## Euclid vs. WFIRST

	Euclid				WFIRST-AFTA			
Aperture	1.2m				2.4m			
Imaging FoV	~0.789×0.709=0.56deg <sup>2</sup>				$0.281 deg^2$			
Area (deg <sup>2</sup> )	15,000				~2000			
Instrument	Optical	NIR			NIR			
Wavelength range (nm)	VIS (550-900)	Y(920-1 146)	J(1146-1 372)	H(1372-20 00)	Y(927-11 92)	J(1131-14 52)	H(1380-77 4)	F(1683-20 00)
Depth	24.5	24	24	24	26.7	26.9	26.7	26.2
Pixel scale(")	0.1	0.3			0.12	0.12	0.14	0.14
Survey regions	All-sky				Mostly southern hemisphere			
Duration	~5yrs				1.3yrs			
Ground-based partners	South: DES+LSST North: not yet determined (Pan-Starrs not preferred, CFHT seems difficult)				Mostly LSST			
Collaboration with the Subaru community	HSC: Proposed to Subaru community to carry out "gri" imaging with HSC for Northern hemisphere (give ~40 scientists to Euclid data)				The Subaru community showed strong interest in participation, especially after WISH (Yamada san is leading the coordination)			

Note:

Euclid and WFIRST both have spectroscopy capability for BAO (R~150 for Euclid, R~550-800 for WFIRST)

Euclid has deep survey regions (40 deg<sup>2</sup>, Y,J,H~26 mag)

WFIRST has IFU capability (3.00×3.15arcsec, 600-2000nm, R~100)

## Ultra-Wide, shallow HSC survey for Euclid

- FYI the DES depth is ~1000 sec. for each of griz
- If the HSC data (g:125sec, r:125sec, i:250sec) is deep enough for Euclid (thanks to 8.2m aperture and much better seeing), how many Subaru nights are needed to cover ~5000 sq. degrees?

$$\frac{5000 \text{ sq. degs.}}{9 \text{ hr/night} \times 60 \text{ min/hr} \times 60 \text{ sec/min}} \times \left[\pi \times (1.5 \text{ deg./2})^2\right] \approx 75 \text{ nights}$$
$$2 \times 250 \text{ sec} + 12 \times 30 \text{ sec}$$
$$\times \left[\pi \times (1.5 \text{ deg./2})^2\right] \text{ w/o weather fac.}$$

- This number ~75 nights is much smaller than what we were requested a few years ago (~200 nights)
- What are benefits?
  - ~40 Japanese scientists can join the Euclid consortium
  - Subaru community should have the full access, at least, to HSC data
  - Cosmology (clusters, WL, large-scale structure, ...)
  - Galactic Archaeology (MW structure up to the halo edge)
  - Time domain astronomy (e.g., a follow up of GW candidates, RR-Lyrae star search in MW halo region)

## Synergy of Ultrawide HSC survey with eROSITA

SRG/eROSITA and HSC surveys

