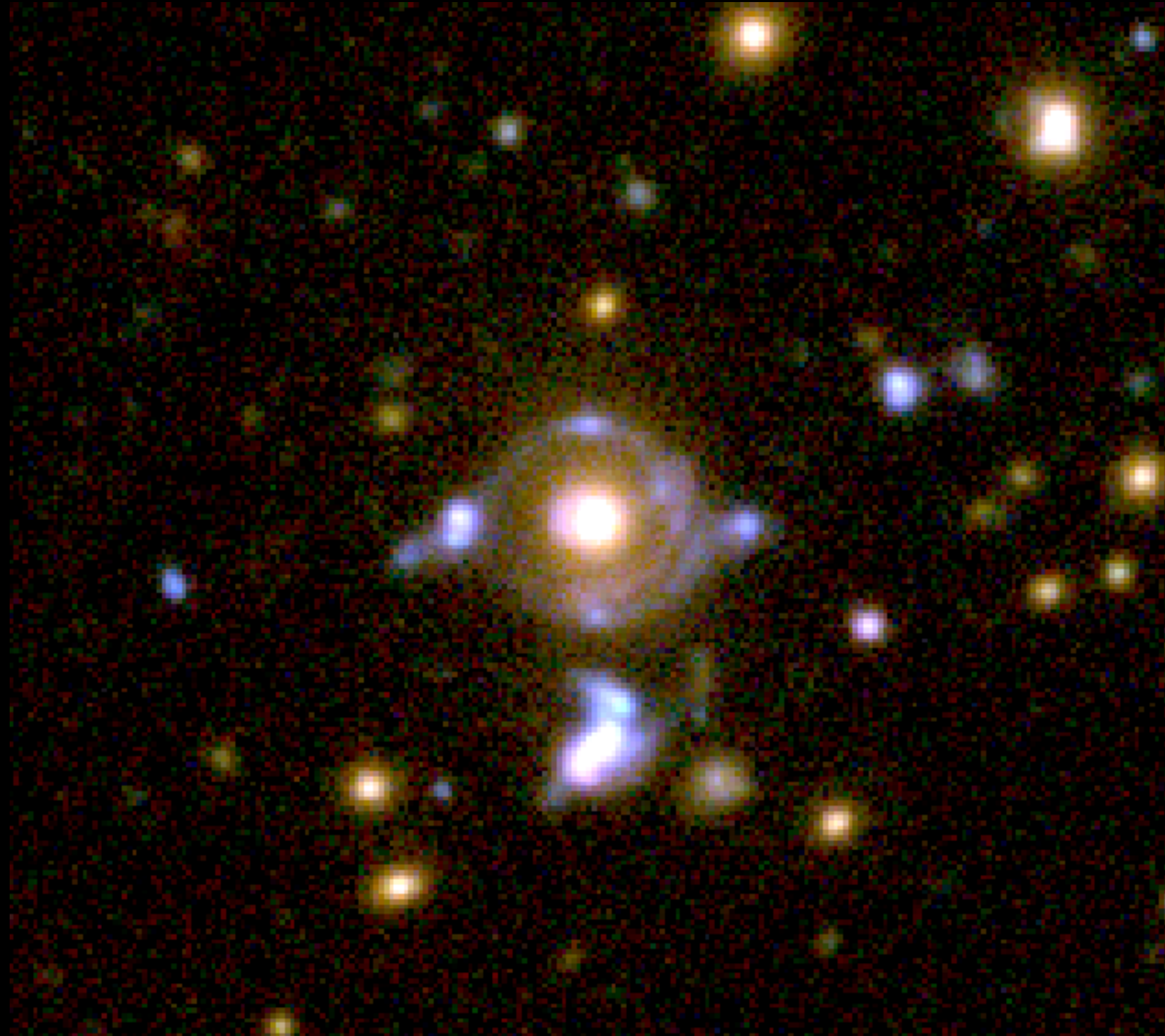


A Spectroscopically-Confirmed Double Source Plane Lens in the HSC SSP

Tanaka, Wong, et al. 2016, ApJ, 826, L19



Kenneth Wong

EACOA Fellow

National Astronomical Observatory of Japan

Masayuki Tanaka (NAOJ)

Anupreeta More (IPMU)

Arsha Dezuka (Kyoto)

Eiichi Egami (Arizona)

Masamune Oguri (Tokyo)

Sherry Suyu (MPA)

Alessandro Sonnenfeld (IPMU)

et al.

Subaru Users' Meeting FY2016

NAOJ

January 11, 2017

Strong Gravitational Lensing

- Background object (source) magnified by foreground object (lens)
- Multiple images → create lens model
- What can we learn from lensing?
 - total mass (within Einstein radius)
 - mass profile slope
 - ellipticity/orientation
 - substructure (both luminous and dark)
 - intrinsic (unlensed) source flux
 - detect/resolve source features by taking advantage of magnification
 - cosmology from time-delay lensed quasars (e.g., Suyu+2016, Wong+2016, Bonvin+2016)
- Surveys to build statistical samples of lenses (e.g., SLACS, SL2S, BELLS)
 - Mostly $z \lesssim 0.4$, up to $z \lesssim 0.8$
 - need deep, wide-area, high-resolution surveys to improve statistics, especially at higher z

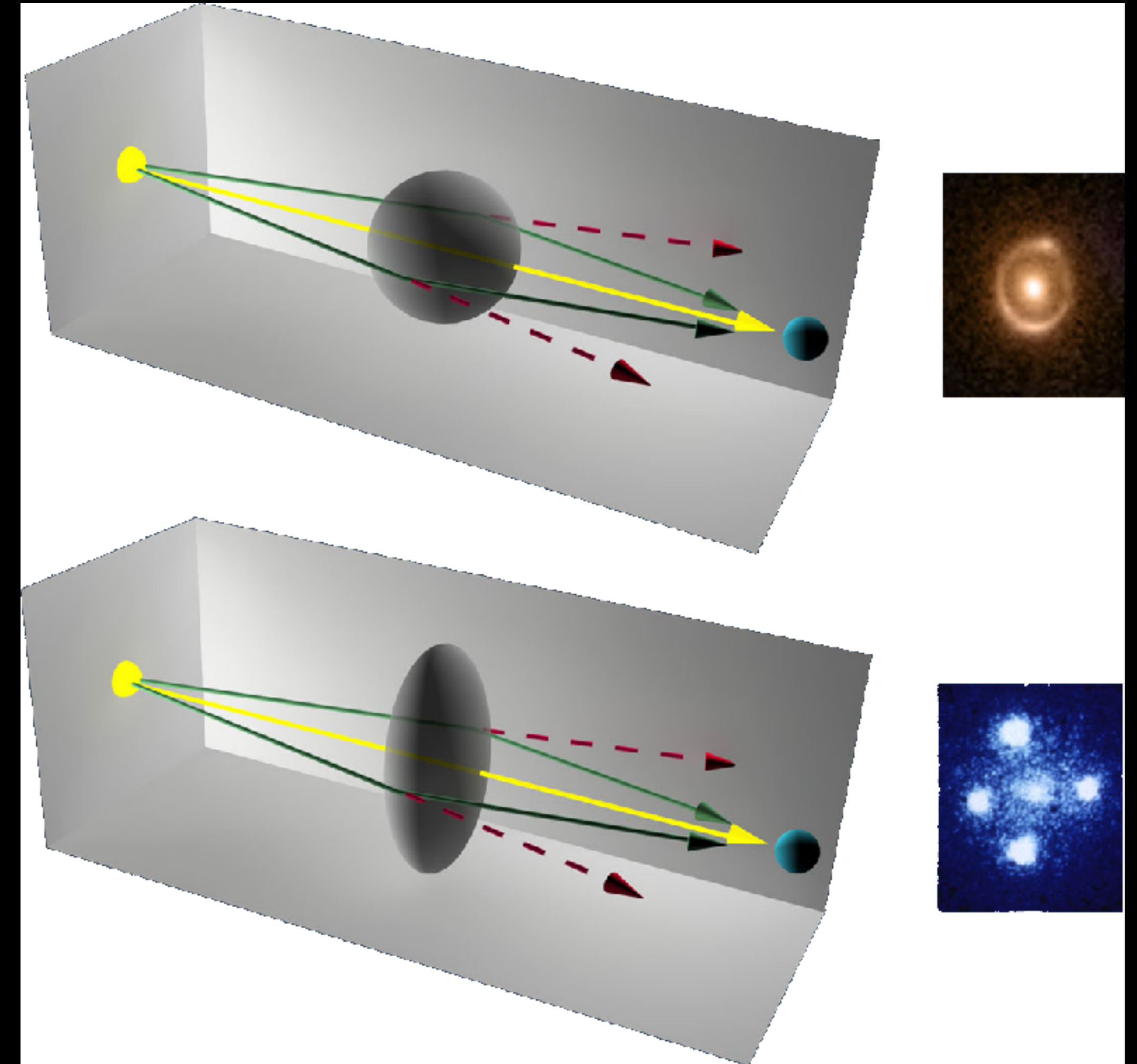
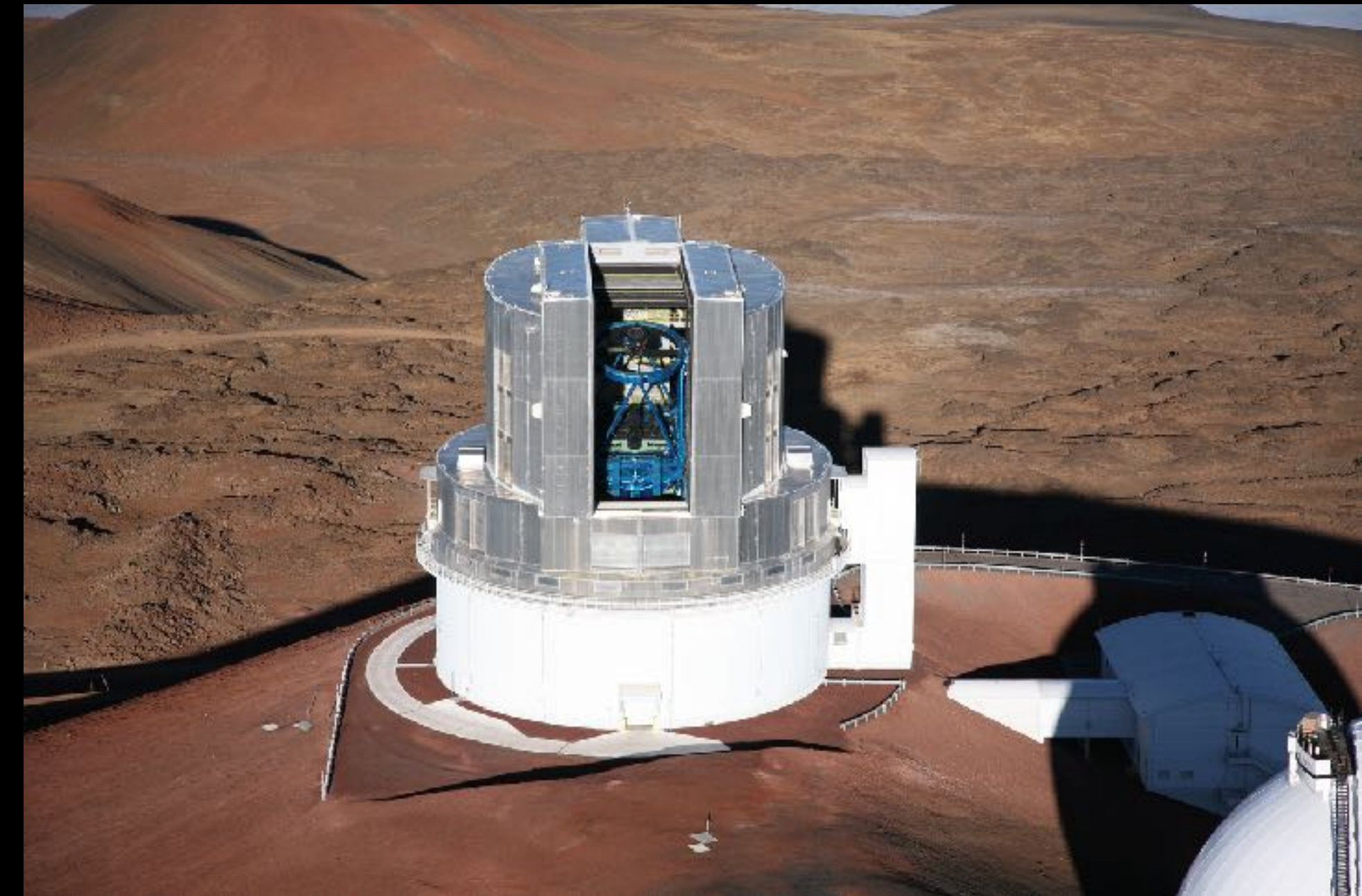


Image credit: ASIAA EPO

Discovering New Lenses with HSC

- Hyper Suprime-Cam SSP
 - 1400 deg² *grizy* imaging to $r_{AB} \sim 26$
- ~30 group/cluster lens candidates
 - found through inspection of CAMIRA clusters (Oguri 2014)
- ~10 lensed quasar candidates
 - CHITAH algorithm (Chan+2015)
- Survey of Gravitational lens Objects in HSC Imaging (see talk by A. Sonnenfeld)
- Many lenses found serendipitously
 - potential to discover exotic lenses



Discovering New Lenses with HSC

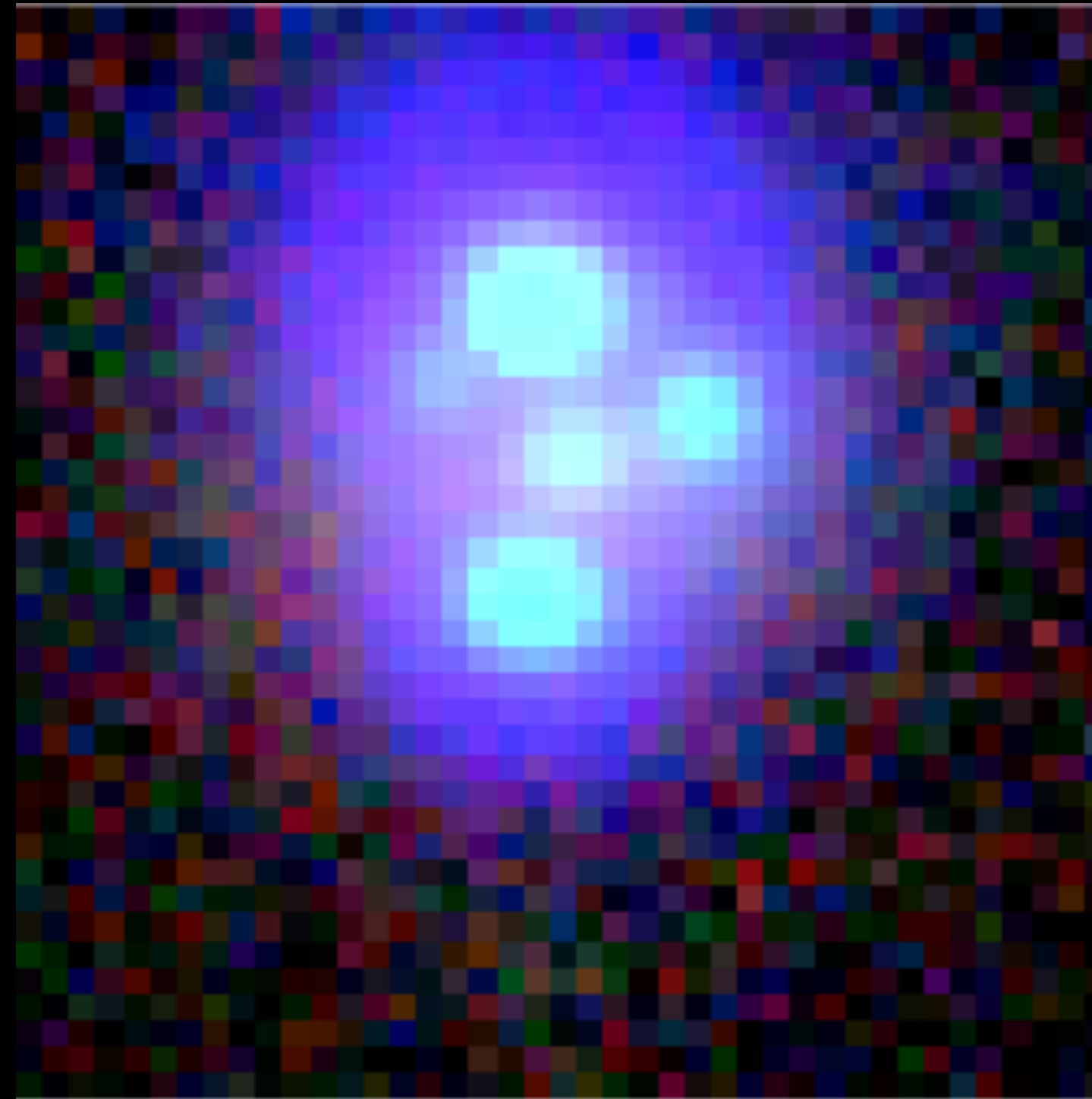
- Hyper Suprime-Cam SSP
 - 1400 deg² *grizy* imaging to $r_{AB} \sim 26$
- ~30 group/cluster lens candidates
 - found through inspection of CAMIRA clusters (Oguri 2014)
- ~10 lensed quasar candidates
 - CHITAH algorithm (Chan+2015)
- Survey of Gravitational lens Objects in HSC Imaging (see talk by A. Sonnenfeld)
- Many lenses found serendipitously
 - potential to discover exotic lenses



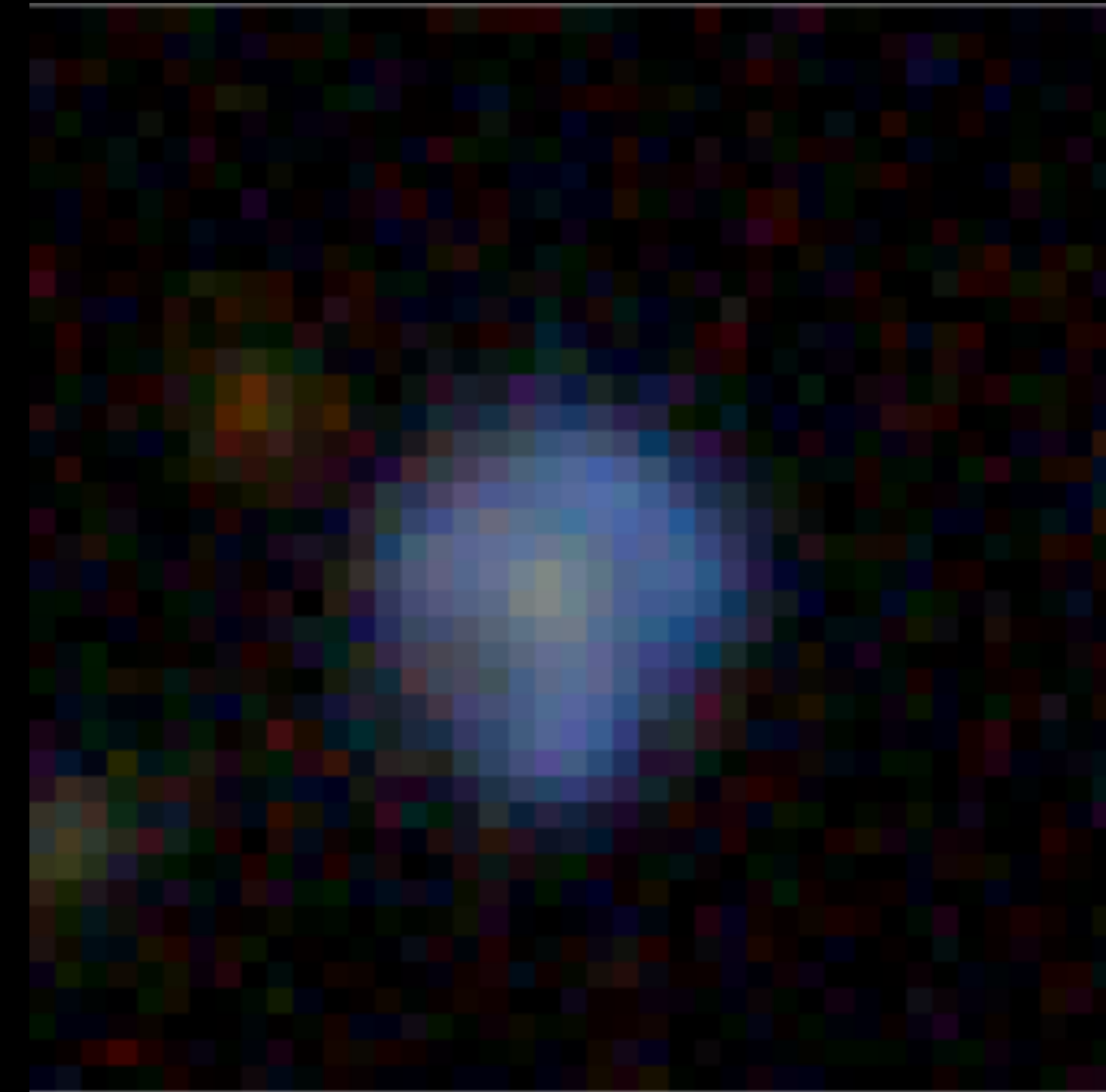
More et al. (in prep)

Discovering New Lenses with HSC

- Hyper Suprime-Cam SSP
 - 1400 deg² *grizy* imaging to $r_{AB} \sim 26$
- ~30 group/cluster lens candidates
 - found through inspection of CAMIRA clusters (Oguri 2014)
- ~10 lensed quasar candidates
 - CHITAH algorithm (Chan+2015)
- Survey of Gravitational lens Objects in HSC Imaging (see talk by A. Sonnenfeld)
- Many lenses found serendipitously
 - potential to discover exotic lenses

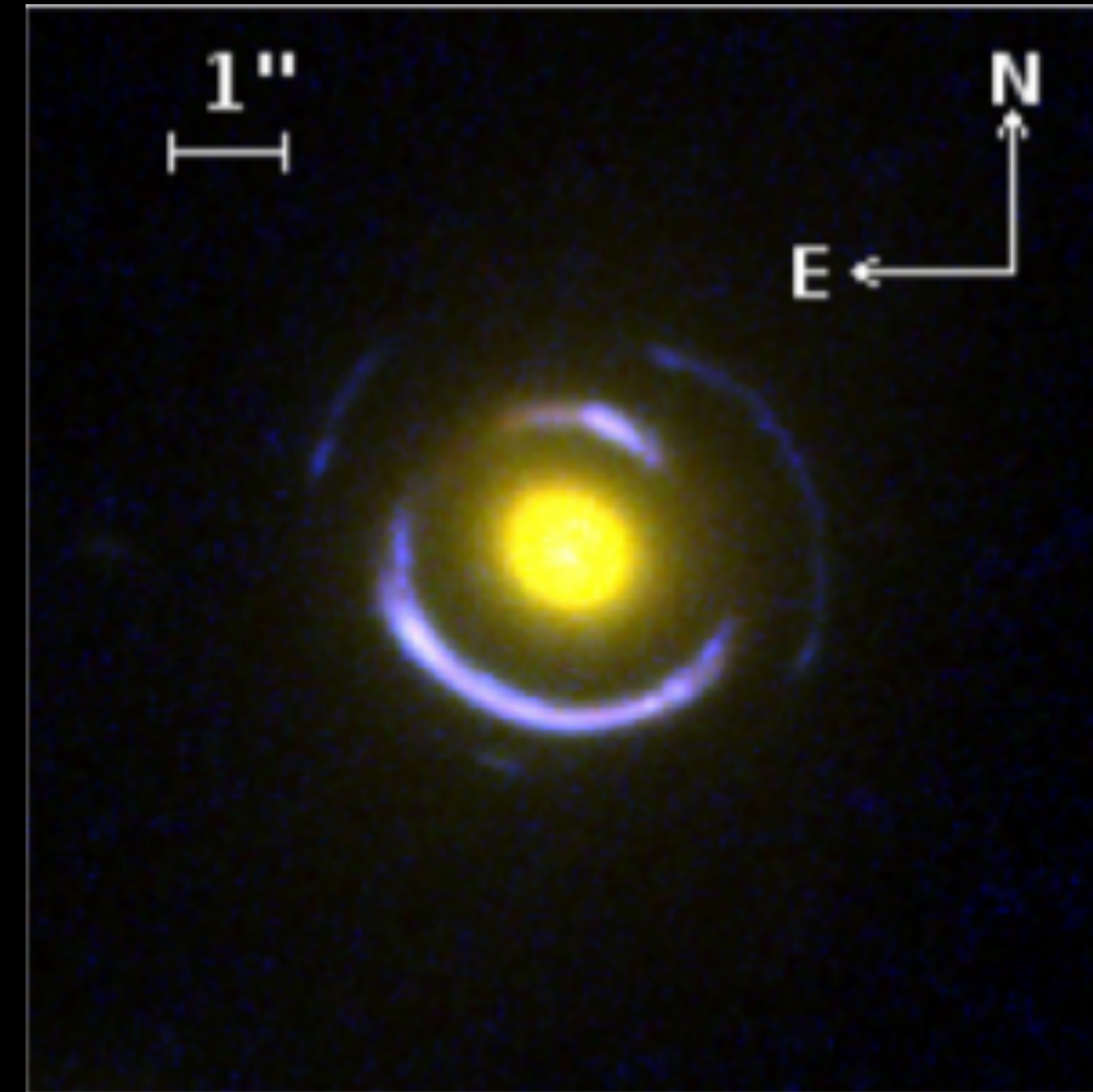


Chan et al. (in prep)

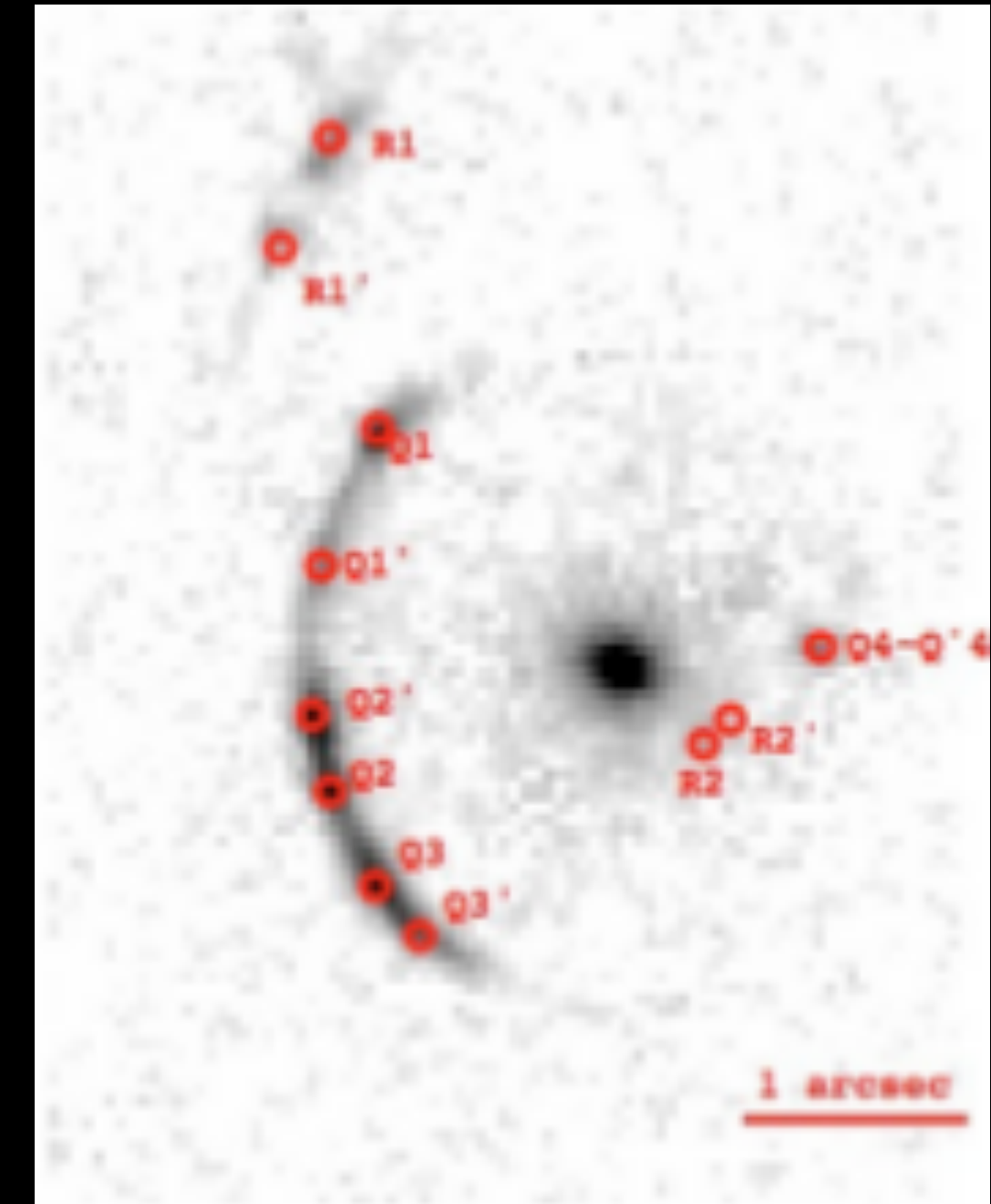


Double Source Plane Lenses

- Double source plane (DSP) lenses
 - two sources at distinct redshifts being lensed by the same galaxy
 - extremely rare, only a handful known (e.g., Gavazzi+2008, Tu+2009)
- Constraints from DSP lenses
 - can constrain Ω_m and w , independent of H_0 (e.g., Collett+2012,2014)
 - can break model degeneracies (although see Schneider 2014), constrain mass structure, IMF (e.g., Sonnenfeld+2012)

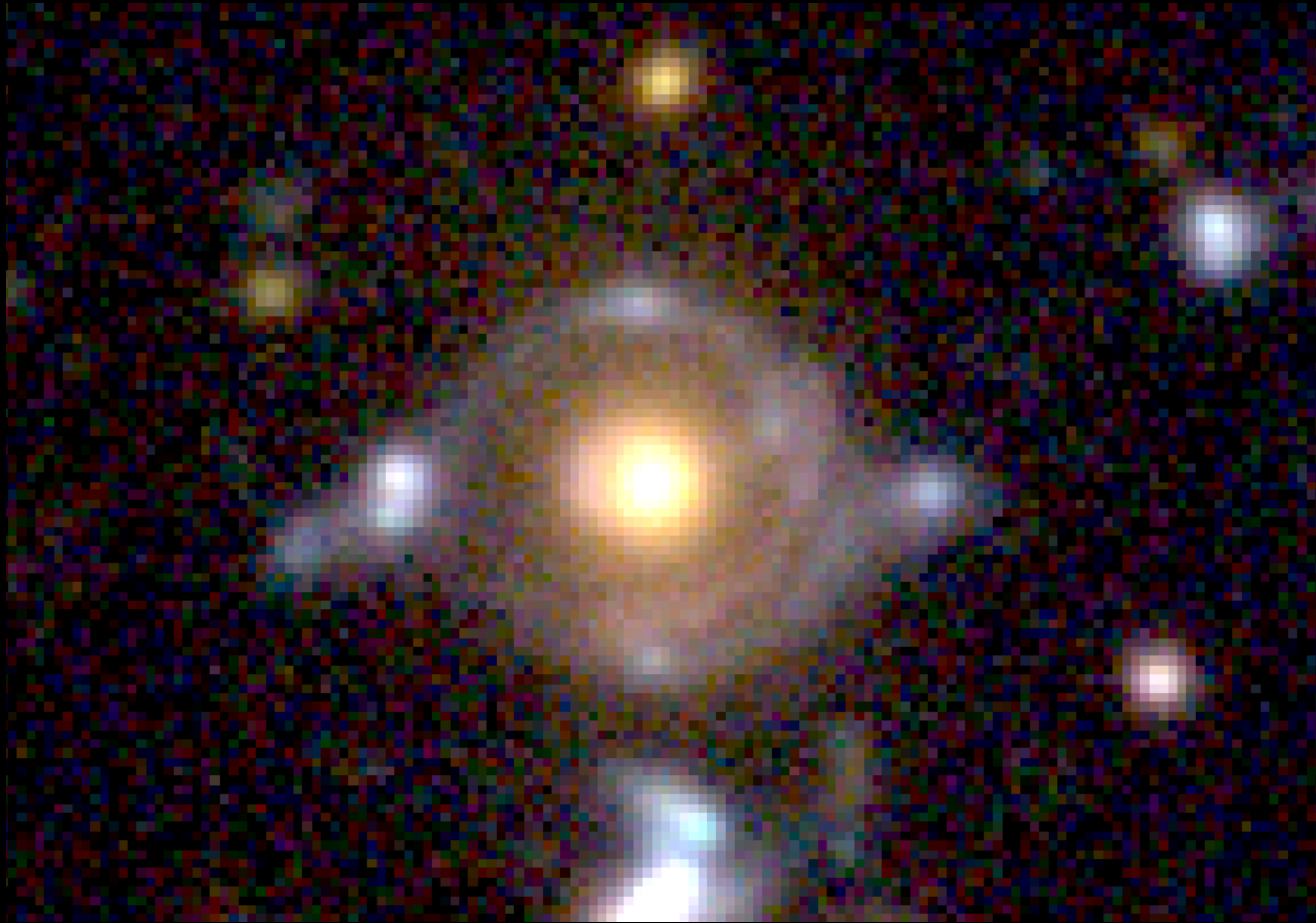


Sonnenfeld+2012



Tu+2009

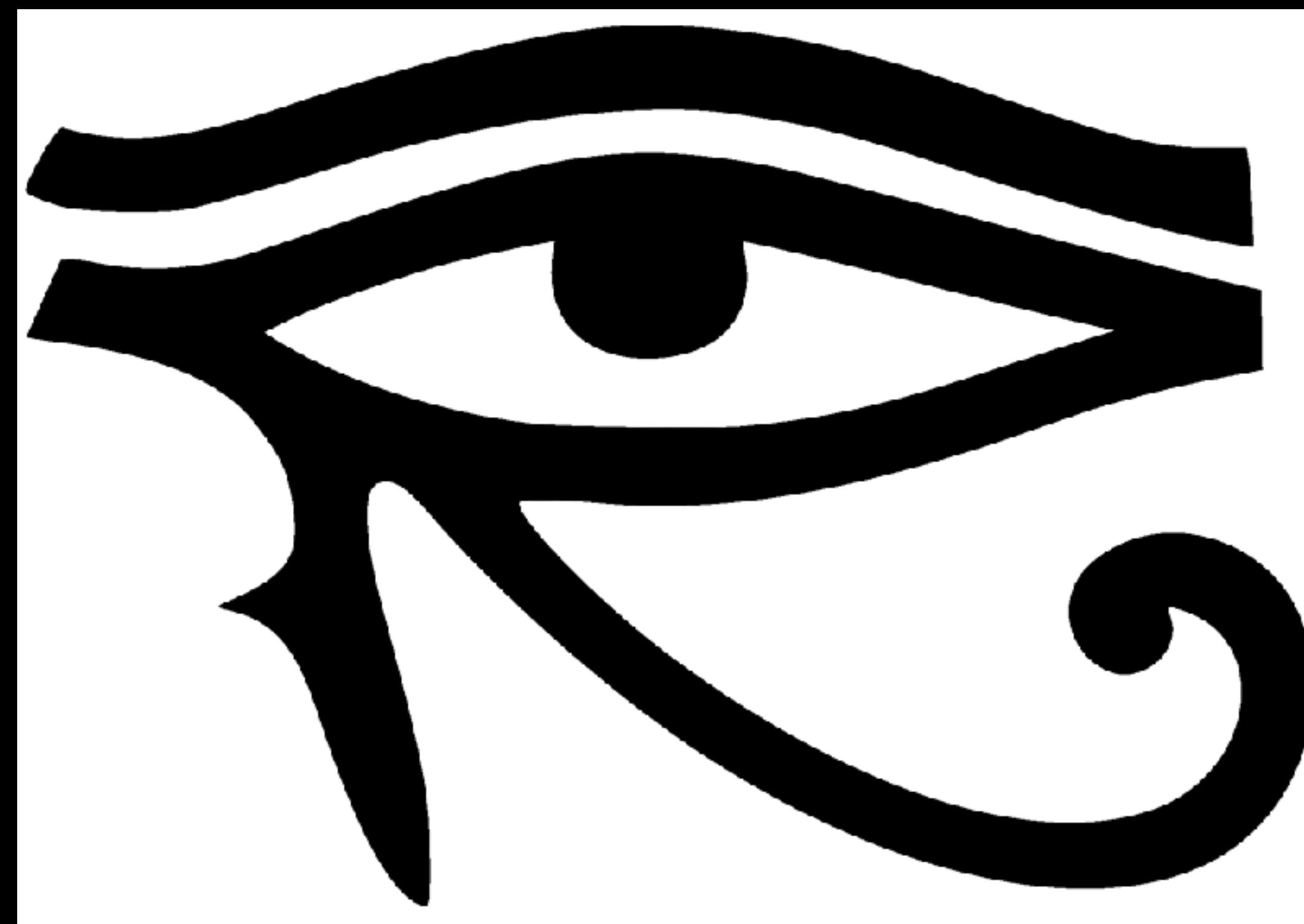
Discovery of the First DSP Lens in the HSC SSP



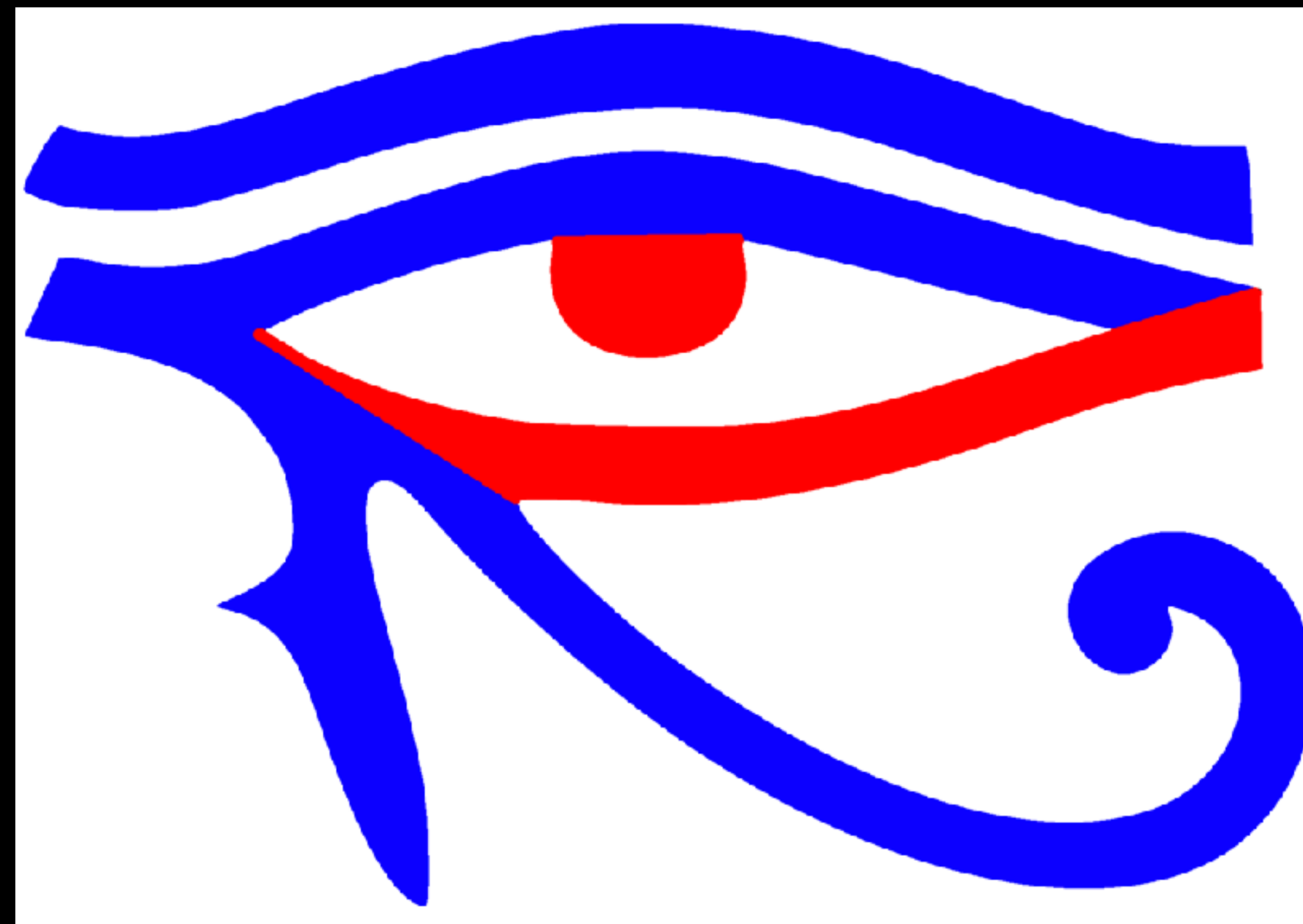
Tanaka, Wong, et al. (2016)

- Double source plane (DSP) lens serendipitously discovered in HSC SSP
 - inner arc and counterimage (S1)
 - outer Einstein ring with central knot (S2)
- All redshifts spectroscopically confirmed from Magellan/FIRE observations
 - $z_L = 0.795$
 - $z_{S1} = 1.30$
 - $z_{S2} = 1.99$
 - first known DSP lens with both source redshifts spectroscopically confirmed
- Press release: naoj.org/Pressrelease/2016/07/25/index.html

The “Eye of Horus”

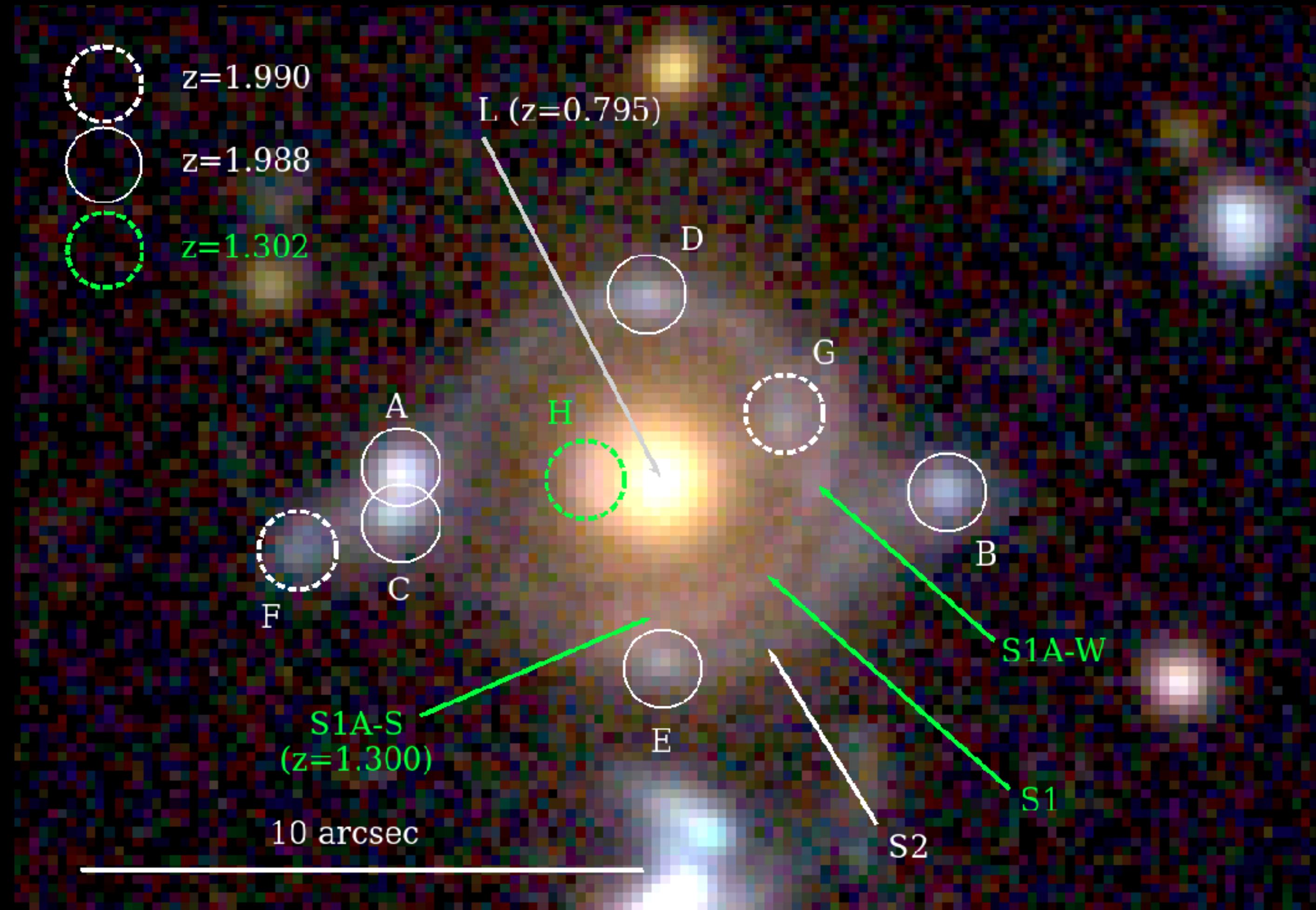


The “Eye of Horus”



A Closer Look Into the Eye of Horus

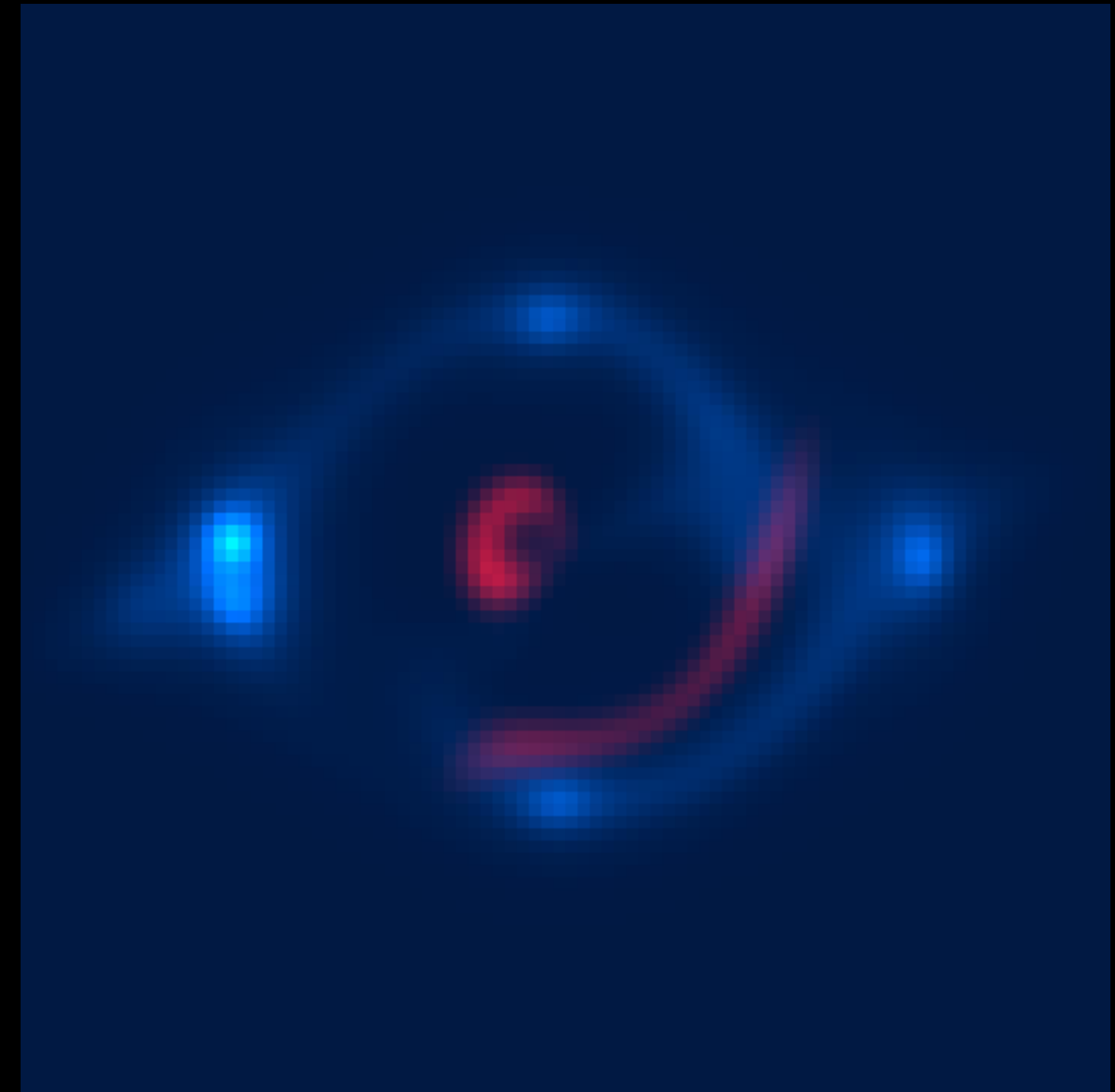
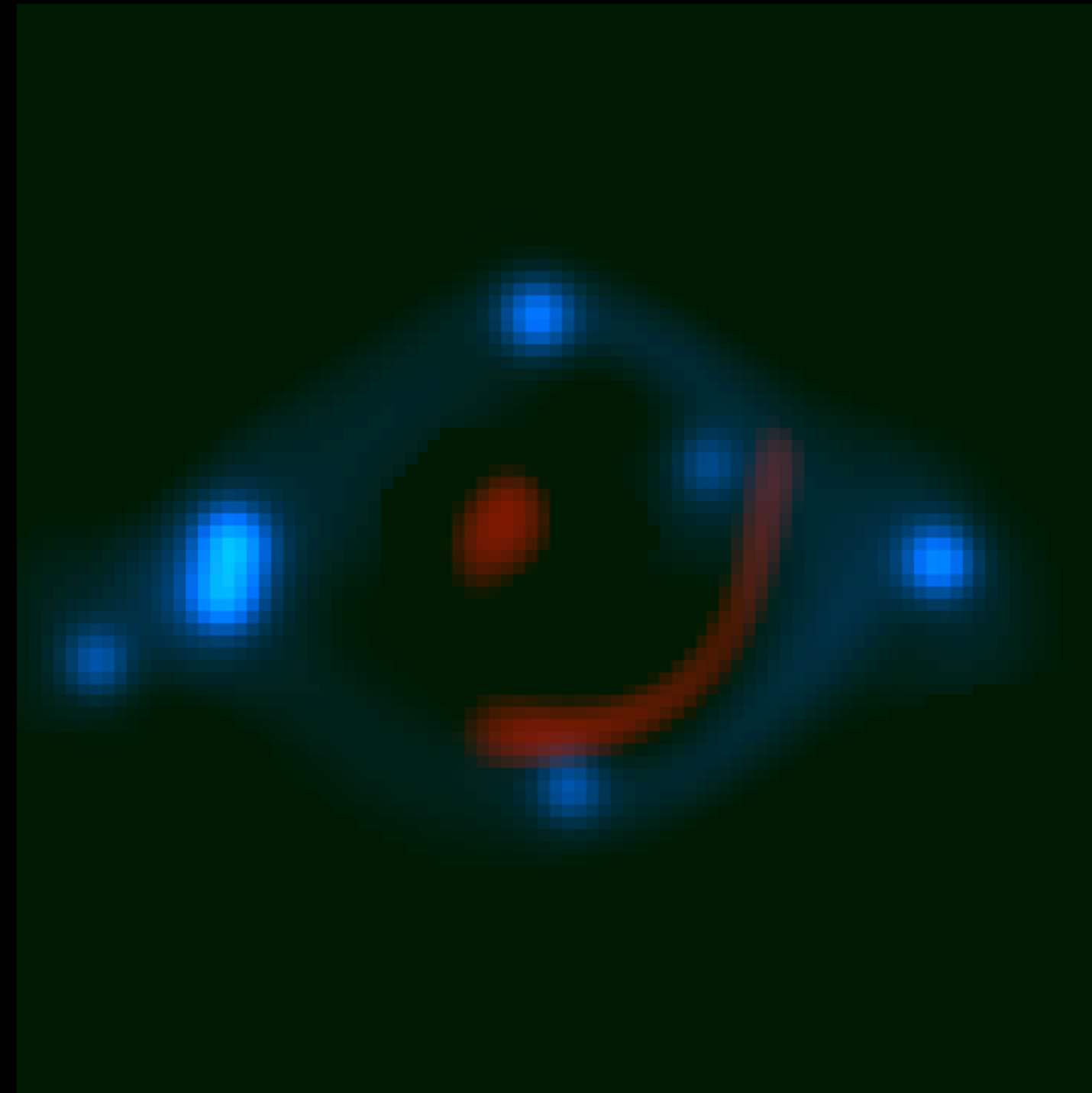
- S1A has slight velocity offset
 - rotating disk?
- Features F and G at $z \approx 1.99$, but slightly offset from rest of S2
 - S2 could be interacting galaxy pair
- One of the images of S2 is split into two distinct peaks (A+C)
 - suggests mass structure causing image splitting
 - no evidence of a galaxy between A and C from ground-based data, need higher resolution



Tanaka, Wong, et al. (2016)

Modeling the Eye of Horus

- Eye of Horus is a “compound” lens
 - S2 is being lensed by both the main lens galaxy and by S1
 - recursive multi-plane effects
- Preliminary models using HSC data
 - broadly reproduce main features
 - require additional mass component to split A+C into two images
- Need higher-resolution data (e.g. *HST*, ALMA, AO) for better constraints
 - ALMA Cycle 4 observations scheduled
 - Subaru/IRCS+AO188 observations in 2017A
- Lens might be in a cluster, need to include environment effects
 - existing multi-object spectroscopy of nearby galaxies
 - upcoming X-ray observations with XMM to probe cluster environment



Science Goals

- Cosmological constraints, complementary to CMB, BAO, time-delay lenses
- Mass structure of lens
 - constrain stellar IMF of early-type galaxy at $z \sim 0.8$
 - structure of satellite galaxy from analysis of high-resolution imaging
- High-resolution studies of source galaxies with ALMA
- Expect to find ~ 10 DSP lenses in HSC survey

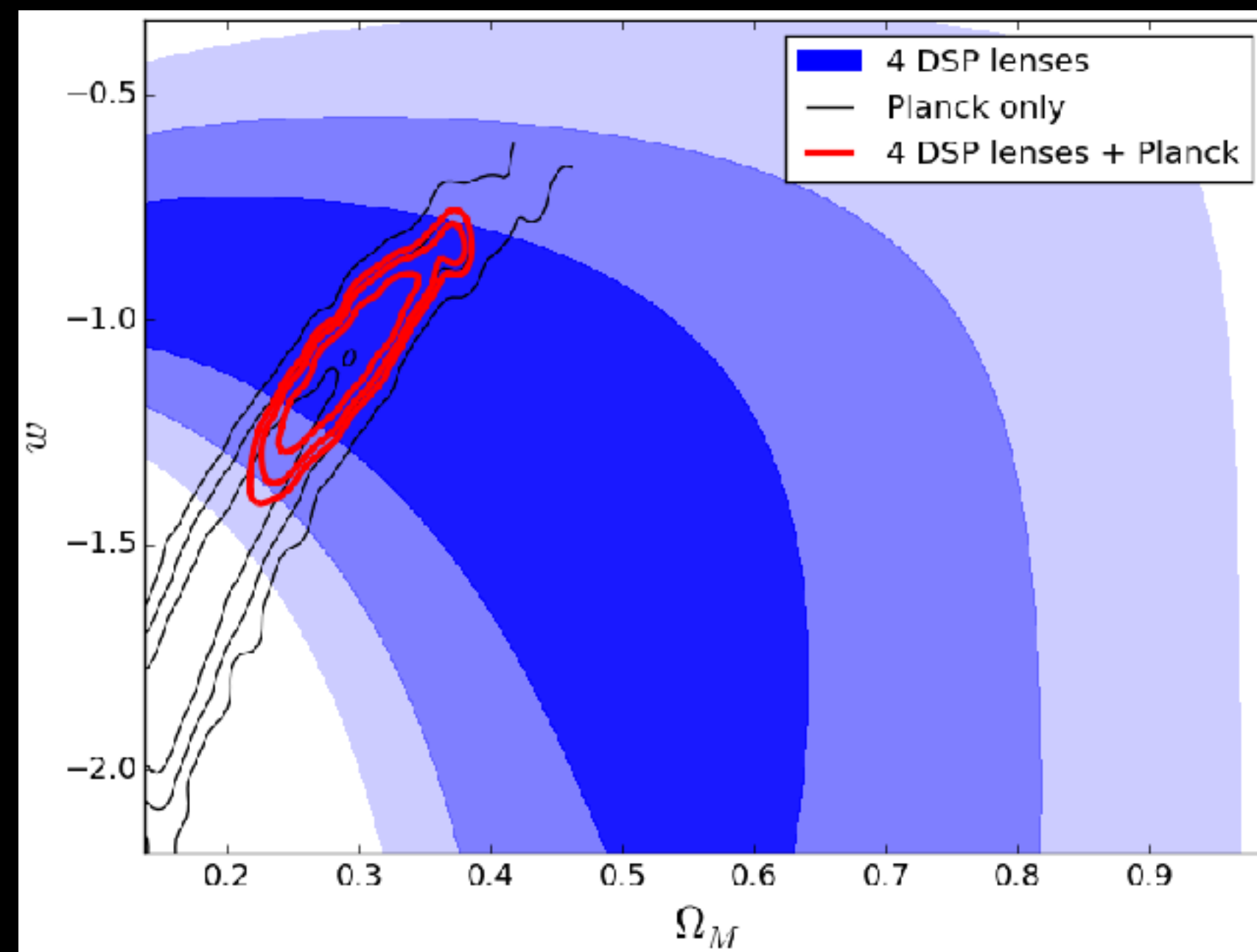


Image credit: A. Sonnenfeld

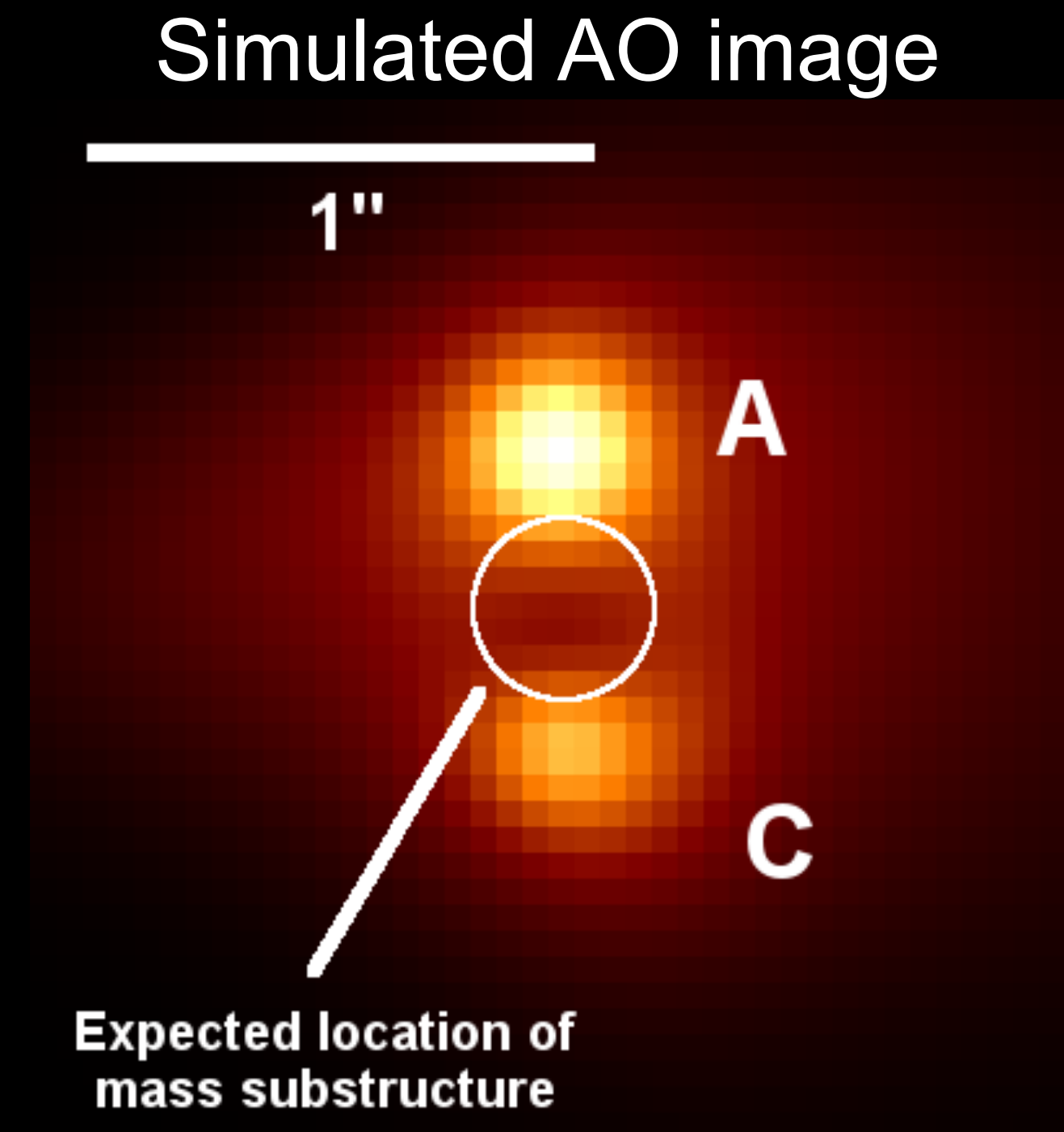


Image credit: A. More

Summary

- HSC SSP will discover hundreds of new lenses at galaxy and group/cluster scales
- Discovery of the “Eye of Horus”, the first double source plane lens found in the HSC SSP (Tanaka, Wong, et al. 2016)
- Eye of Horus is the first DSP lens with all redshifts spectroscopically confirmed
 - $z_L = 0.795$, $z_{S1} = 1.30$, $z_{S2} = 1.99$
- Possible substructure causing additional image splitting of A+C
- High-resolution data (e.g., *HST*, ALMA, AO) needed for more detailed modeling
 - ALMA Cycle 4 observations scheduled
 - 2017A IRCS+AO188 observations scheduled
- Modeling efforts will make this system useful for cosmology and galaxy evolution studies

