

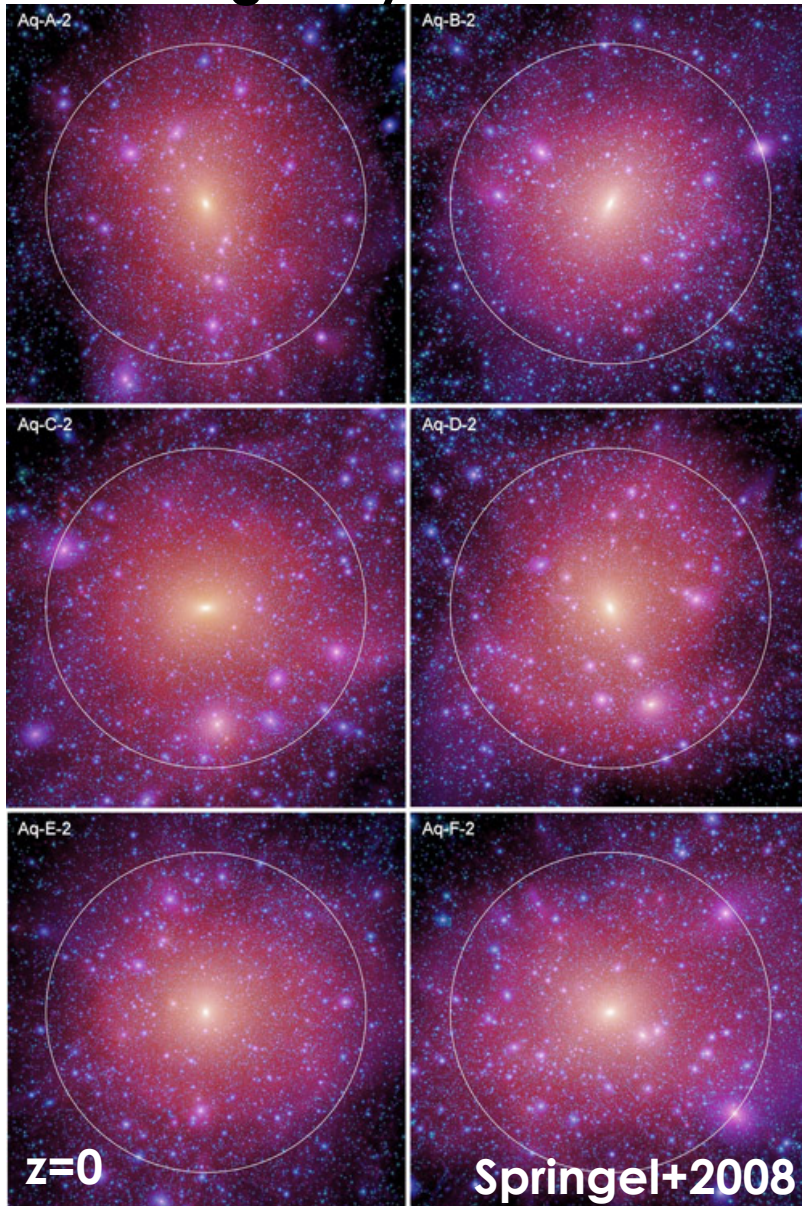
Subaru UM, Jan, 19, 2016

M81; Near field cosmology beyond the Local Group

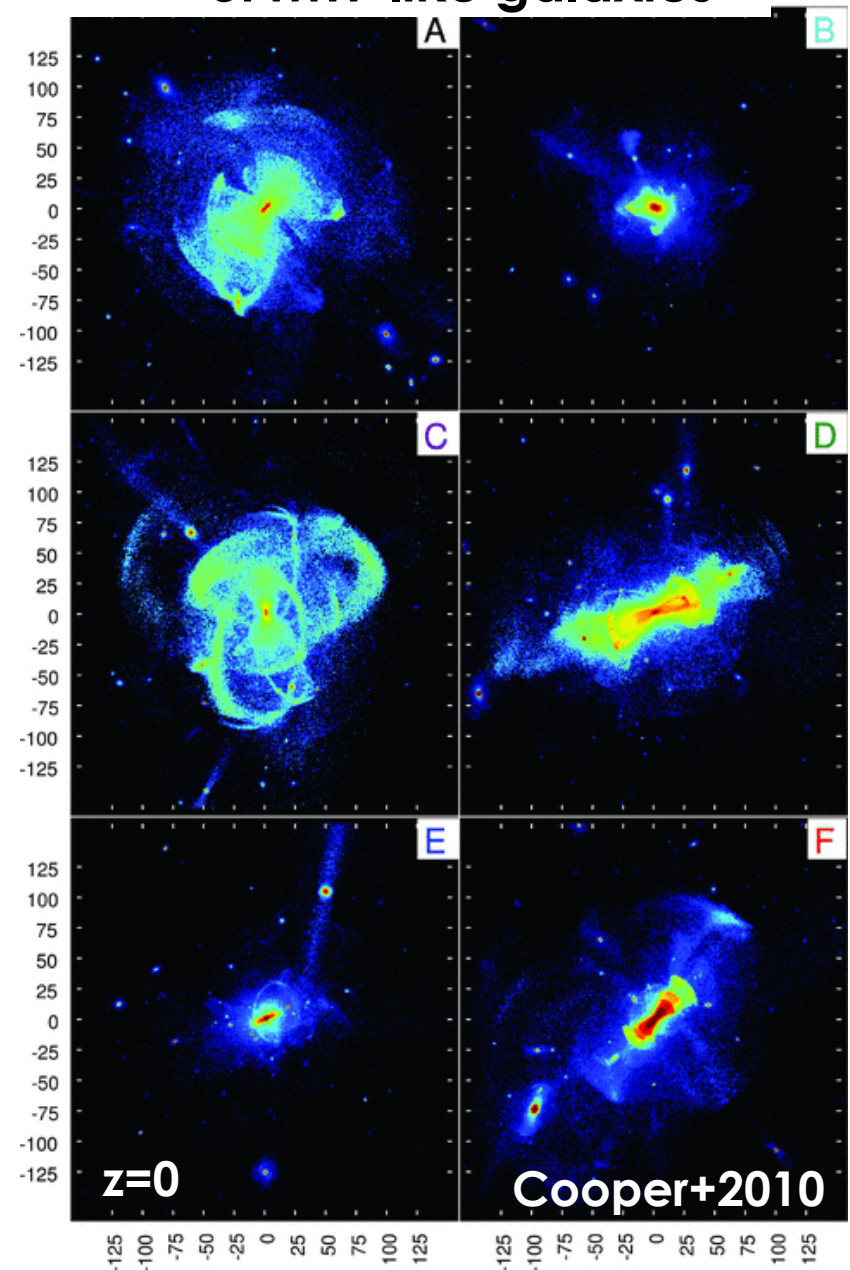
Sakurako Okamoto
Shanghai Astronomical Observatory

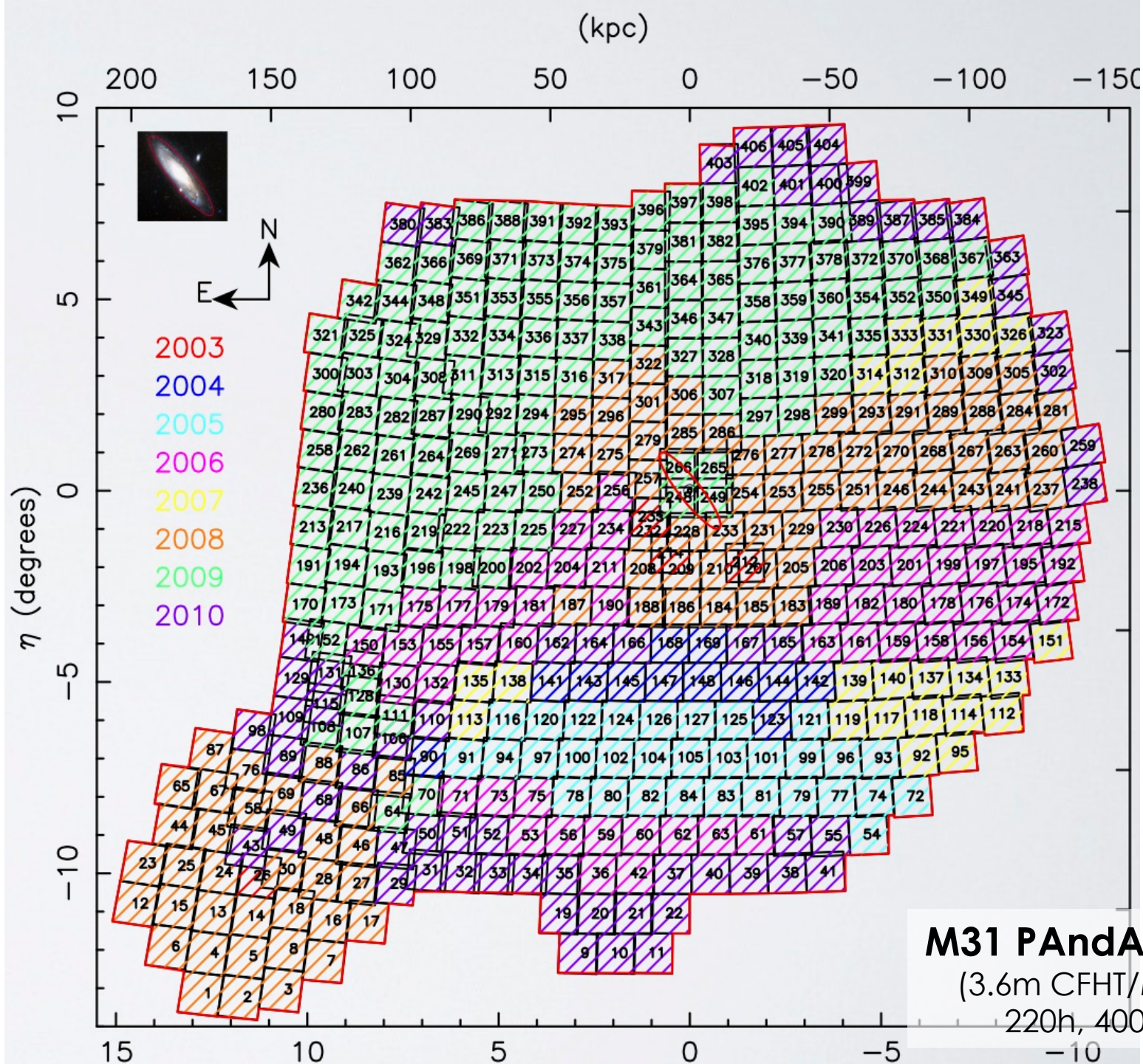
Collaborators: N. Arimoto (NAOJ), A. Ferguson (Edinburgh), M. Irwin (Cambridge)
E. Bernard (Edinburgh), Y. Yamada (NAOJ), U. Utsumi (Hiroshima)

DM densities of MW-like galaxy halos



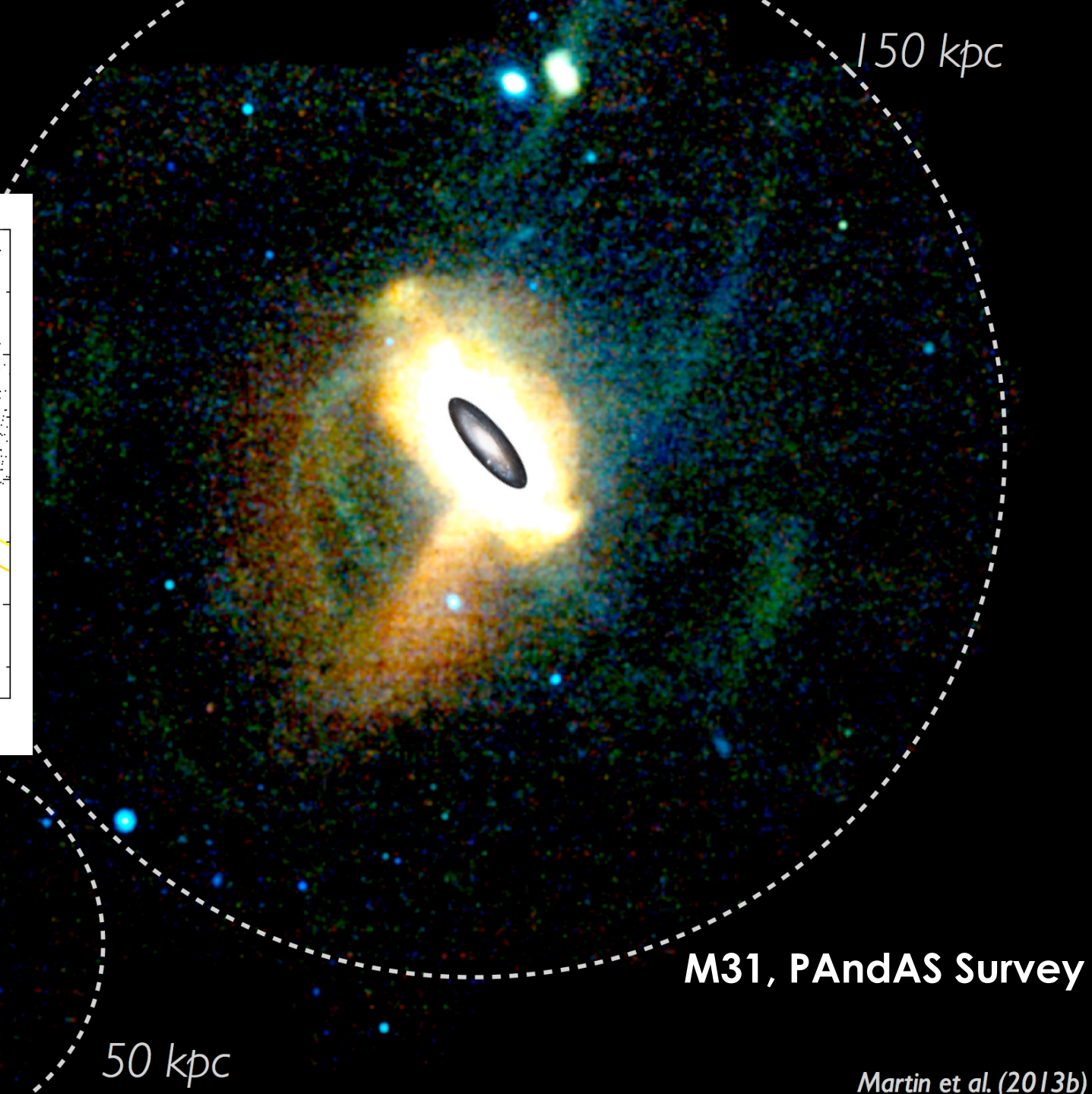
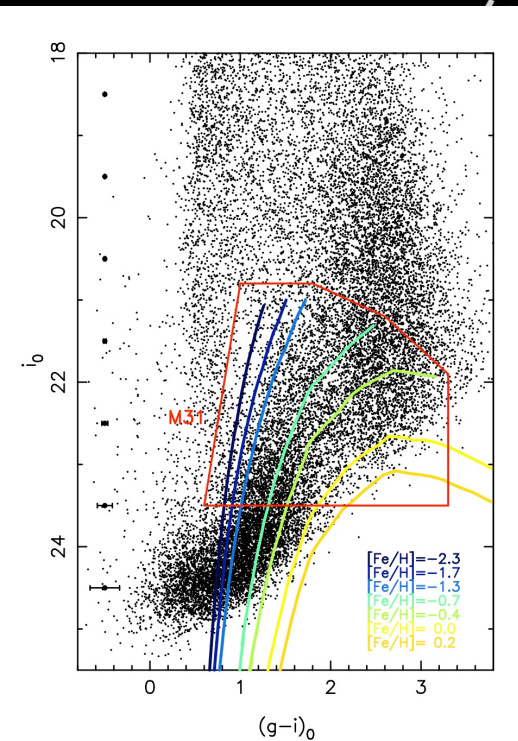
Stars formed in satellites of MW-like galaxies





M31 PAndAS Survey
(3.6m CFHT/Megacam
220h, 400 pointings)

$[\text{Fe}/\text{H}] \sim -2.3$
 $[\text{Fe}/\text{H}] \sim -1.4$
 $[\text{Fe}/\text{H}] \sim -0.7$



M31, PAndAS Survey

Martin et al. (2013b)

Dwarf galaxies

Zucker et al. (2004; 1)

Martin et al. (2006; 3)

Majewski et al. (2007; 1)

Ibata, Martin et al. (2007; 2)

Irwin et al. (2007; 1)

Zucker et al. (2007; 1)

McConnachie et al. (2008; 3)

Martin et al. (2009; 2)

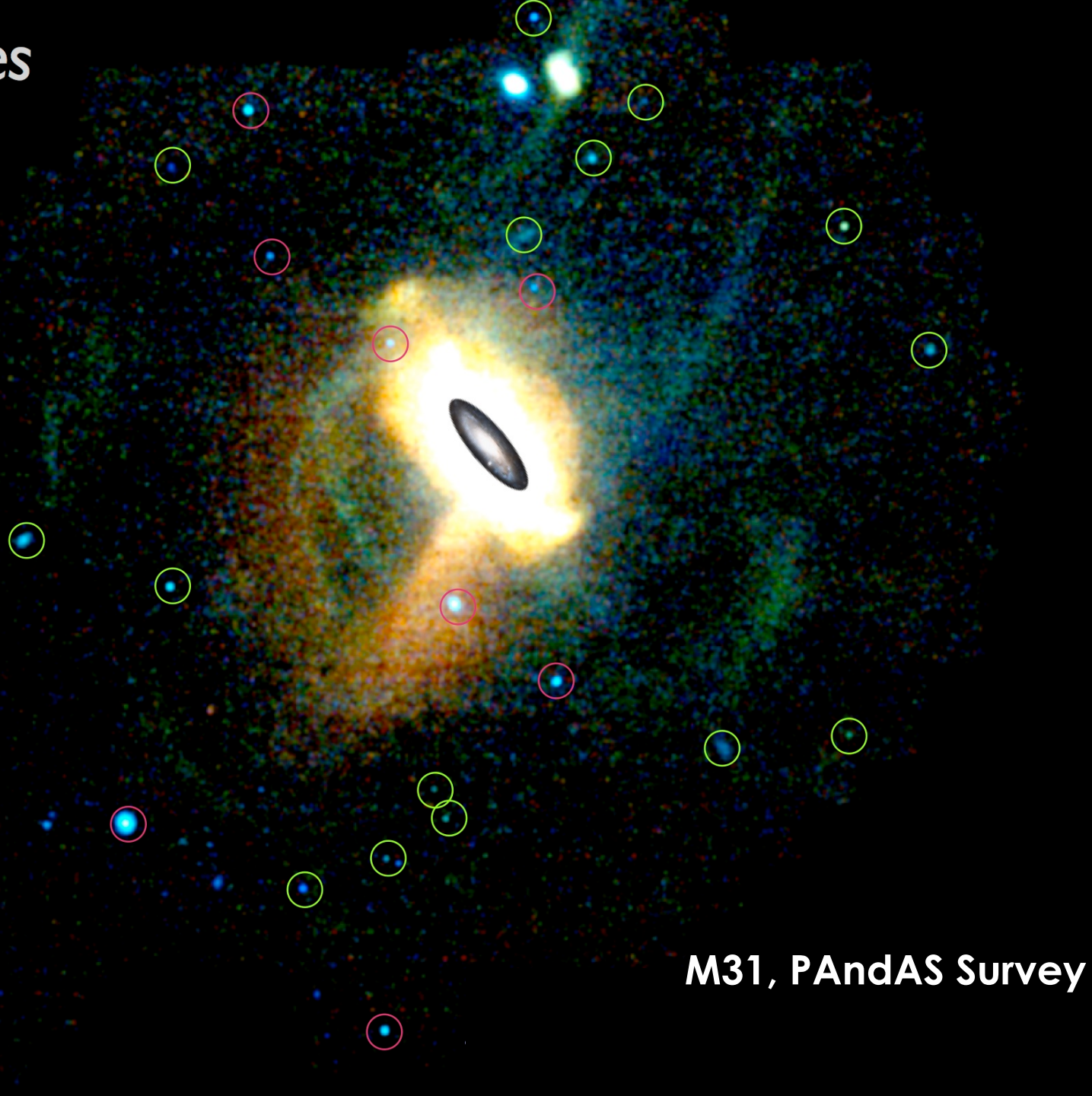
Richardson et al. (2011, 5)

Bell, Slater & Martin (2011, 1)

Slater, Bell & Martin (2011, 1)

Martin et al. (2013a, 2)

Irwin et al. (in prep, 1)



M31, PAndAS Survey

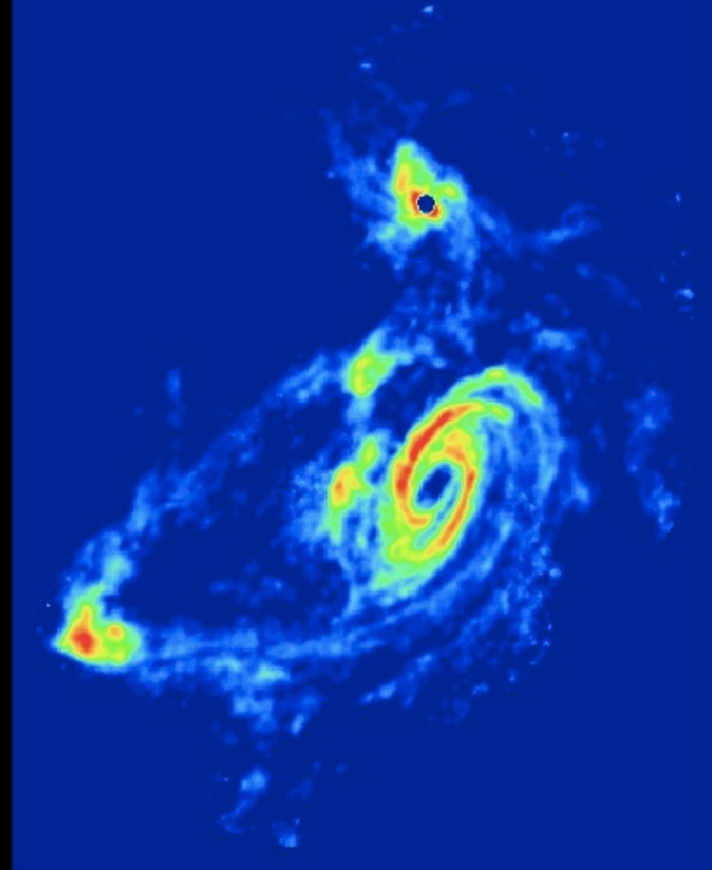
Near field cosmology beyond the LG with HSC/Subaru

M82
 $M_V = -19.4$

M81
 $M_V =$

NGC3077
 $M_V = -17.2$
(LMC: -17.6)

HI distribution (Yun et al. 1994)



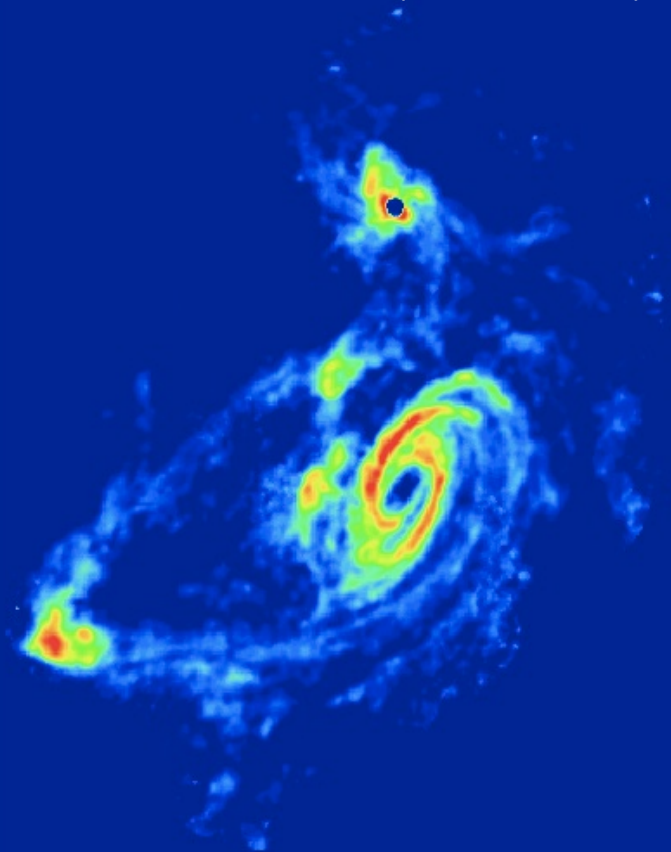
Near field cosmology beyond the LG with HSC/Subaru (S14A-184, S14B-101, S15B-073, S16A-096, PI: S.Okamoto)

M81, MW analogue lies at 3.6Mpc

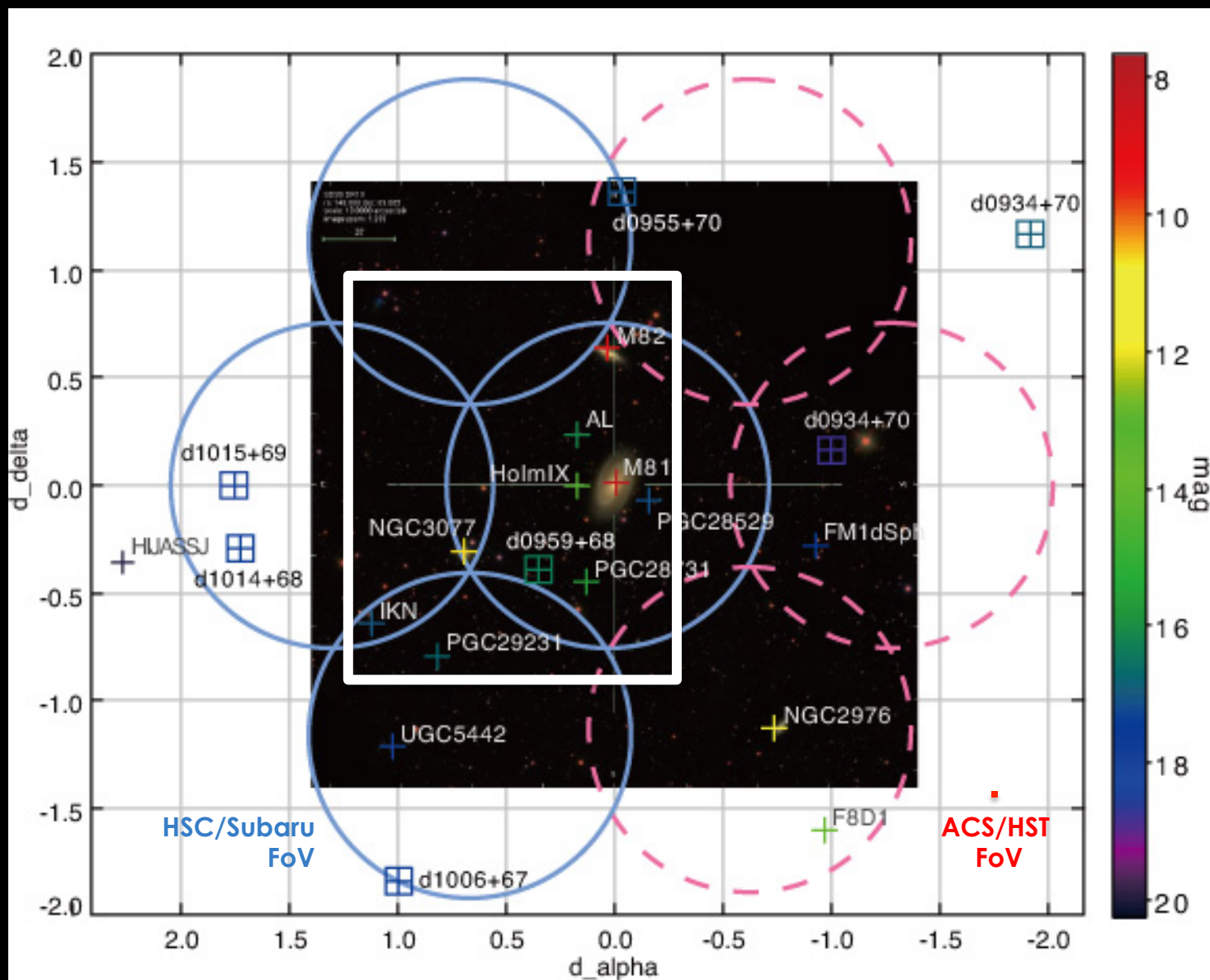
- ✧ True nature of young/old stellar contents in the large scale structures
- ✧ Age and metallicity constraints for the extended component
- ✧ New stellar debris, satellites, streams, arcs in the entire region
- ✧ Globular clusters in the M81 group



HI distribution (Yun et al. 1994)



Near field cosmology beyond the LG with HSC



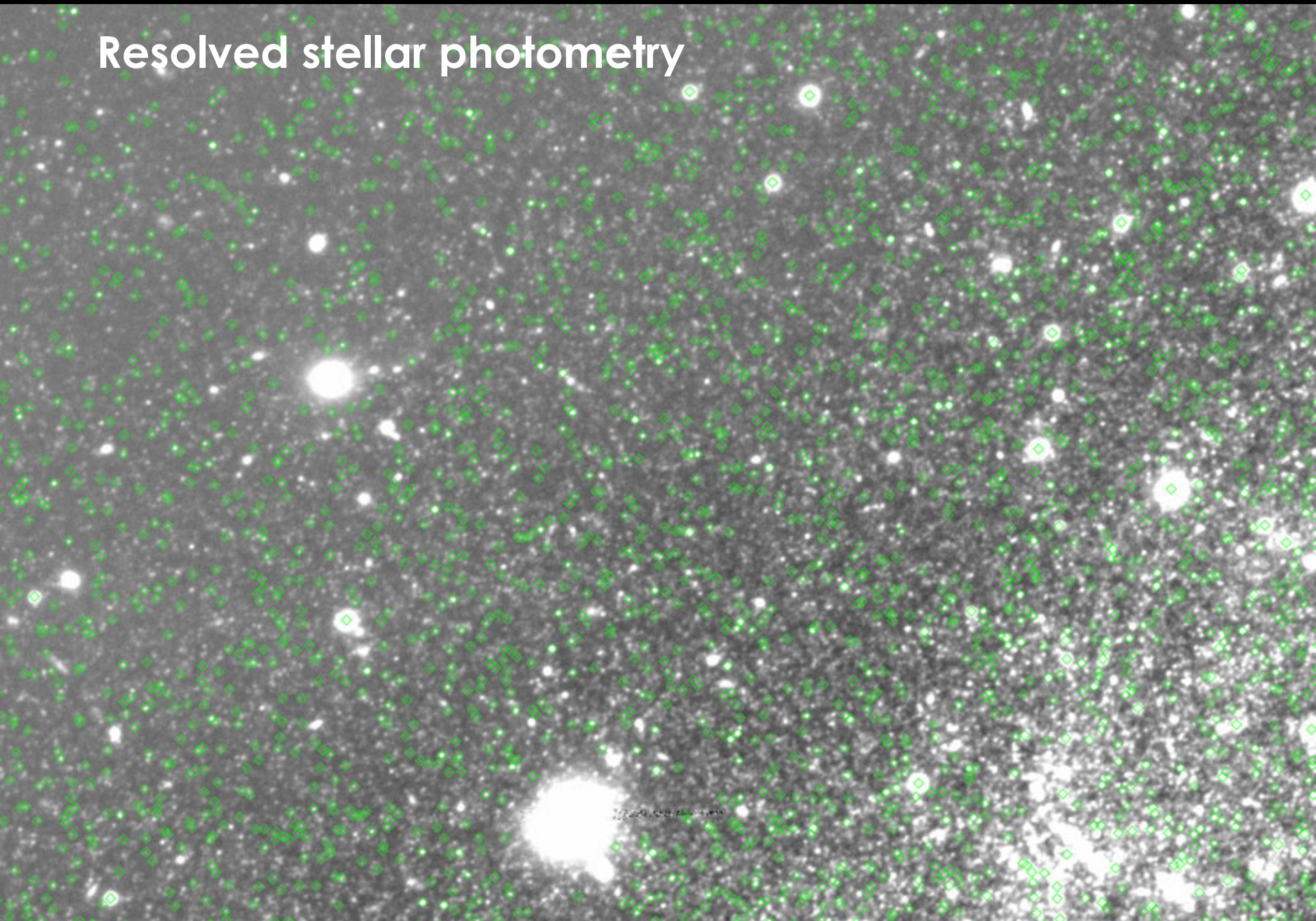
M81 photometry

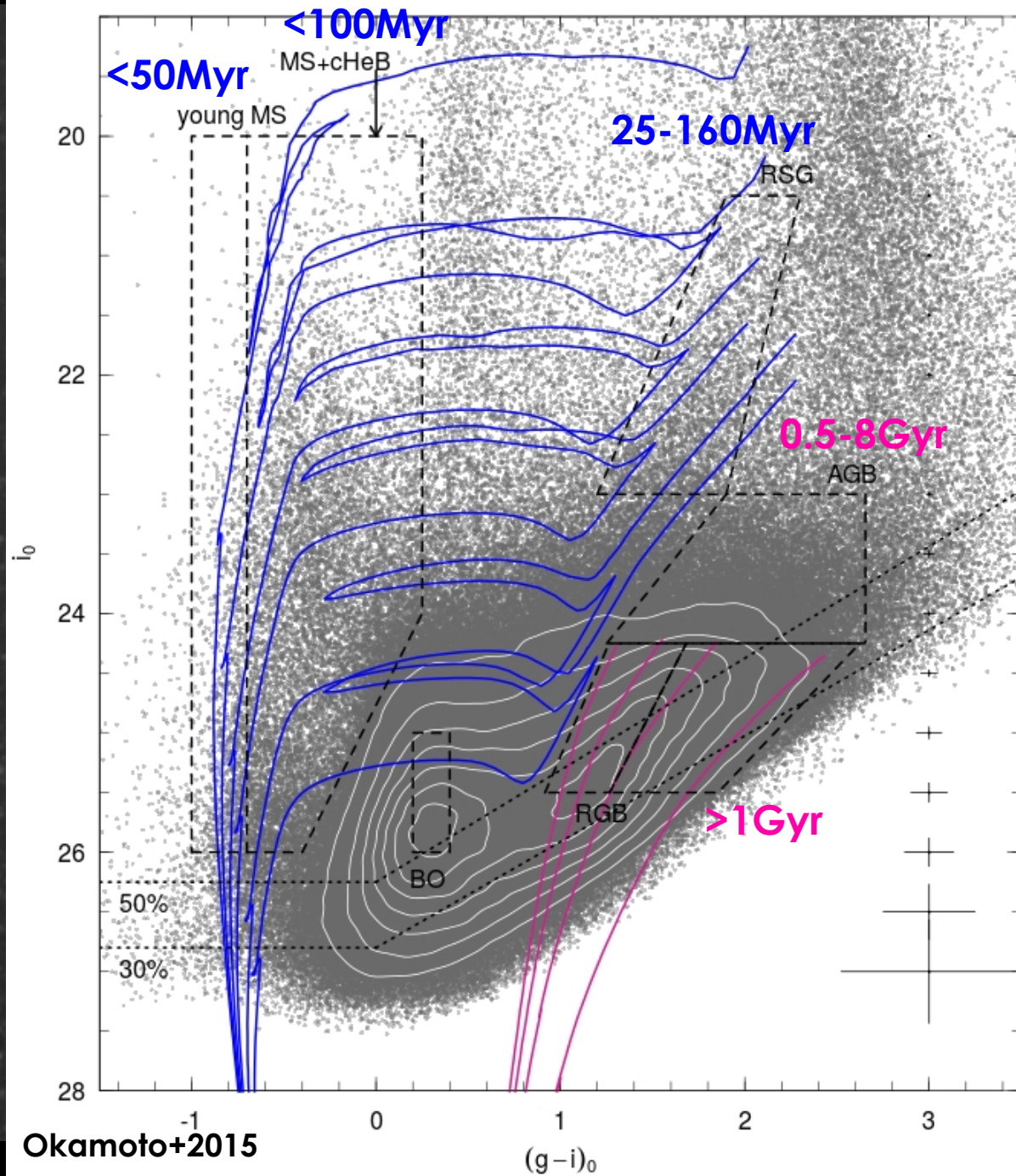


M81 photometry

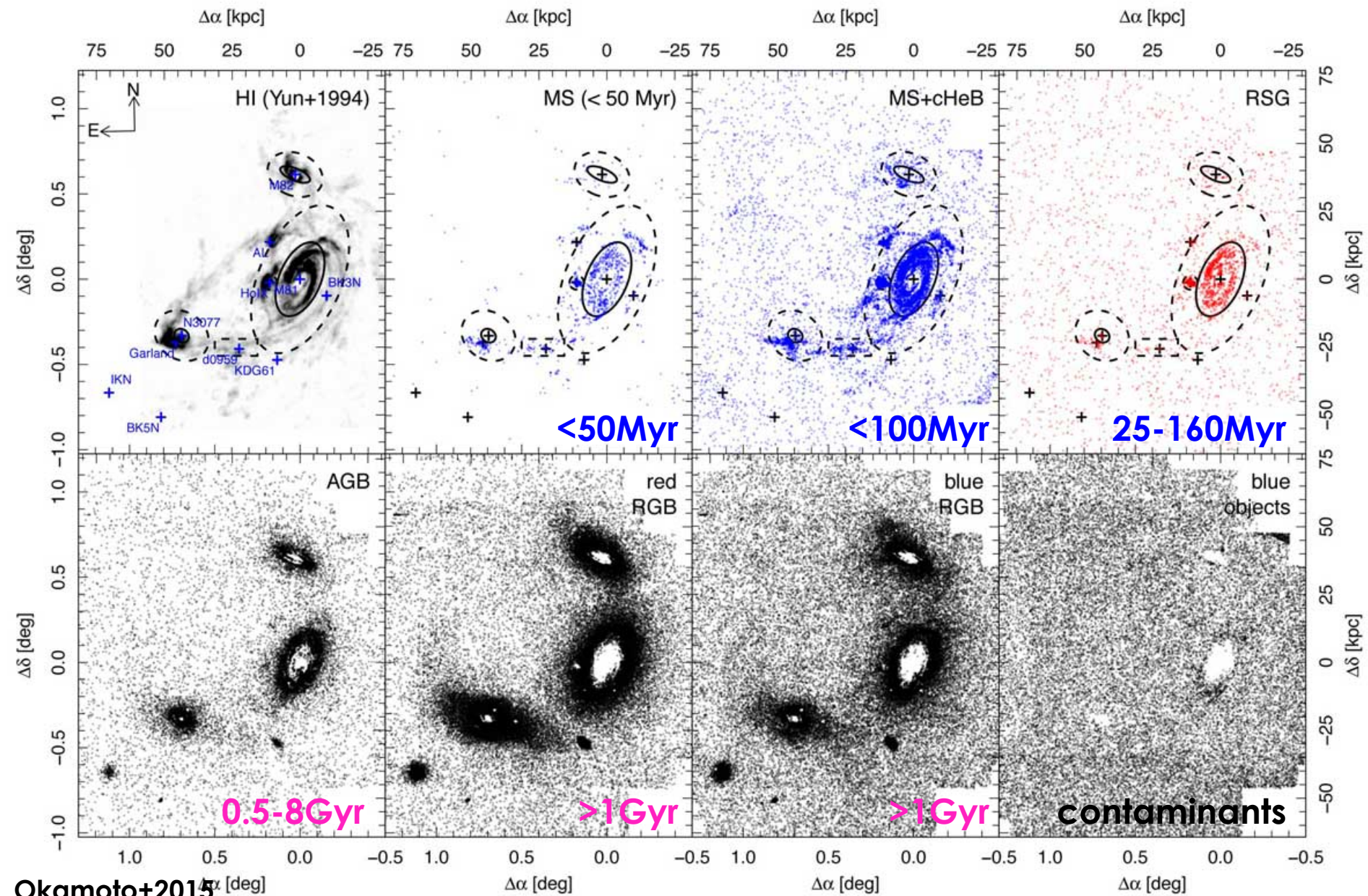


Resolved stellar photometry



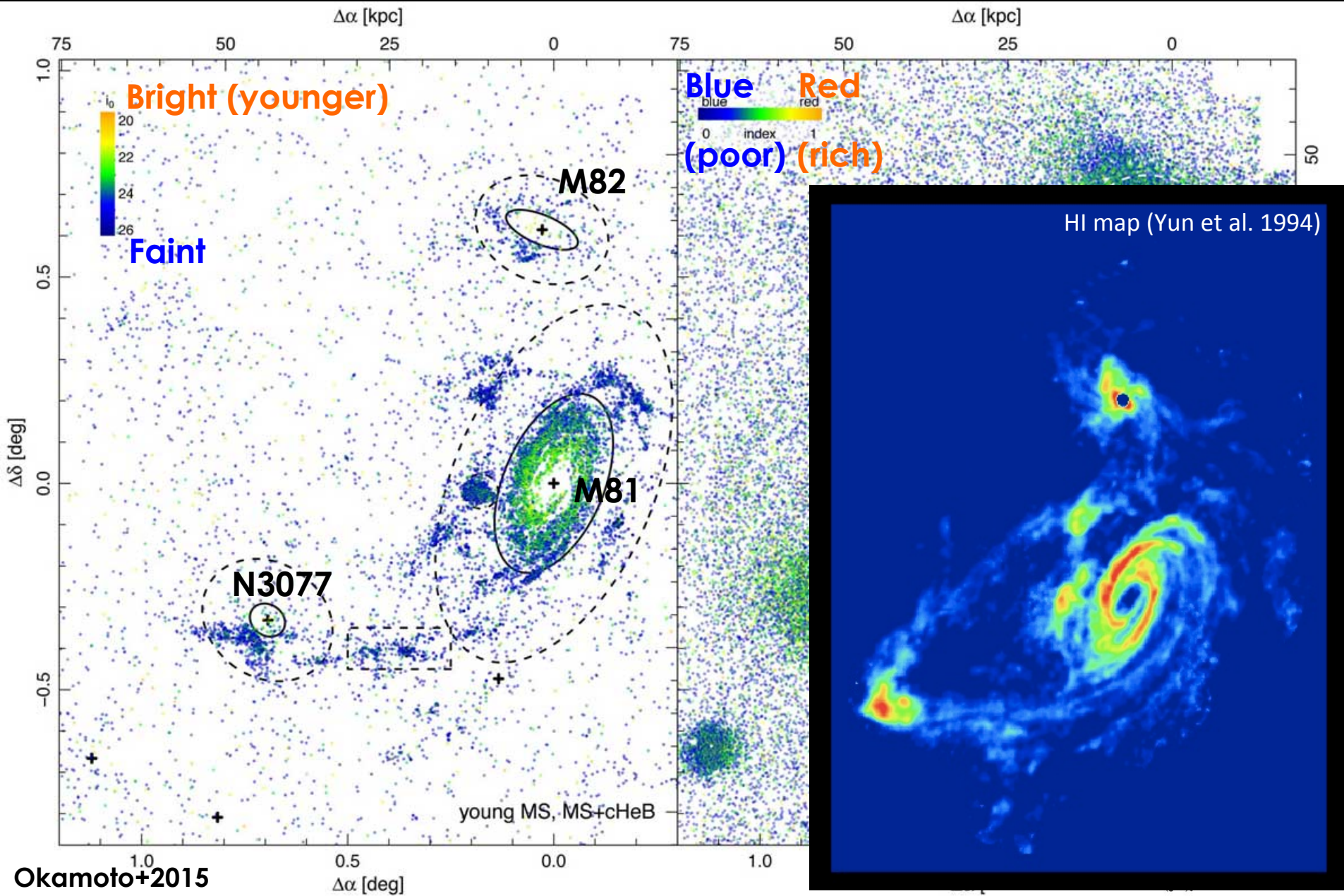


Spatial distributions of each populations



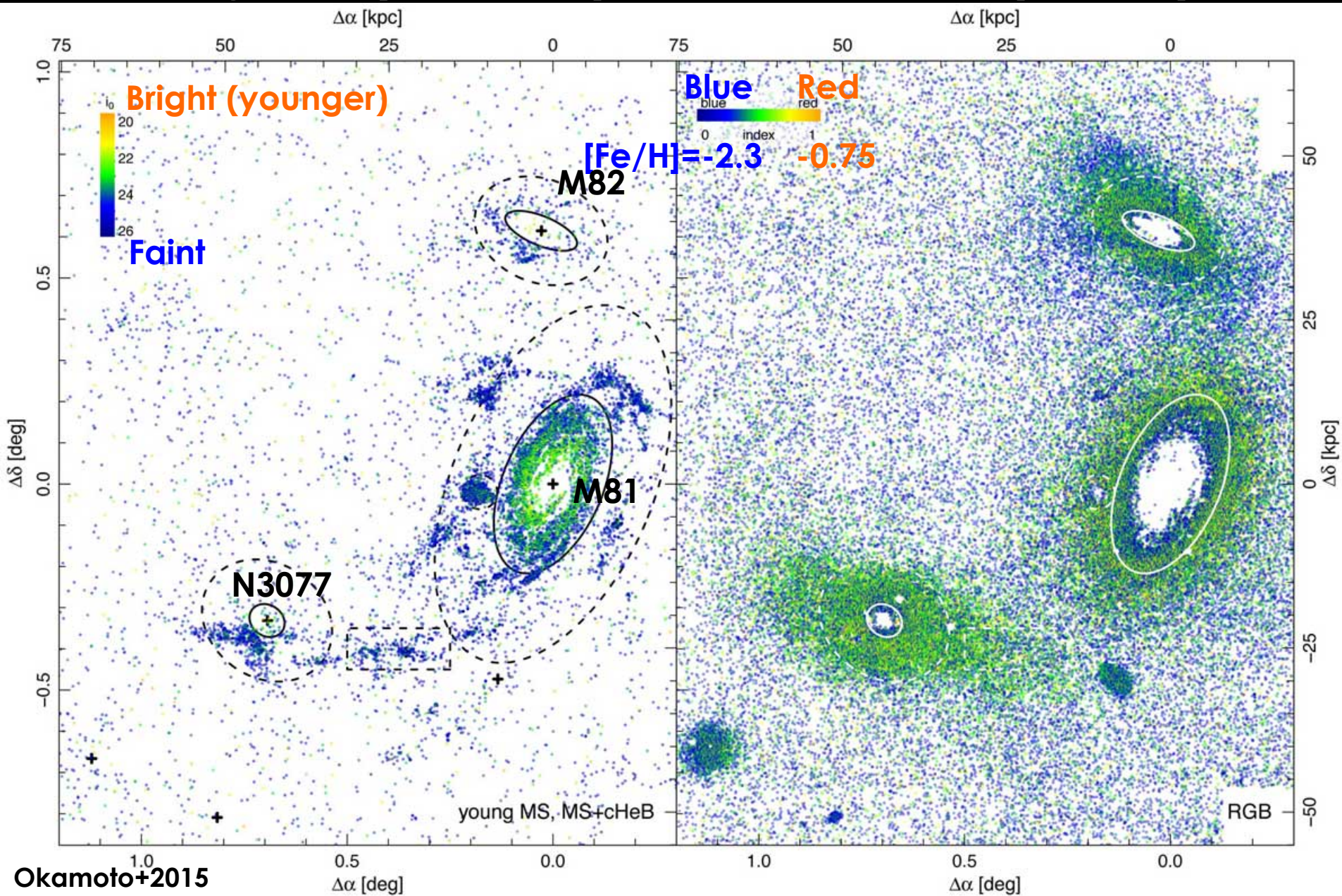
Young stars (<100Myrs)

Old stars (>1Gyrs)

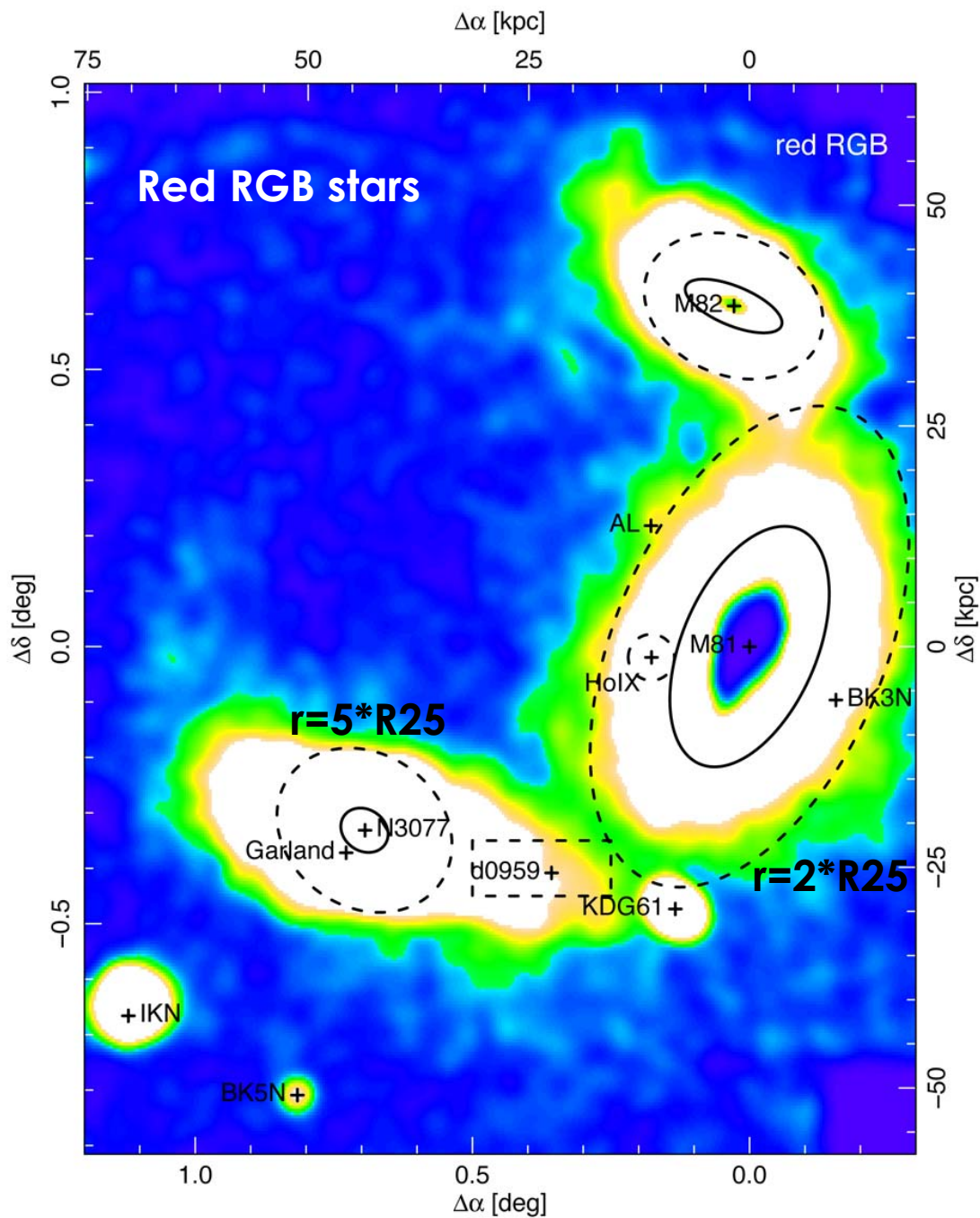
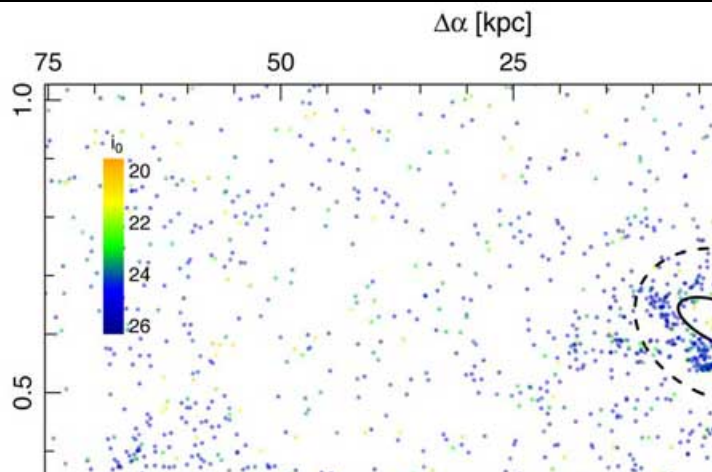


Young stars (<100Myrs)

Old stars (>1Gyrs)

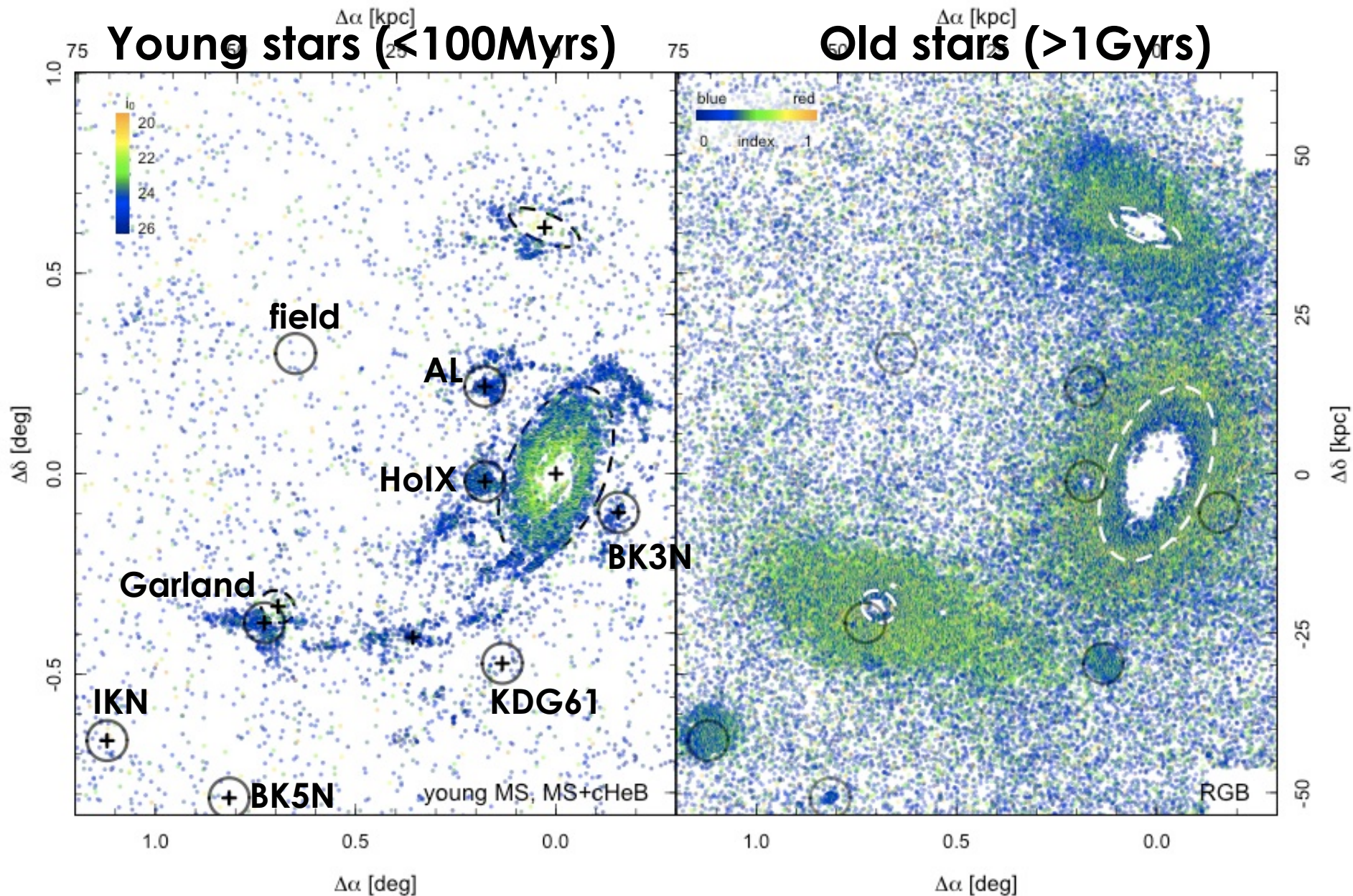


Spatial distributions



Okamoto+2015

Dwarf galaxies in M81 group



Summary

M81 galaxy archeology : We present the first results from a deep wide-field imaging survey of the M81 group center that we are conducting with HSC/Subaru, which reveals the first truly panoramic view of the low surface brightness stellar component.

- ✧ **Young intra-group pop follow the HI filamentary distributions.**
- ✧ **Young stars in a stellar stream between M81 and NGC3077, AL, HolX, Galand and in numerous outlying stellar associations, which SF was synchronized each other (i.e. tidal dwarfs).**
- ✧ **Very extended ($>2 \times R_{25}$) old stellar halos of M81, M82, NGC3077**
- ✧ **Highly disturbed morphologies of M82 and NGC3077 halo**
- ✧ **Bluer (more metal-poor) outer halo of M82**

On-going works....

Analyses of halo profiles & individual dwarf galaxies,
Exploration of new satellites/streams/substructures,
Data processing, Photometry, Observations,
Writing the follow-up proposals & papers...