/researchers : 9
  - Graduate students : 7

# Subaru-TESS Synergetic Science Cases Under Consideration

- Exoplanets
- Galactic Archaeology
- Superflare Stars

# Simulated TESS Planet Detections



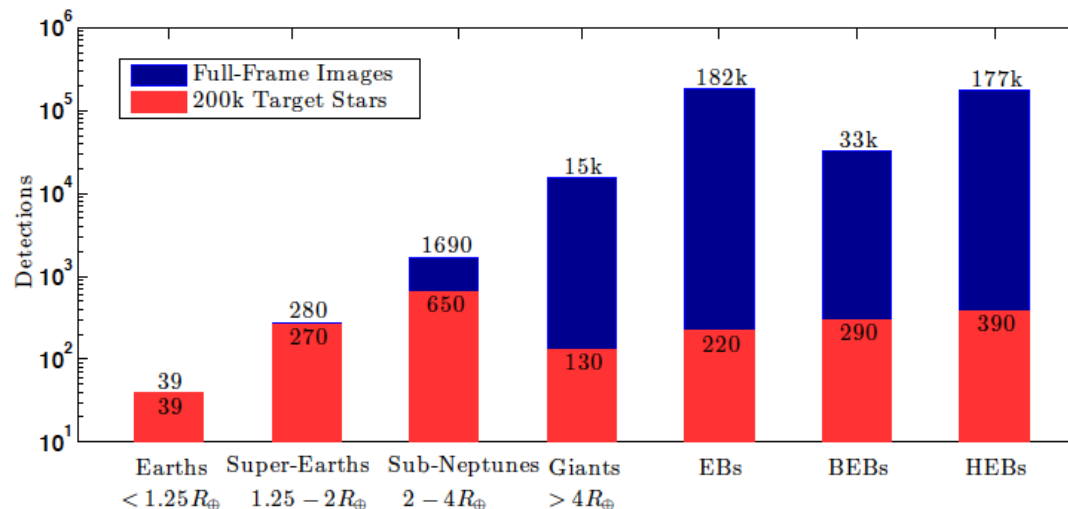
Sullivan et al. (2015)

- detectable planets around pre-selected target stars
- detectable planets around other stars in full-frame images



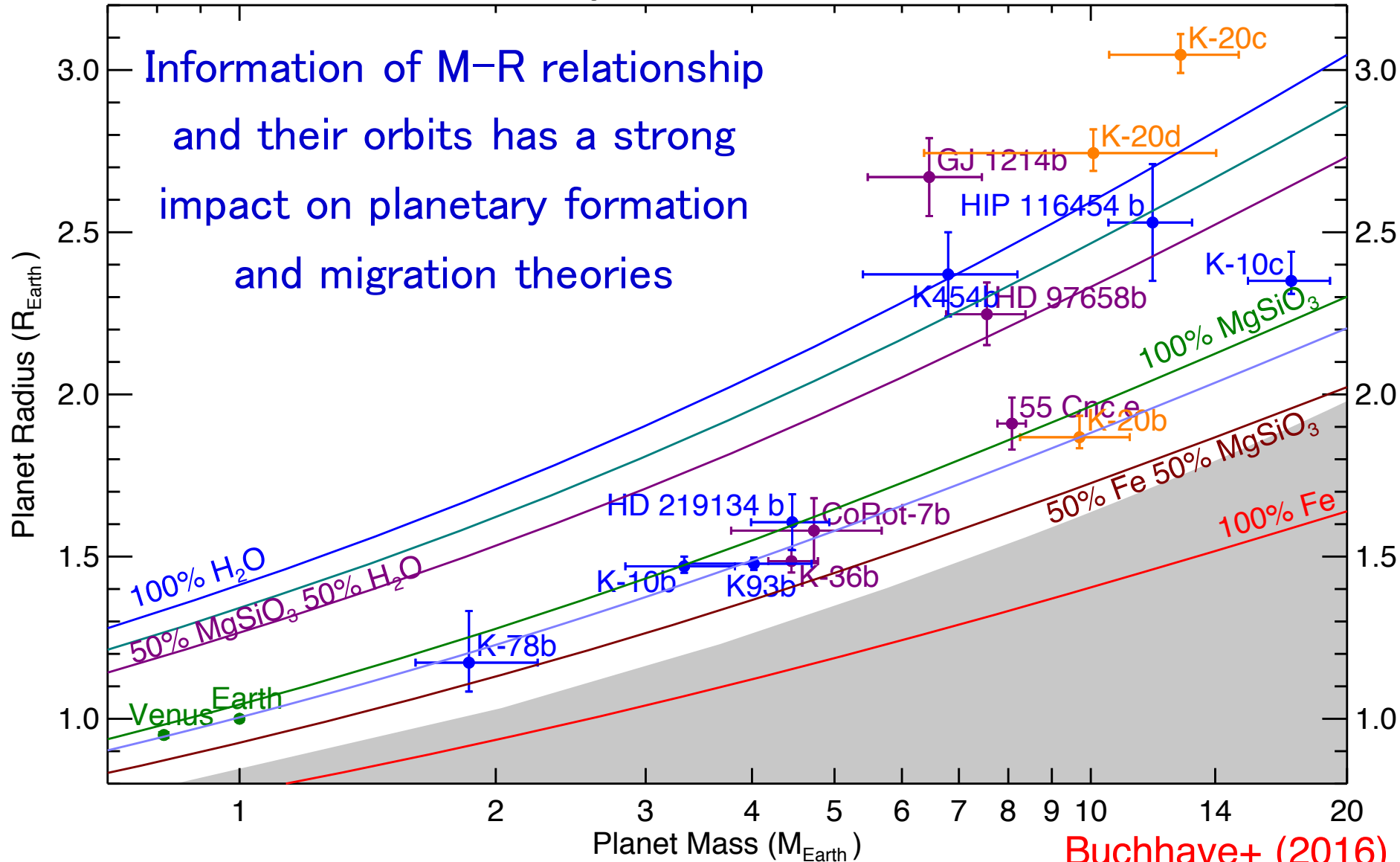
# TESS Planet Yield Simulation (Sullivan+2015)

- about 300 planets: smaller than  $2 R_{\text{earth}}$ 
  - among those, ~165 planets are orbiting M dwarfs
  - ~100 planets are orbiting  $L < 10$  host stars
  - about 20 planets are orbiting within or near habitable zone
- about 650 planets: between  $2 R_{\text{earth}}$  and  $4 R_{\text{earth}}$ 
  - boundary of “super-Earth” and “mini-Neptune”
  - also important targets for planetary science



# Subaru/IRD can reveal mass-radius relationship of small planets in detail

Information of M-R relationship  
and their orbits has a strong  
impact on planetary formation  
and migration theories



Buchhave+ (2016)

# How many planets are observable from Hawaii?

Mid-late M dwarfs with  $J < 11$  and max. elevation  $> 45$  deg, from Hawaii

	Number	$< 2R_E$	$2-4R_E$	$< 2R_E$ $0.2 < S/S_E < 1.5$	$< 2R_E$ $0.2 < S/S_E < 0.8$
$\leq 3000K$	4	2	2	2	1
$\leq 3100K$	3	2	1	2	2
$\leq 3200K$	1	1	0	1	1
$\leq 3300K$	9	7	2	7	5
$\leq 3400K$	30	20	10	4	0
Total	47	32	15	16	9

# Desired Subaru Instruments/Nights

- Concentrate on 30 favorable & interesting planets
  - Less than 3400K,  $J < 11$ , not rapidly rotating or active
  - (mostly  $J > 10$  for which 8m class telescope are needed)
  - 15 planets each for 1-2  $R_{\text{Earth}}$  and 2-4  $R_{\text{Earth}}$  planets
  - Including  $\sim 10$  planets in the habitable zone with  $< 2 R_{\text{Earth}}$
  - $\sim 360$  hours with Subaru/IRD are desired to determine the mass and eccentricity of the planets
- About 60 nights including weather success rate



# Other Exoplanet Science Cases

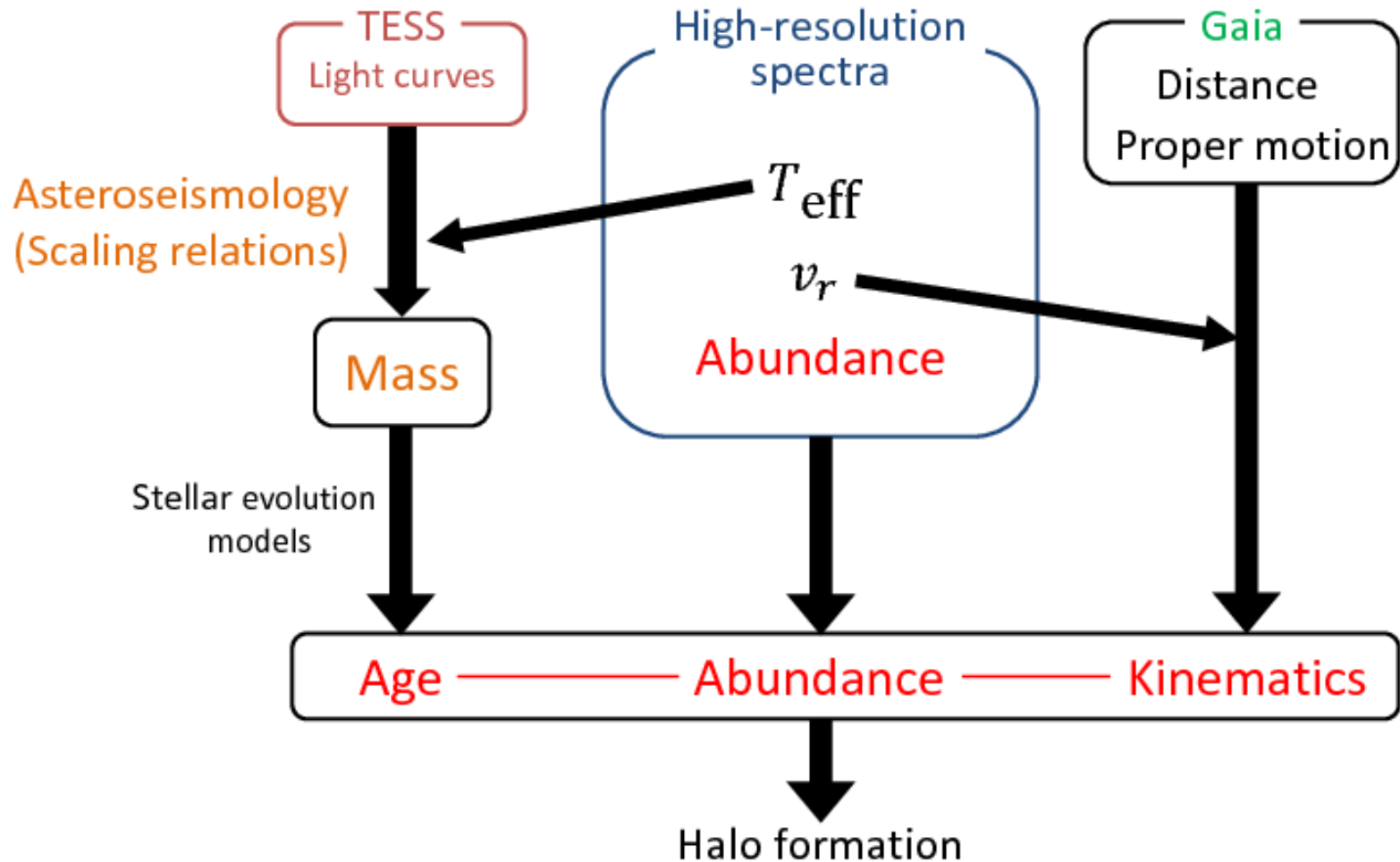
- Characterizations of planetary atmospheres/orbits
  - Transmission spectroscopy to probe planetary atmospheres
  - Doppler tomography to measure spin-orbit misalignments
- Detecting Niche Exoplanets
  - giant planets around giant (III) stars
  - giant planets around early-type (BA) stars
  - planets around young stars
  - planets around stars in stellar clusters

# TESS and Galactic Archaeology

- TESS is also useful for asteroseismology
  - 60 targets with 20 sec sampling
  - 750 targets with 2 min sampling
  - mainly red giants in the Milky Way galaxy
  - One can propose more 2 min sampling targets in GIP
- This is **the first opportunity** to investigate stars in the galactic halo by space-photometric asteroseismology
  - CoRoT, Kepler, K2 observed only limited volume

# Galactic Archaeology with TESS+Subaru+GAIA

Courtesy of T. Matsuno (Sokendai/NAOJ)

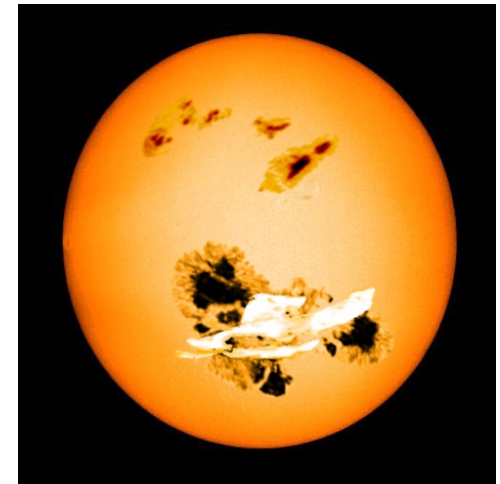
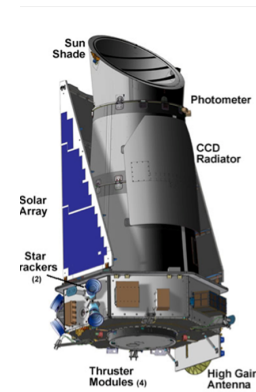
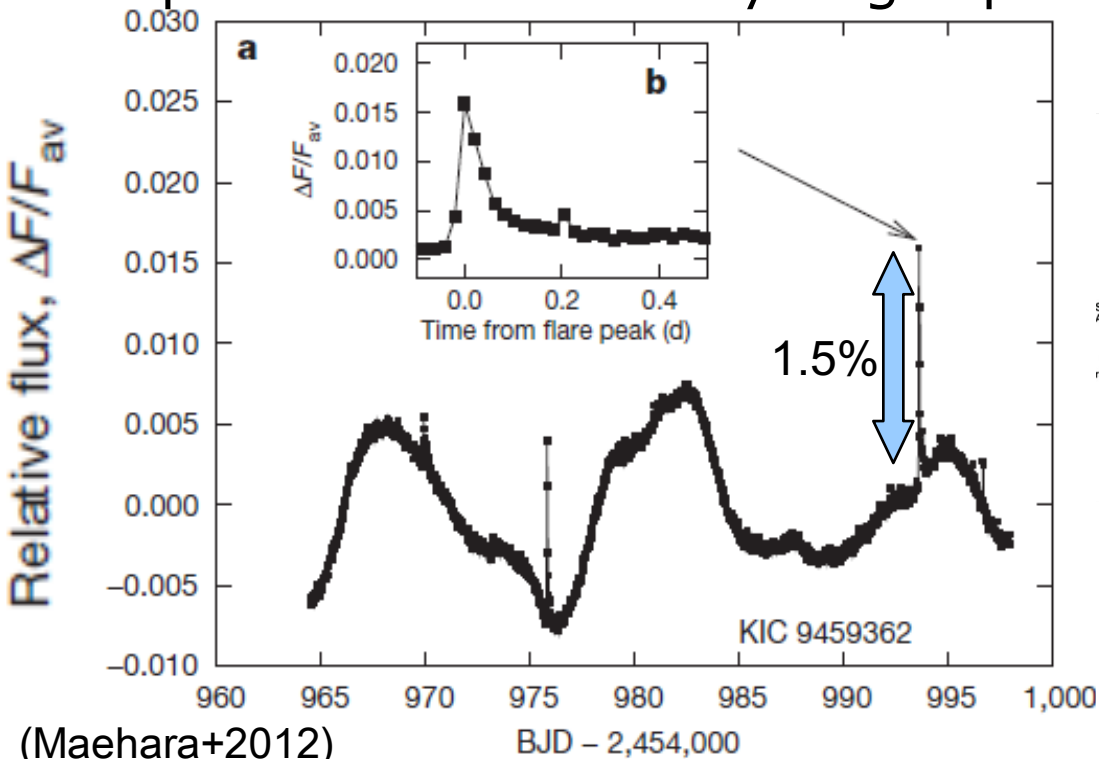


# Desired Subaru Instruments/Nights

- Subaru HDS offers high-spectral resolution and high-SNR with the new image slicers (Tajitsu et al. 2012)
- Can determine  $T_{\text{eff}}$ ,  $v_r$ , and abundance of 100 or more red giants in the galactic halo
- About 15 nights (for 100 stars) are desired (including weather success rate)

# Superflare Stars

- Previous studies with Kepler data (Maehara+2012 Nature, etc)  
⇒ Discoveries of about 300 superflare stars (10-10,000 times energetic than maximum Solar flares)
- Quasi-periodic variations suggest the presence of spots
- Variations enable to estimate rotation period and size of spots  
⇒ suggest Solar-like slowly rotating stars may also show superflares if extremely large spots exist



artist's illustration  
of a superflare star

Courtesy of Y Notsu (Kyoto)



# TESS and Superflare Stars



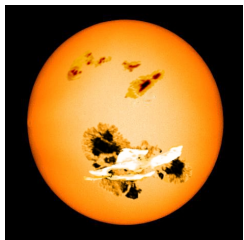
TESS will reveal

- occurrence rate, stellar rotation period, and size of spots
  - (demerit) only photometric & only  $\geq 30$  days observation  
⇒ Follow-up observations are necessary
  - (merit) nearby & all-sky superflare stars can be detected
- ⇒ For especially interesting targets
- Long-term photometric monitoring for spot monitoring
  - Direct time-resolved spectroscopy of flares
  - Direct measurements of magnetic field using Zeeman effect

More detailed studies are possible than faint Kepler stars



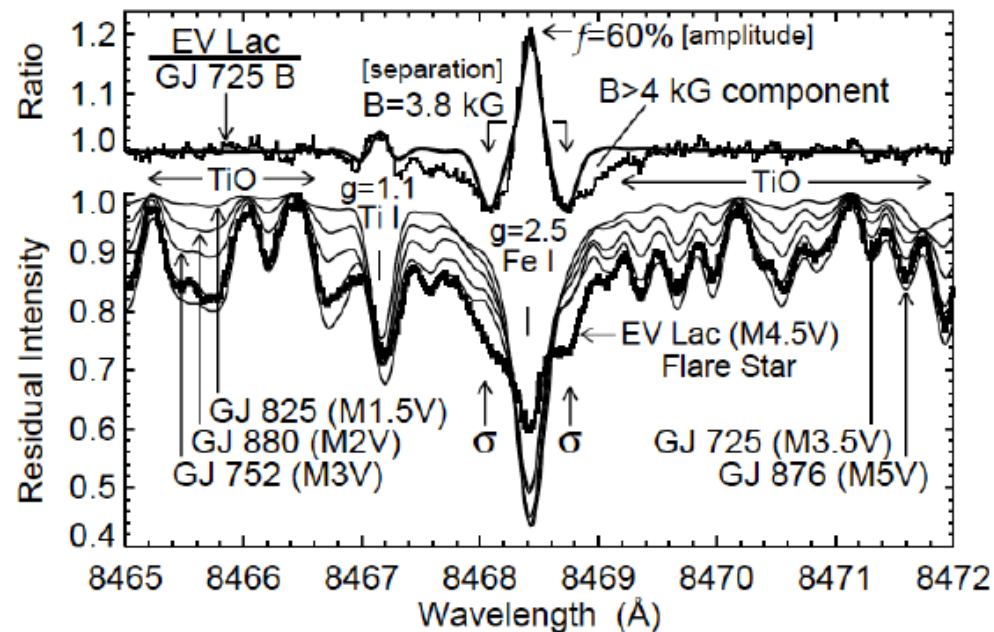
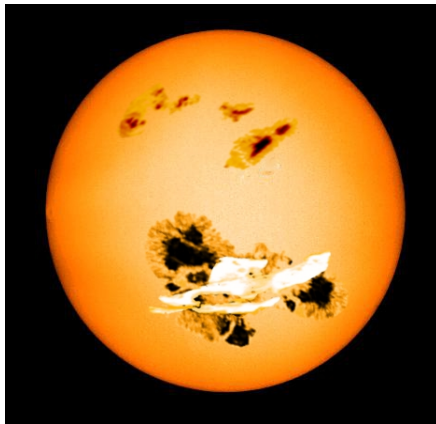
Can uncover a dynamo mechanism which generates extremely large spots



# Follow-up Study with Subaru/HDS

- Subaru/HDS can provide high spectral resolution ( $R \sim 160,000$ ) and high SNR ( $\text{SNR} \sim 200\text{-}300$ )
- It is not capable with 3-4m telescopes (including Kyoto 3.8m)
- High spectral resolution and high SNR spectroscopy for 20 MK superflare stars ( $V < 10$ ) to measure the line splitting caused by the Zeeman effect
- Enables direct measurements of the strength of the magnetic field of photospheres for the superflare stars
- Make comparisons with the strength of the magnetic field of corona, which can be derived from flare energy and duration (Namekata+2017 in prep)

→ About 10 nights are desired



# Summary of Desired Instruments/Nights

- Exoplanets
  - Detecting small planets around M dwarfs : IRD, 60 nights
  - Atmospheres of small planets : MOIRCS, 10 nights
  - Spin-orbit misalignments of small planets : IRD, 10 nights
  - Detecting giant planets around giants : HDS, 15 nights
  - Detecting giant planets around BA stars : HDS, 10 nights
  - Detecting planets around young/cluster stars : IRD/HDS, 10 nights
- Superflare stars
  - Measurements of the Zeeman effect : HDS, 10 nights
- Galactic Archaeology
  - Studying galactic halo with TESS/Subaru/GAIA : HDS, 15 nights
- Science cases are still growing

# What Subaru can do with TESS?

- Subaru can follow-up new transiting planets and other interesting objects (red giants for asteroseismology, superflare stars, etc) soon after the data releases in a timely and efficient manner
- High scientific productivity is expected

We hope a chance of call for a medium size program  
similar to WFIRST

# Summary

- TESS is the next flagship exoplanet mission to search for interesting exoplanets in the solar neighborhood
- Not only exoplanets, but also other fields (such as galactic archeology, superflare stars, transients, etc) can receive benefits from TESS data
- Subaru-TESS synergetic campaign would benefit Subaru and Japanese astronomical community